

GAIA MISSION

General news and Science of SSOs

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PSL

Observatoire
de Paris



Acknow.: P. Tanga, F. Spoto, F. Mignard, CU4-SSO

Jan. 2020 – ИПА, Санкт Петербург, Россия



Outline

- *The Gaia mission, the satellite, the DPAC*
- *Science - general results with DR2*
- *SSOs (and exoplanets)*
- *Future*

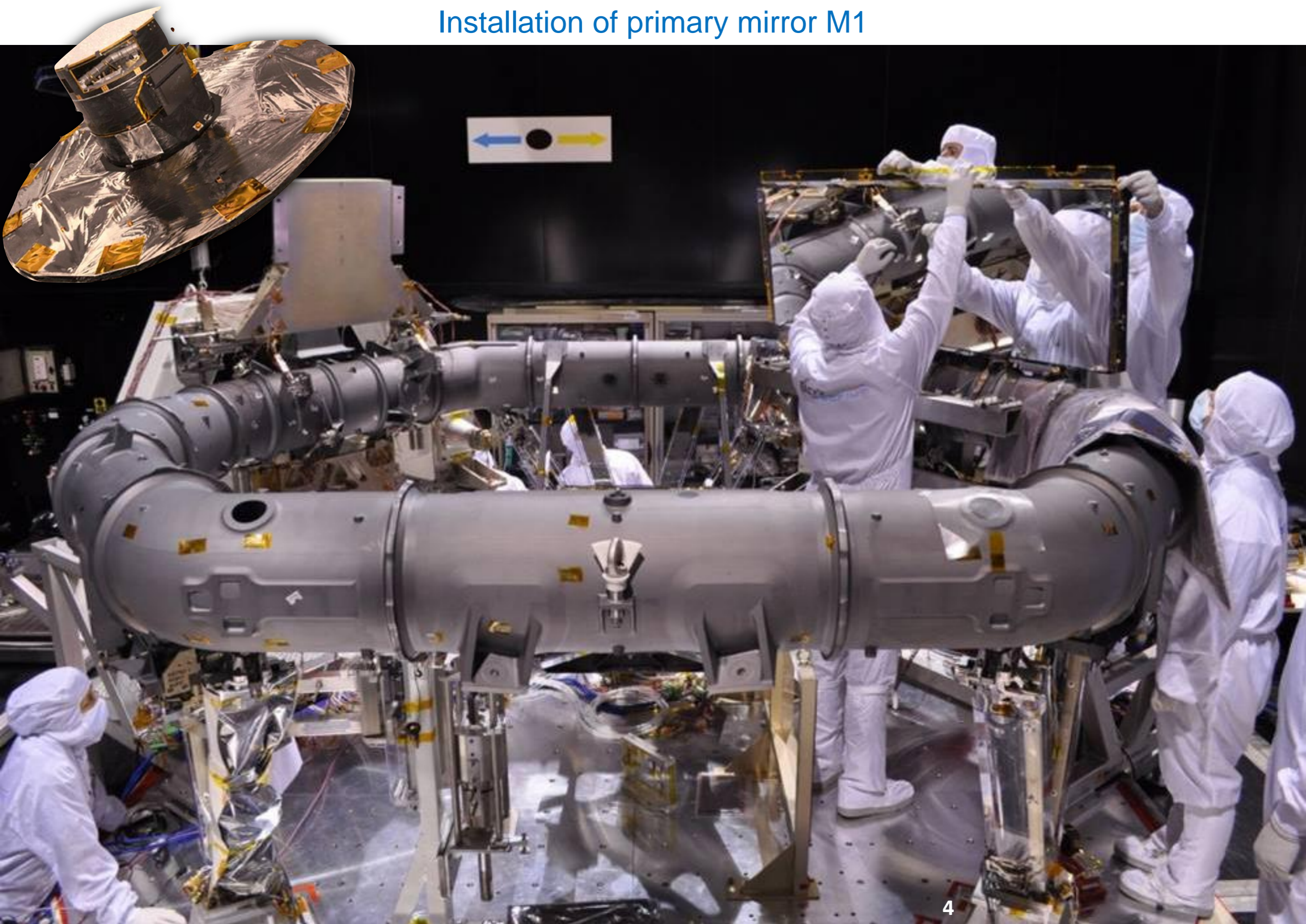
Gaia Mission

3D (6D) Census of the Milky Way
unravelling its dynamical and chemical history

Perform global astrometry with high precision (π , μ , rv)
Observing the whole sky (North and South) with same instrument
Continuously scanning the sky in space (no atmosphere),

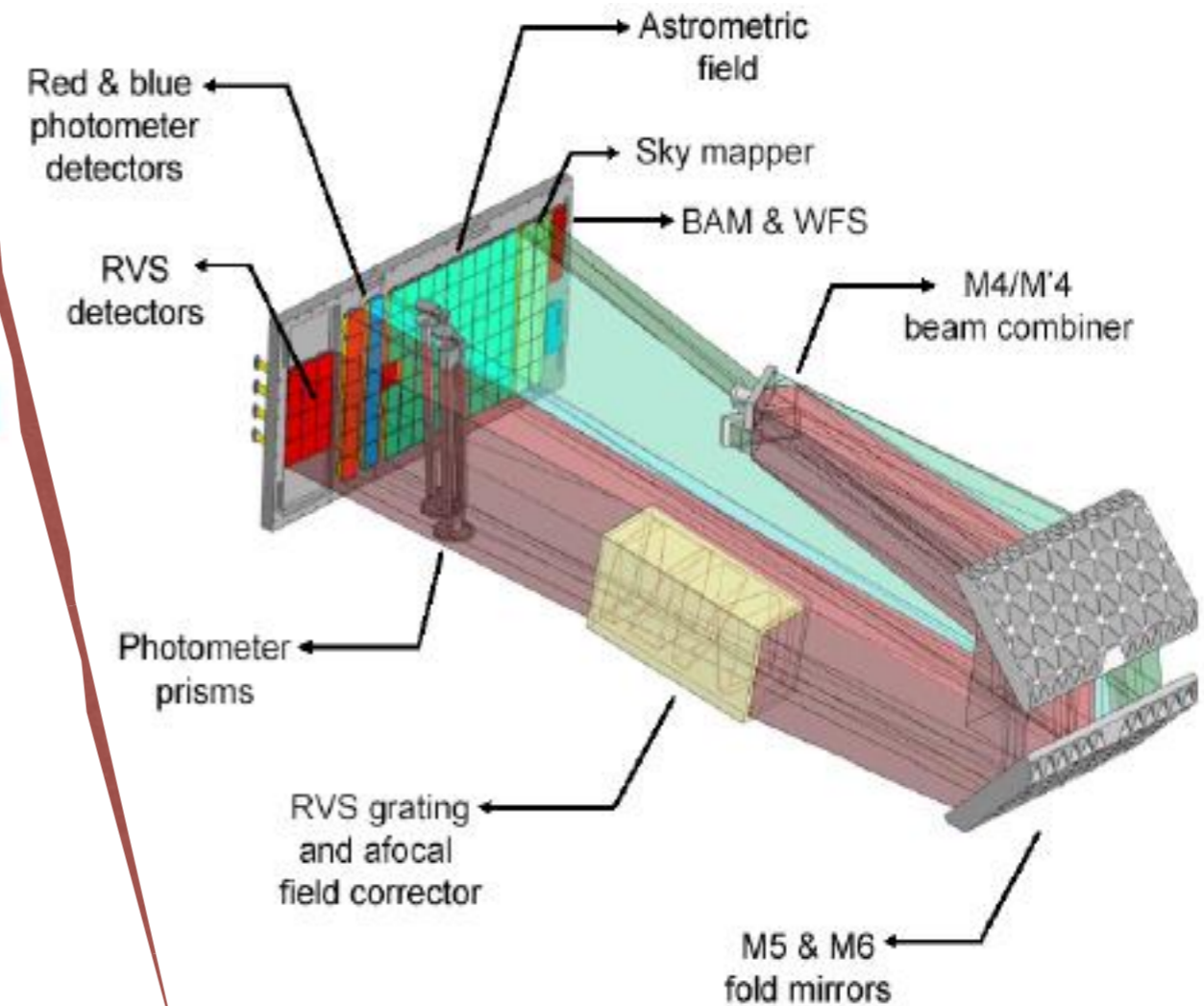
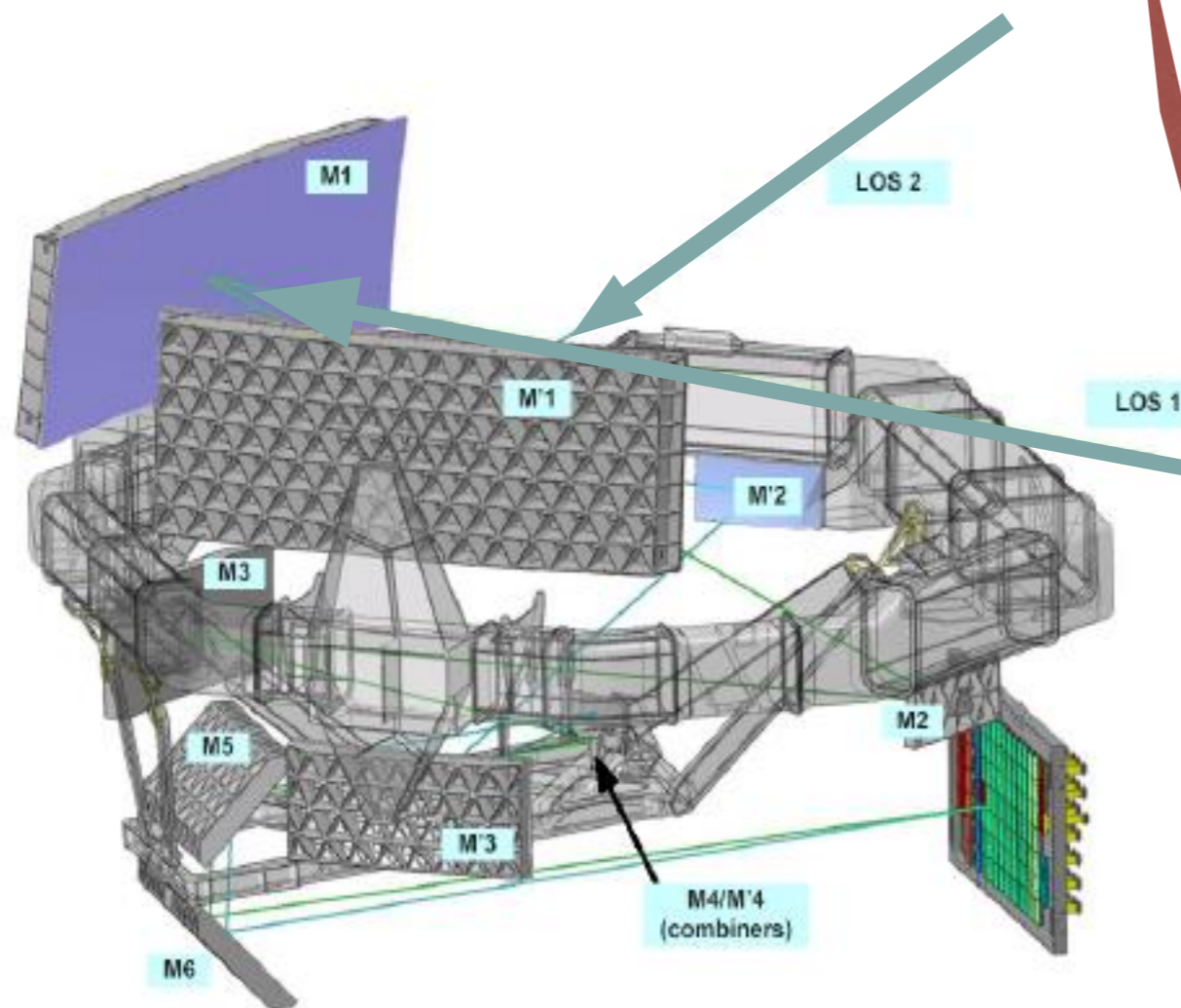
Measuring epoch positions & brightness over years
and their variation: motion & variability

Installation of primary mirror M1





Focal plane - Payloads



*One single focal plane ; 3 instruments
astrometry,
spectro-photometry,
spectrometry*

Focal plane - Payloads

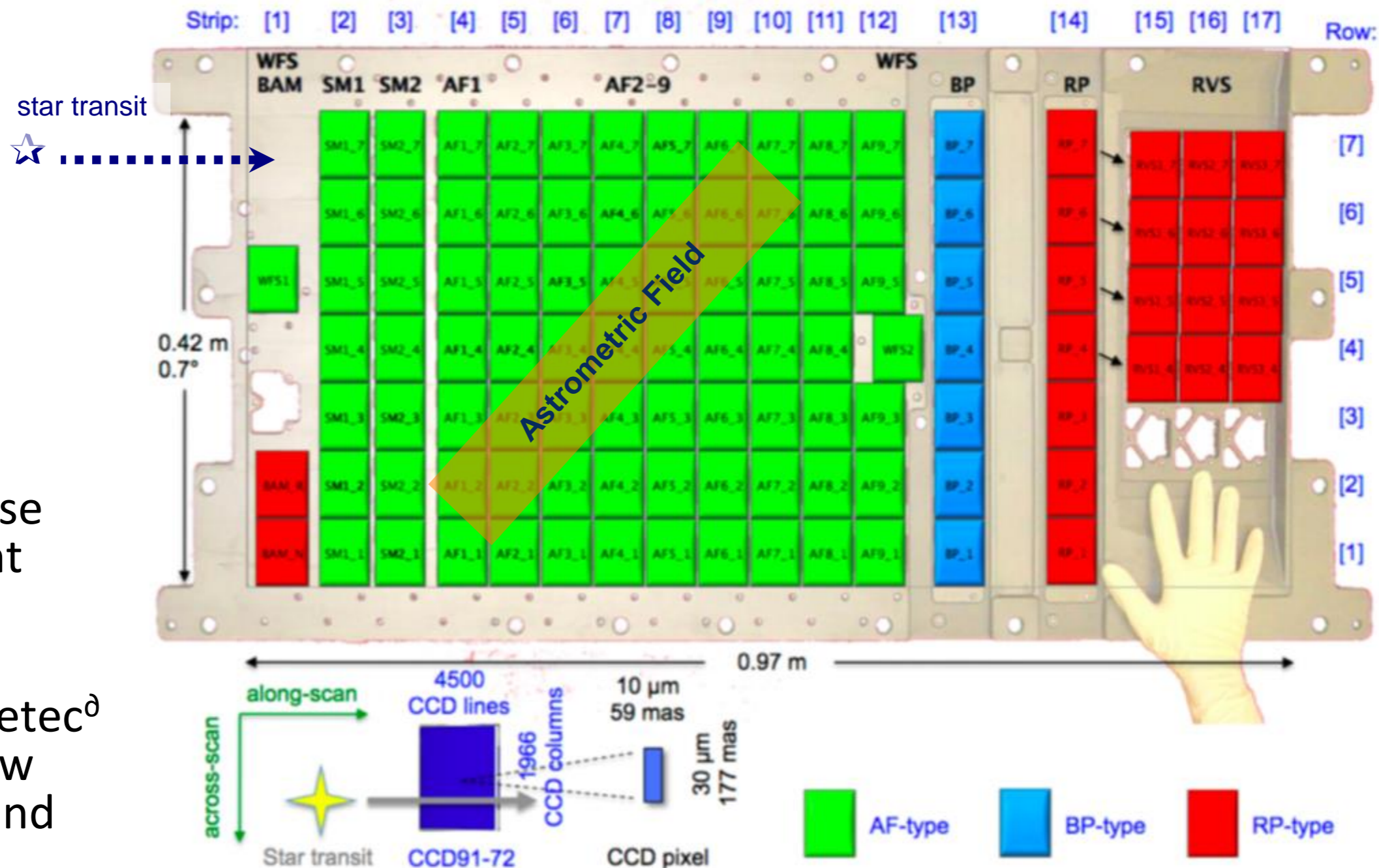


Astrometry/photometry

Spectro-photom.

Spectrometry

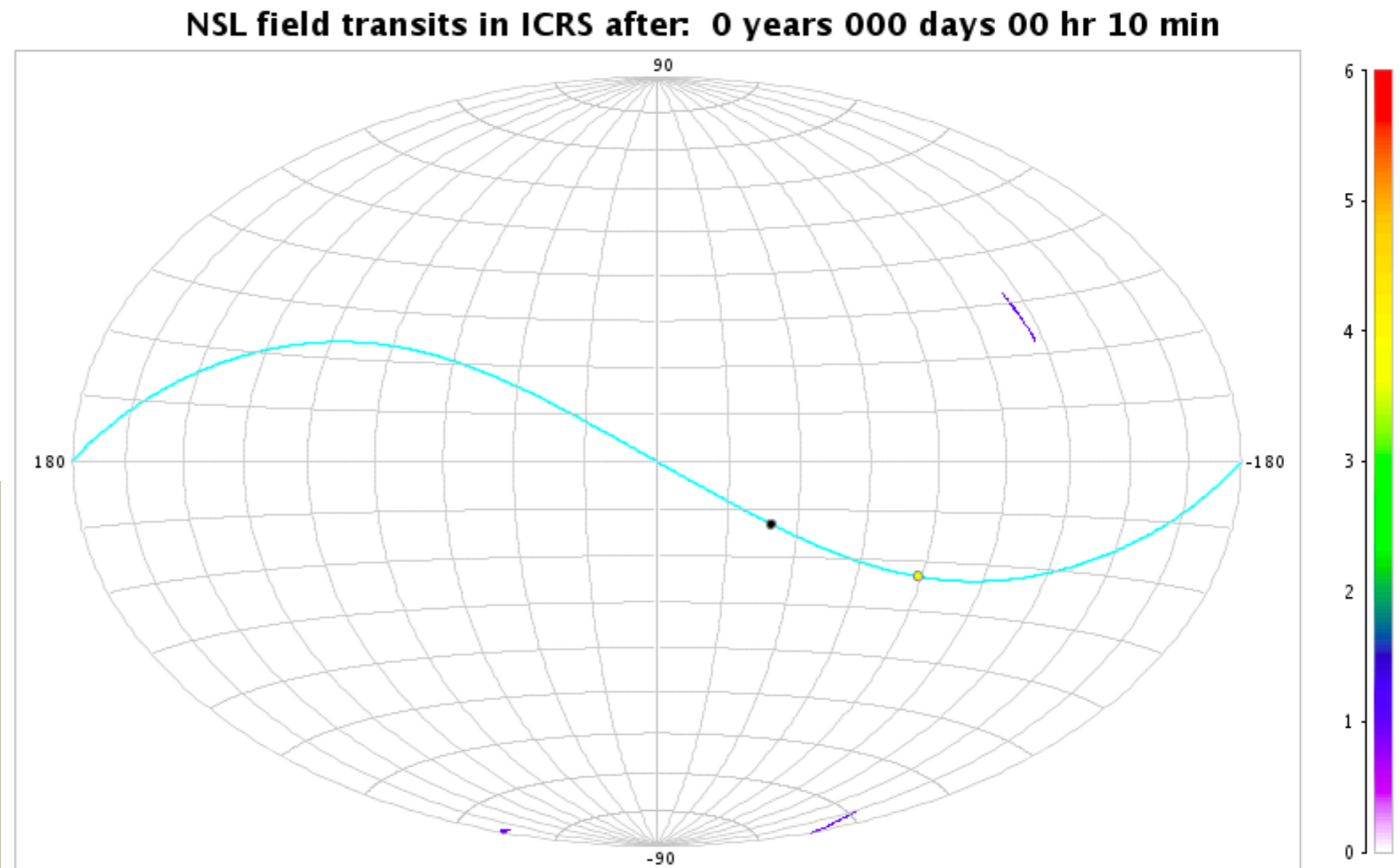
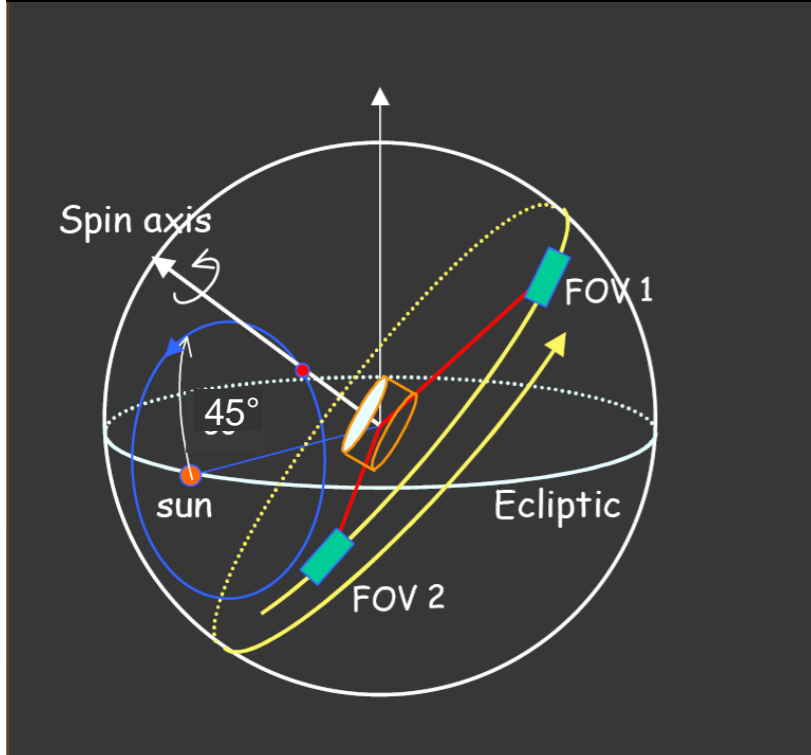
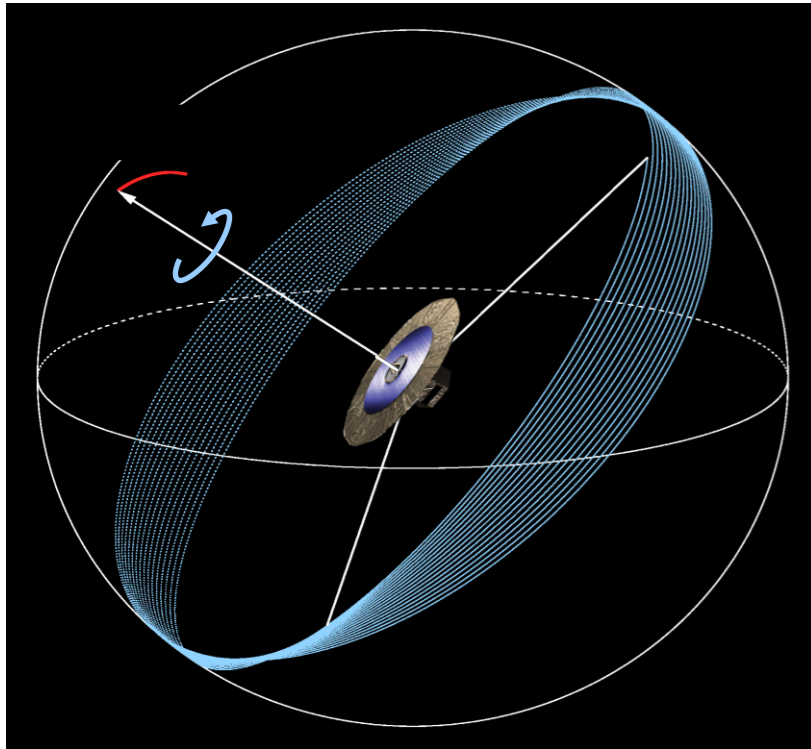
- Large mosaic
>100 CCDs
- TDI mode
- Binning
 - Reduced noise
photon count
 - 1D epoch
astrometry
 - Automatic detec^o
patch window
send to ground



Scanning law - Mapping



- *Scanning - no pointing*
- *2 FOVs - global astrom.*

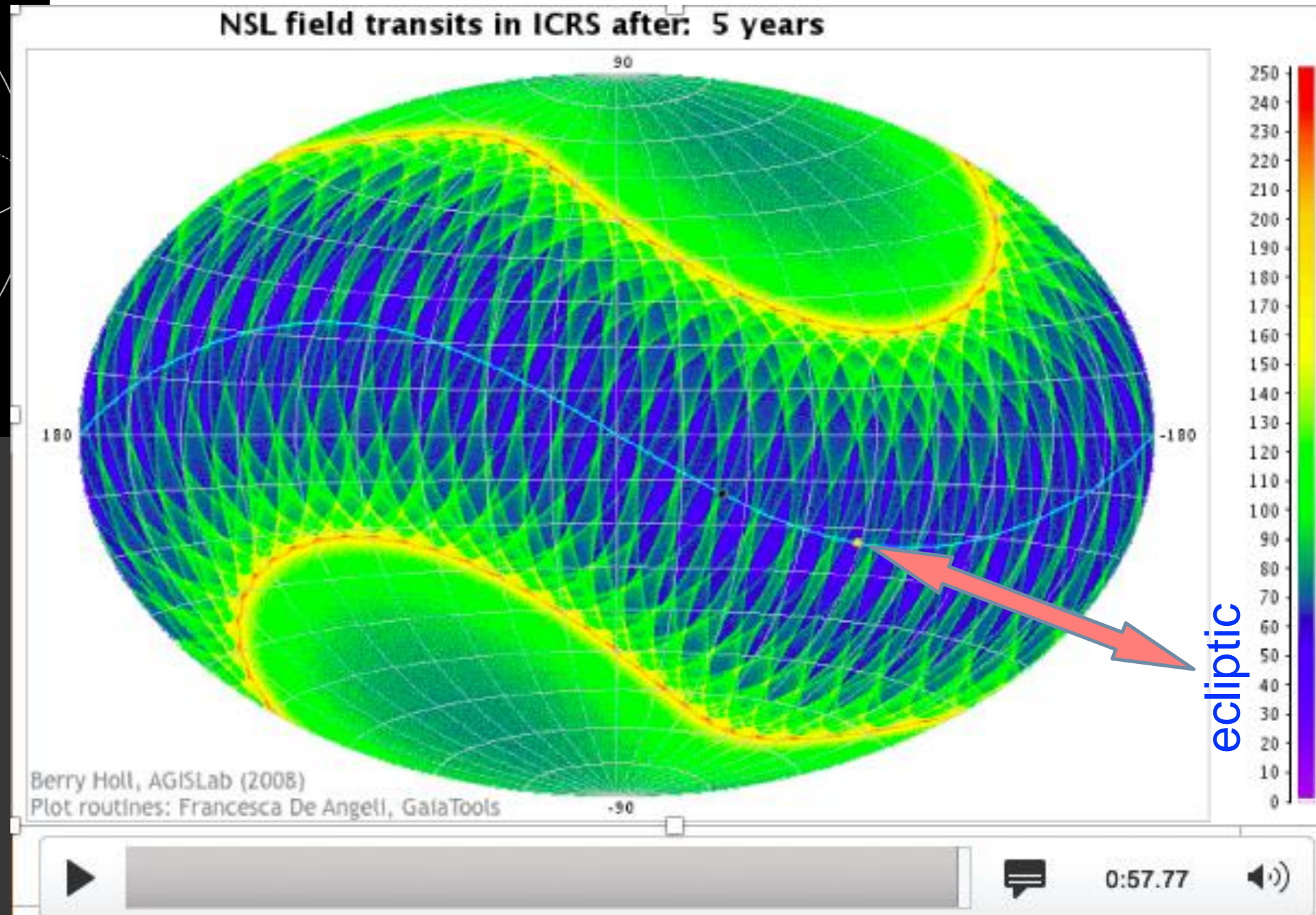
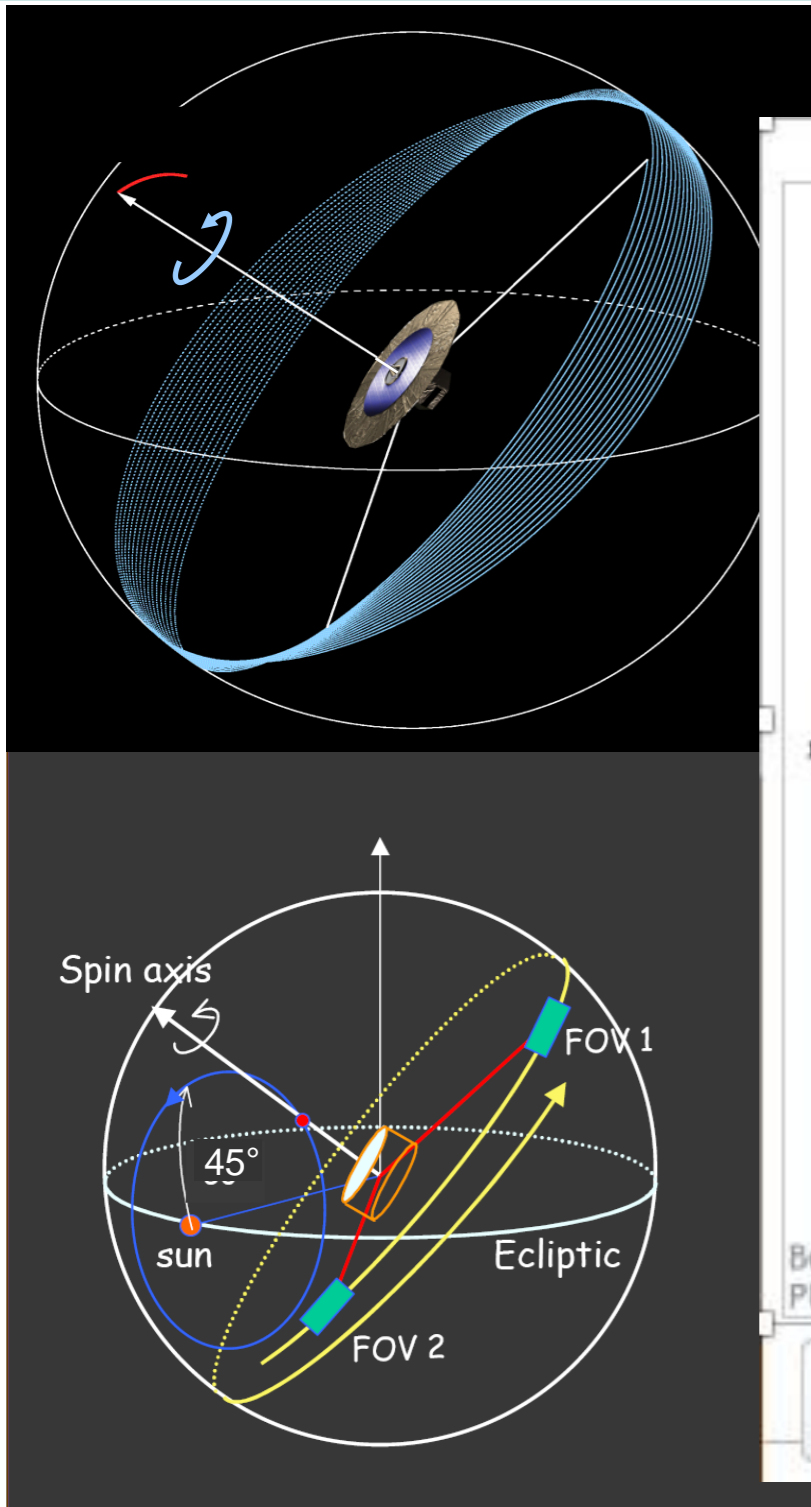


Scanning law - Mapping



- *Scanning - no pointing*
- *2 FOVs - global astrom.*

Movie showing sky coverage over time



Gaia 2nd data release DR2

– overview –

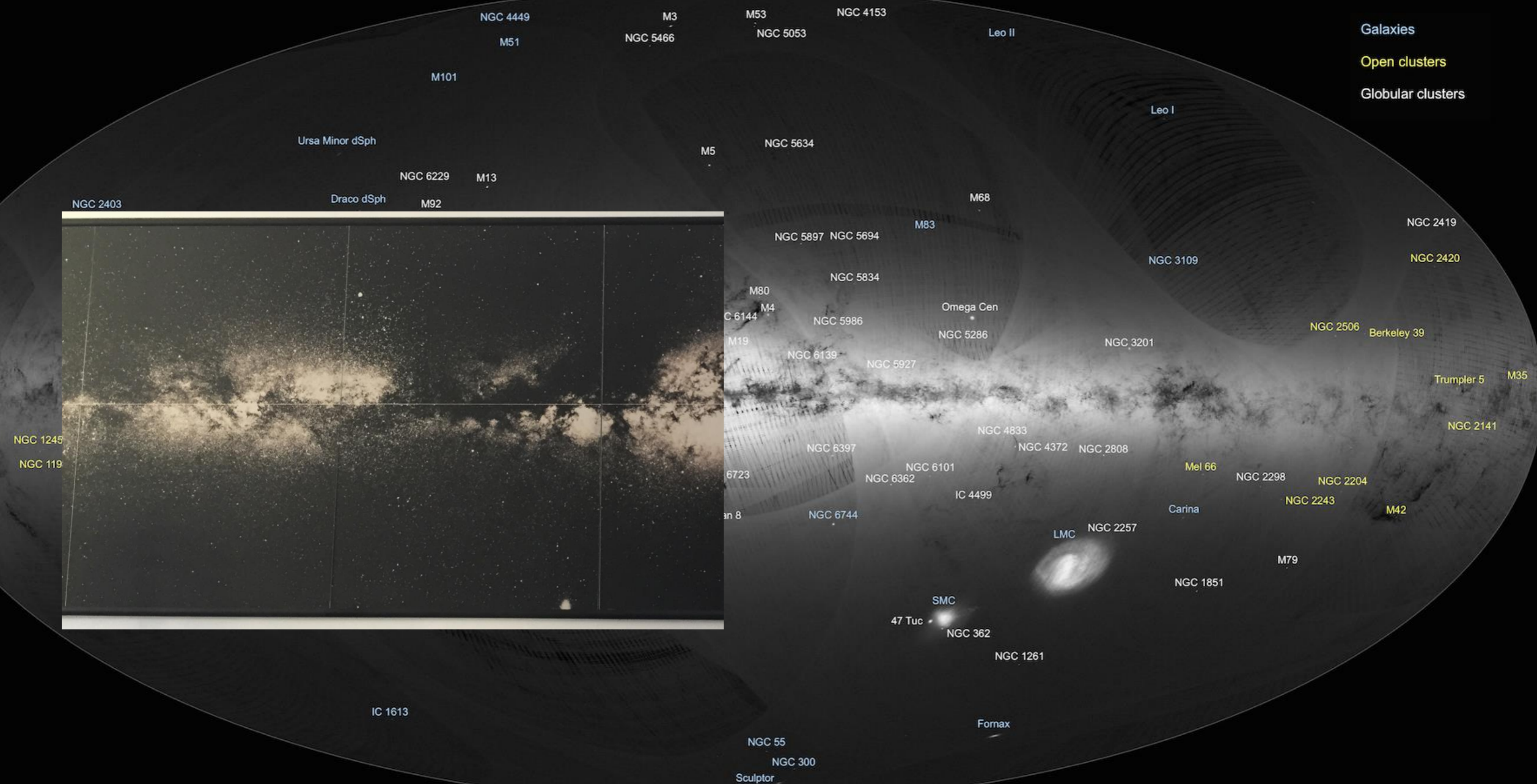


Gaia DR1 (2017) – a teaser

→ GAIA'S FIRST SKY MAP



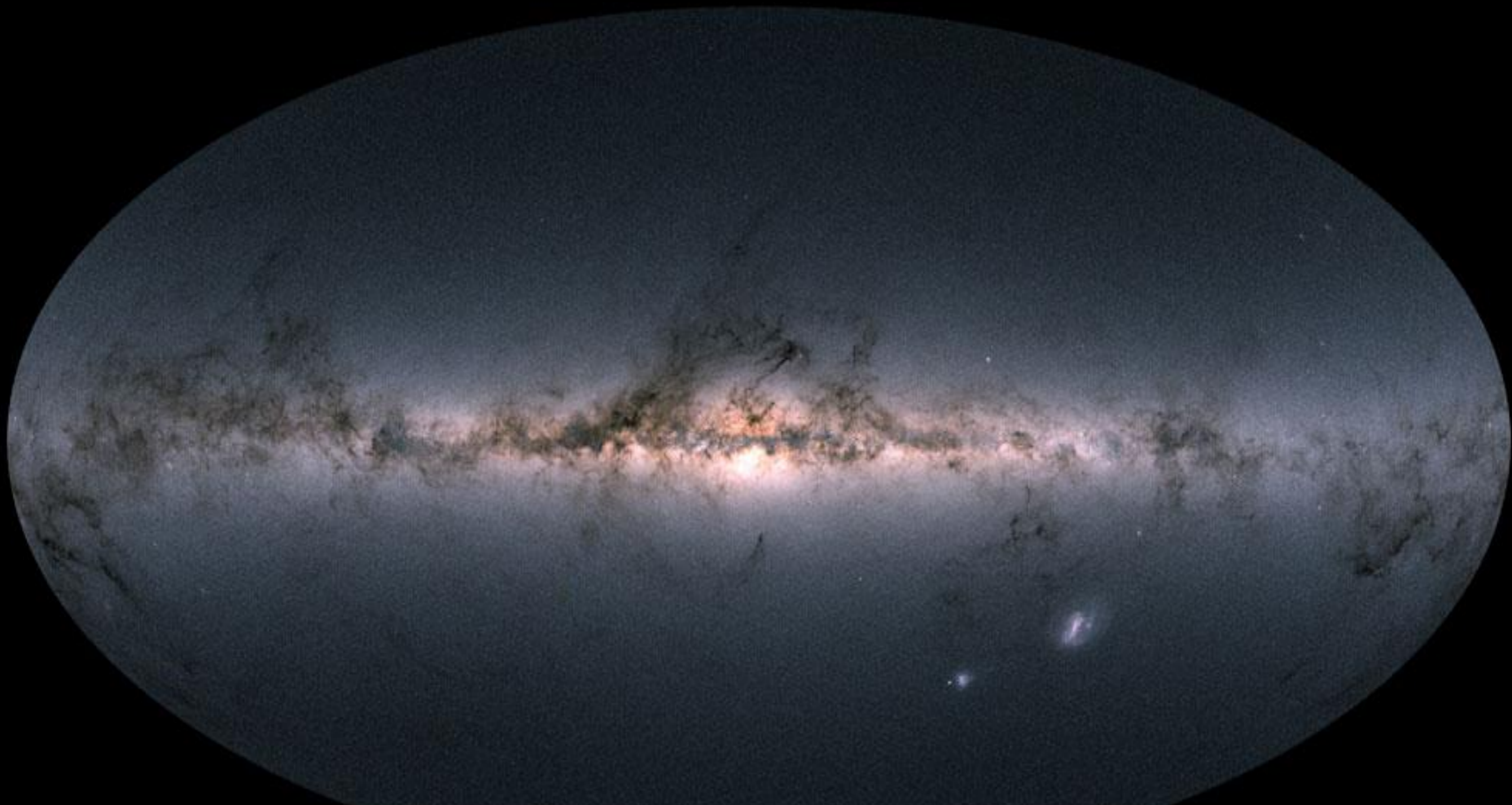
Galaxies
Open clusters
Globular clusters



Mostly positions & magnitudes, TGAS



Gaia DR2 – April 2018

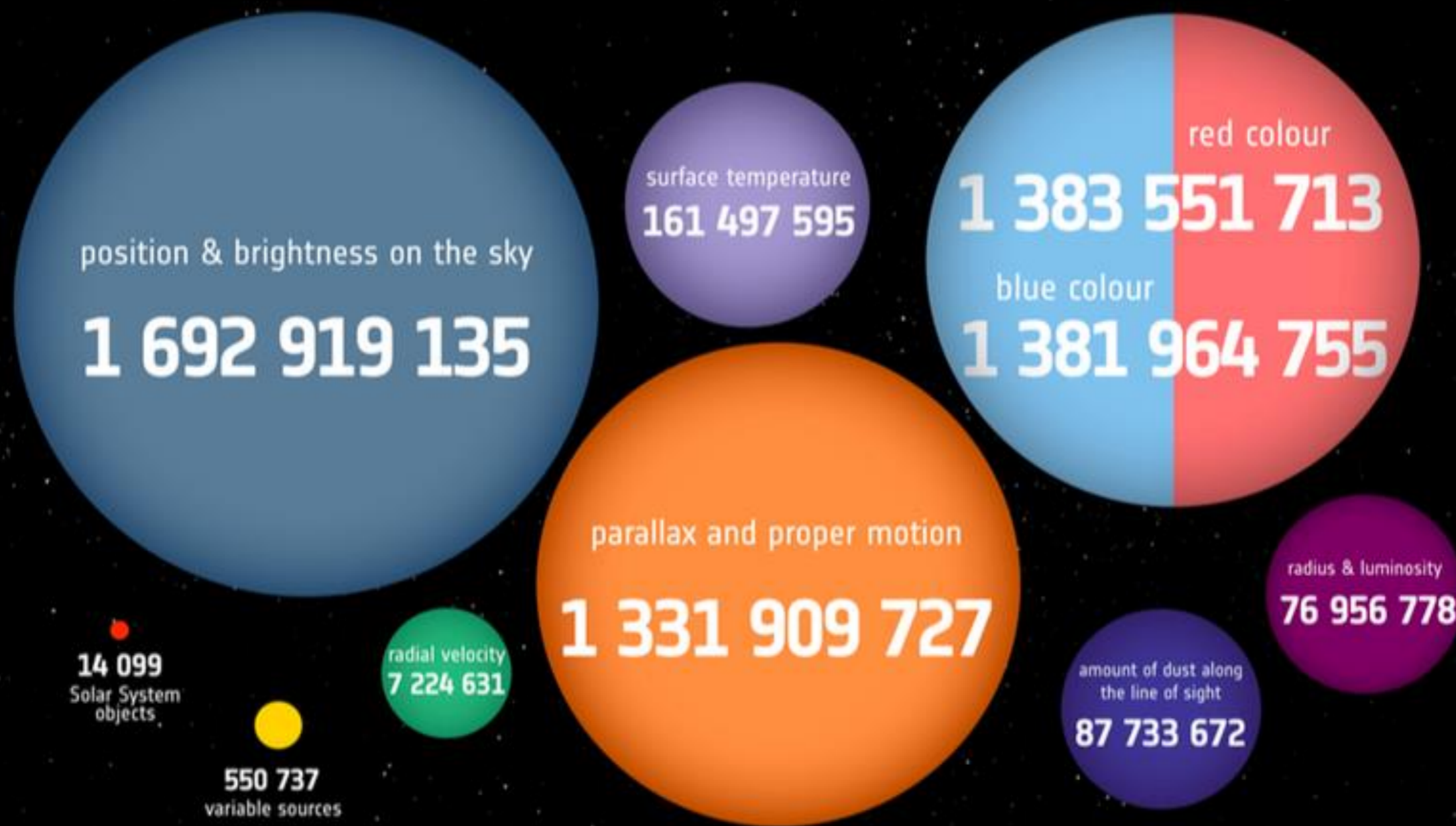


not only an image :: a map



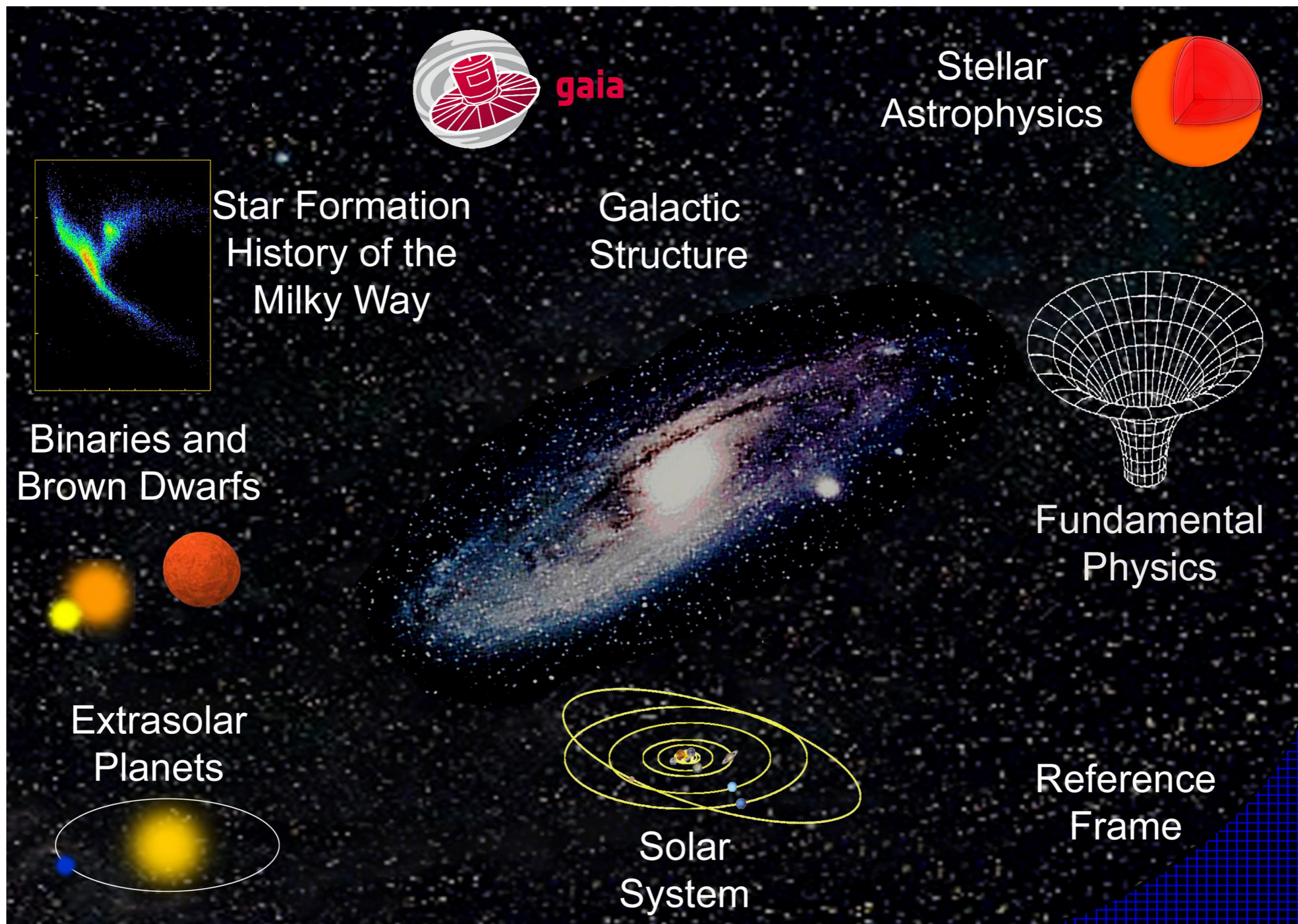
Gaia DR2 – in numbers

Over 22 months (2014.6 to 2016.4)



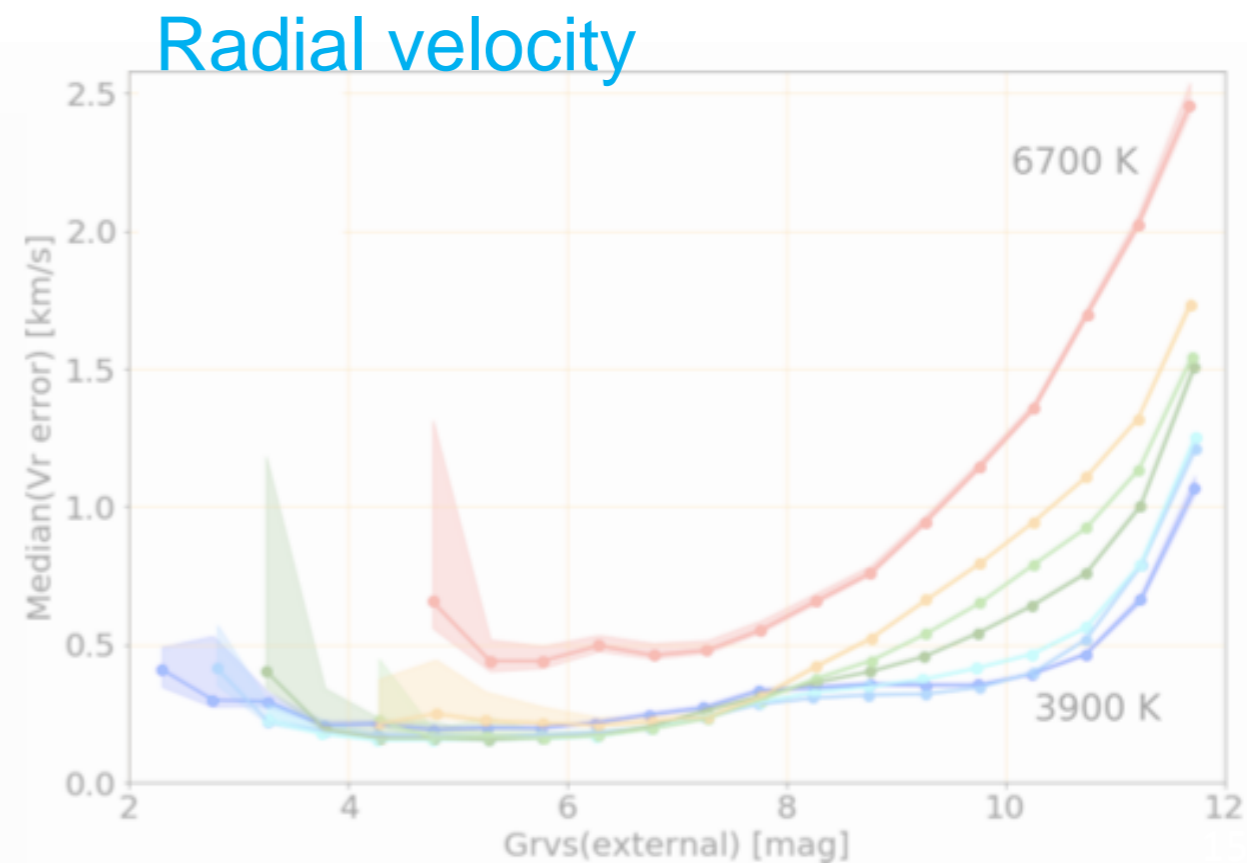
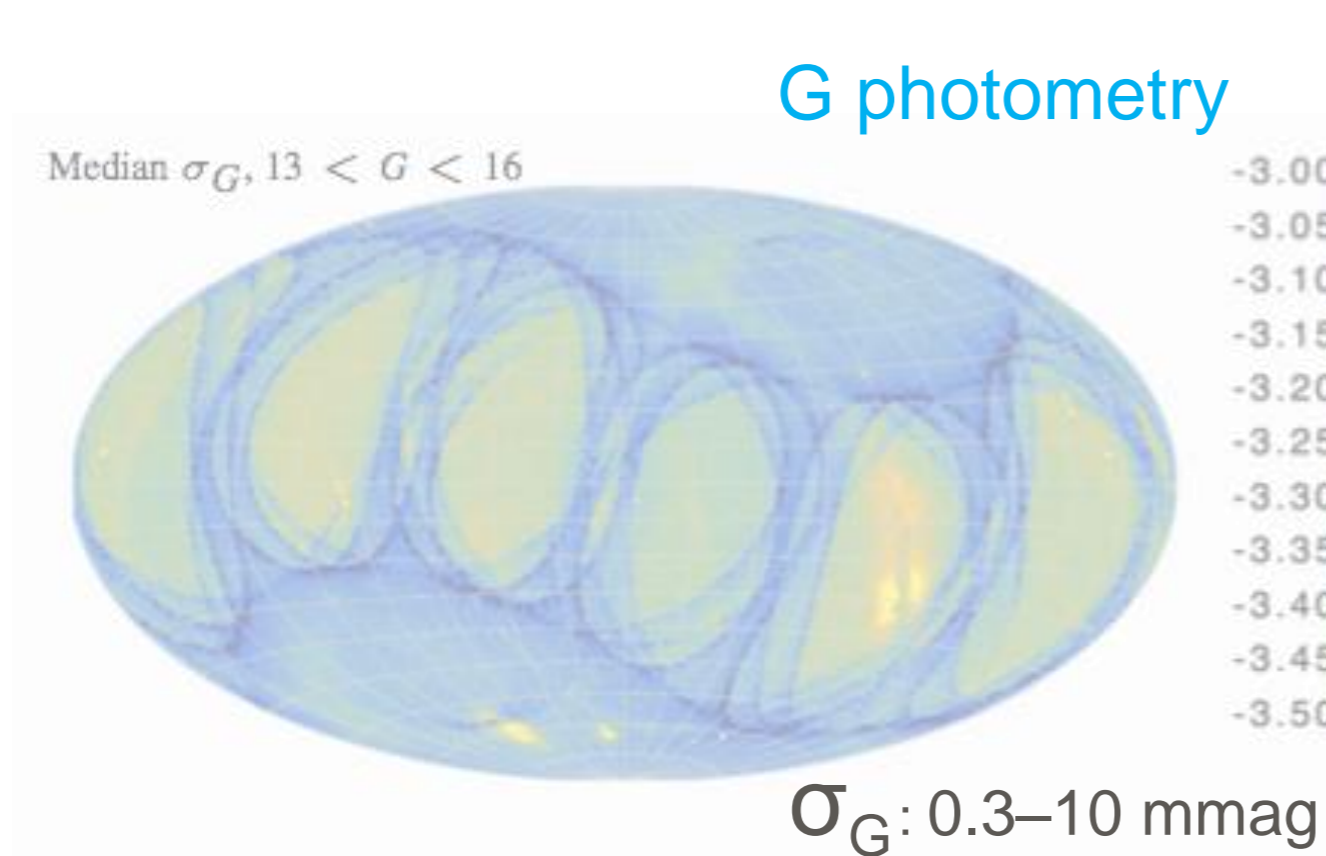
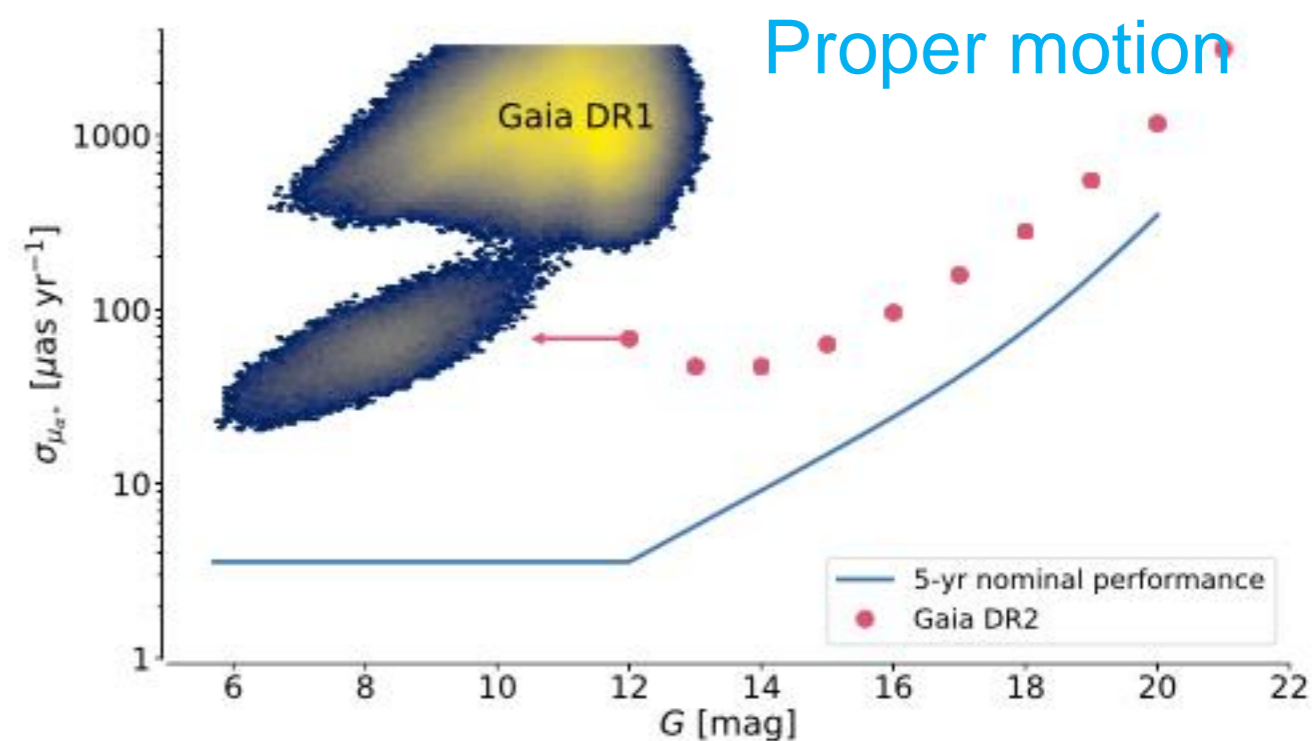
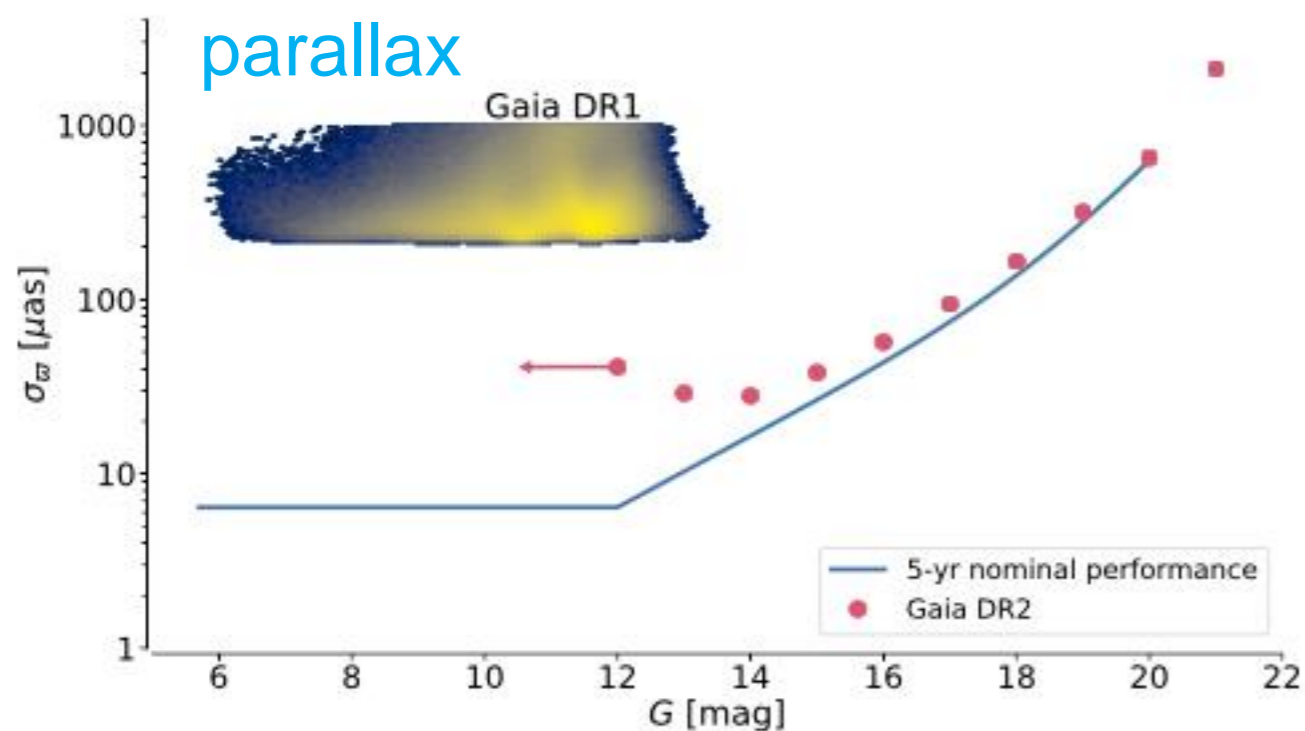


Gaia – Scientific Domains



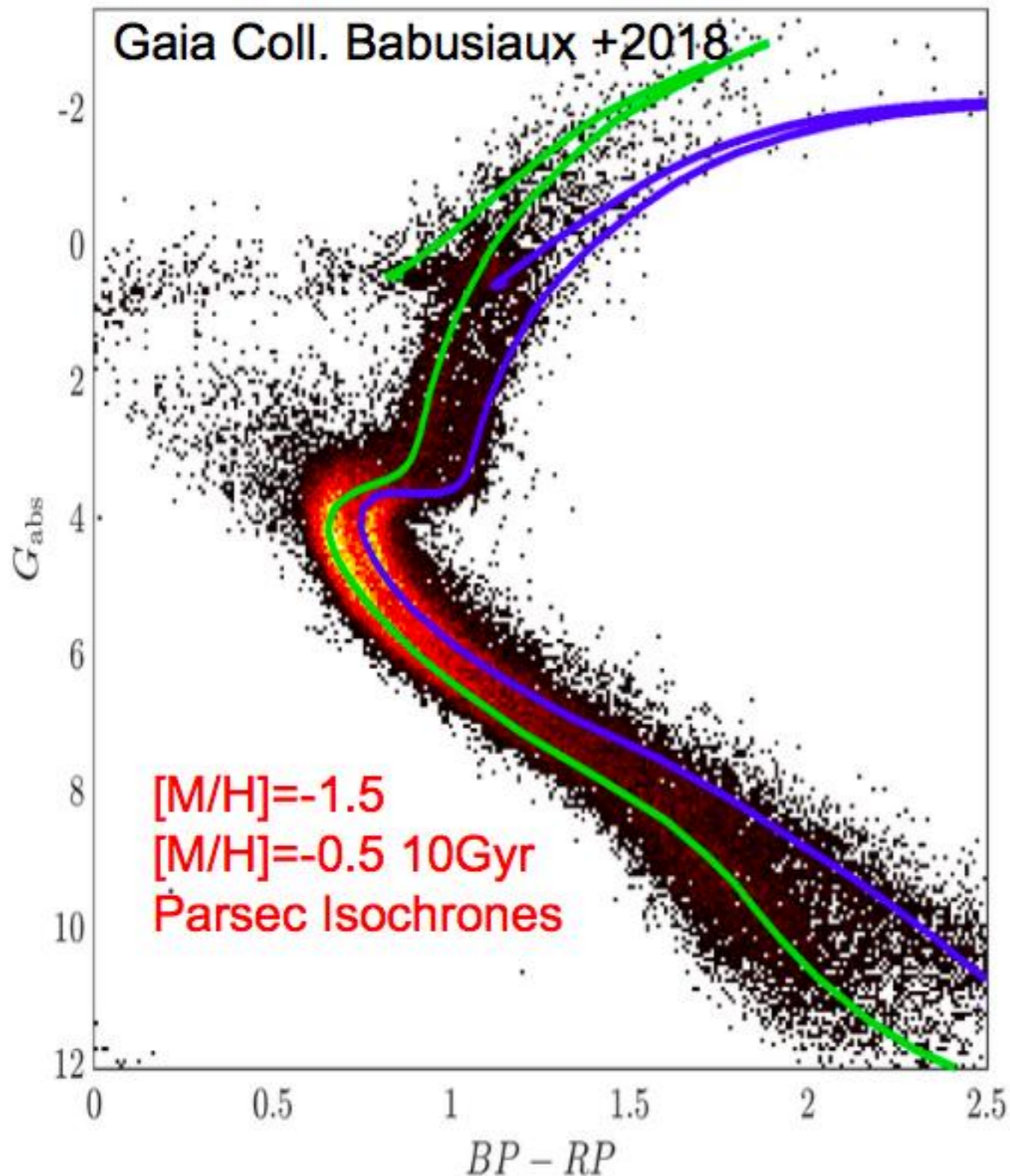


DR2 – Performances





Gaia DR2 – H-R Diagram



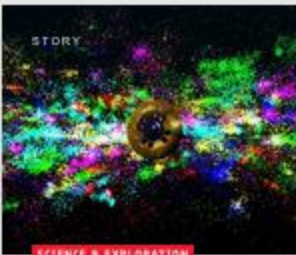
- Fine structures in H-R diagrams
- New views of
 - halo
 - Wdwarfs
 - clusters
 - mergers...
- Stellar evolution,
- Kinematics
- Extinction
- Combination with other surveys

(Babusiaux et al. 2018, A&A)

SCIENCE & EXPLORATION

Gaia

http://www.esa.int/Science_Exploration/Space_Science/Gaia/



SCIENCE & EXPLORATION

Gaia untangles the starry strings of the Milky Way

25/08/2019 8017 views 122 Links

READ



SCIENCE & EXPLORATION

Astronomers spy Europa blocking distant star - thanks to Ga...

25/07/2019 8393 views 119 Links

READ



SCIENCE & EXPLORATION

Gaia starts mapping our galaxy's bar

16/07/2019 8319 views 88 Links

READ



ENABLING & SUPPORT

Gaia's biggest operation since launch

15/07/2019 6433 views 140 Links

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SCIENCE & EXPLORATION

Observing Gaia from Earth to improve its star maps

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Gaia clocks new speeds for Milky Way-Andromeda collision

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Gaia reveals how Sun-like stars turn solid after their demi...

09/01/2019 11044 views 160 Links

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SCIENCE & EXPLORATION

Galactic ghosts: Gaia uncovers major event in the formation...

31/10/2018 10054 views 183 Links

READ



SCIENCE & EXPLORATION

Gaia spots stars flying between galaxies

02/10/2018 17873 views 180 Links

READ



SCIENCE & EXPLORATION

Gaia finds candidates for interstellar 'Oumuamua's home

25/09/2018 7032 views 107 Links

READ



SCIENCE & EXPLORATION

Gaia hints at our Galaxy's turbulent life

18/09/2018 12130 views 185 Links

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SCIENCE & EXPLORATION

Infant exoplanet weighed by Hipparcos and Gaia

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SCIENCE & EXPLORATION

Gaia creates richest star map of our Galaxy - and beyond

25/04/2018 30730 views 455 Links

READ



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Call for media: Second data release from ESA's Gaia mission

13/04/2018 3025 views 27 Links

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Close encounters of the stellar kind

31/08/2017 12897 views 111 Links

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SCIENCE & EXPLORATION

Artificial brain helps Gaia catch speeding stars

26/05/2017 7411 views 81 Links

READ



SCIENCE & EXPLORATION

The future of the Orion constellation



SCIENCE & EXPLORATION

Gaia's billion-star map hints at treasures to come



SCIENCE & EXPLORATION

Watch Gaia first data release media briefing



SCIENCE & EXPLORATION

Call for media: First data release from ESA's Gaia mission



SCIENCE & EXPLORATION

Celebrity comet spotted among Gaia's stars



SCIENCE & EXPLORATION

Gaia's first year of scientific observations



SCIENCE & EXPLORATION

Counting stars with Gaia



SCIENCE & EXPLORATION

Gaia discovers its first supernova



Gaia DR2 – in numbers

More than 2000 refereed papers
More than 300 arXiv preprints on the road ...

ADS Public Library

Gaia

Refereed Gaia papers since launch

Number of Papers:

2620

Date Created:

Apr 2 2019, 2:18pm

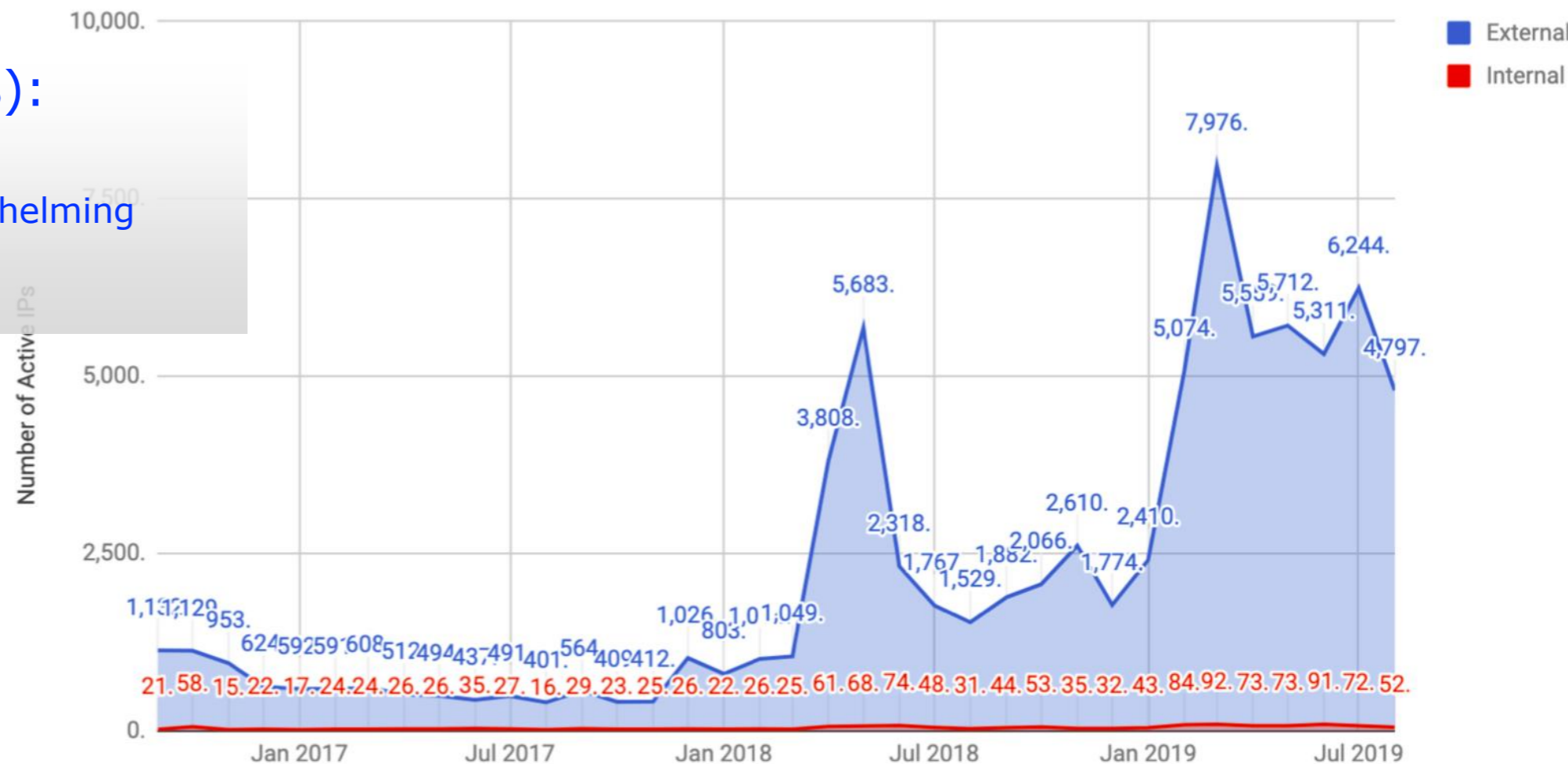
Date Last Modified:

Nov 24 2019, 9:43pm

T. Prusti (conclusions):

- Spacecraft is working
- Science production overwhelming

Archive Users (different IP Addresses) downloading data per month



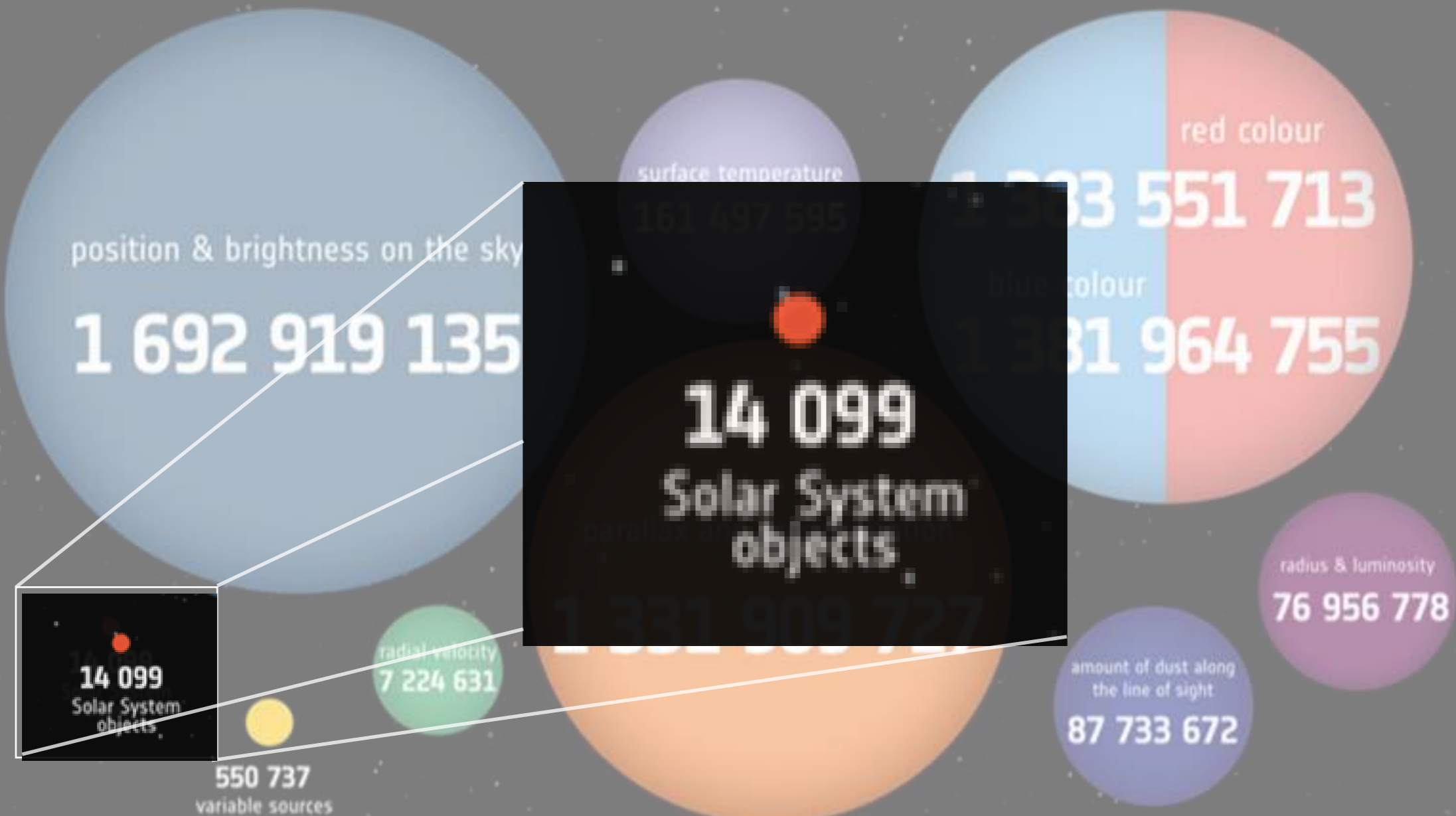
Solar System Objects +exoplanets

Treated in CU4 (a core unit of DPAC consortium)
With other peculiar objects
(NSS, extended objects, galaxies, gravitational lenses, etc.)



Gaia DR2

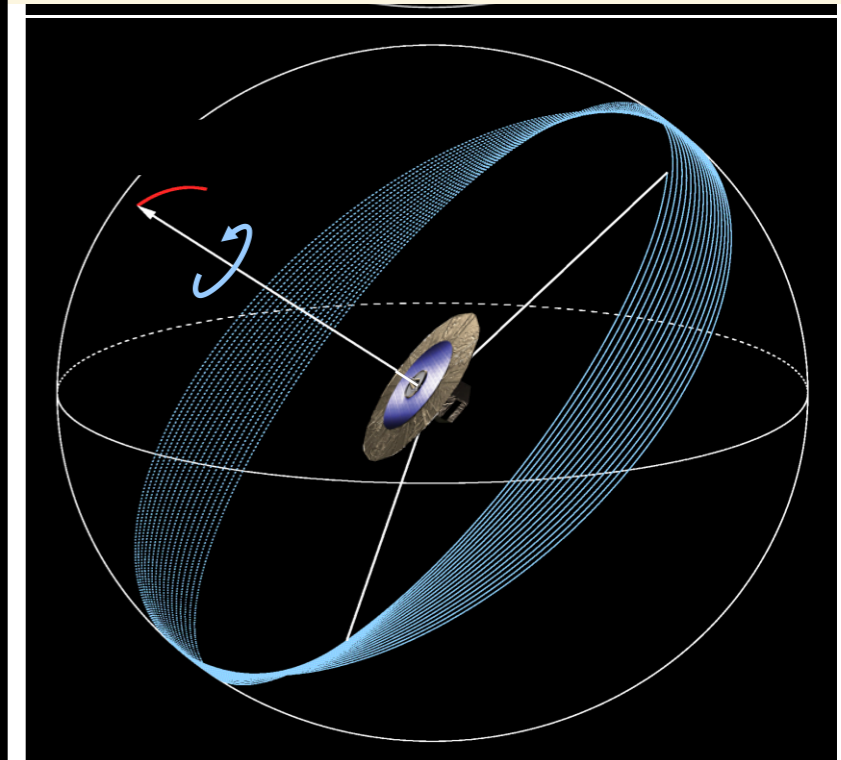
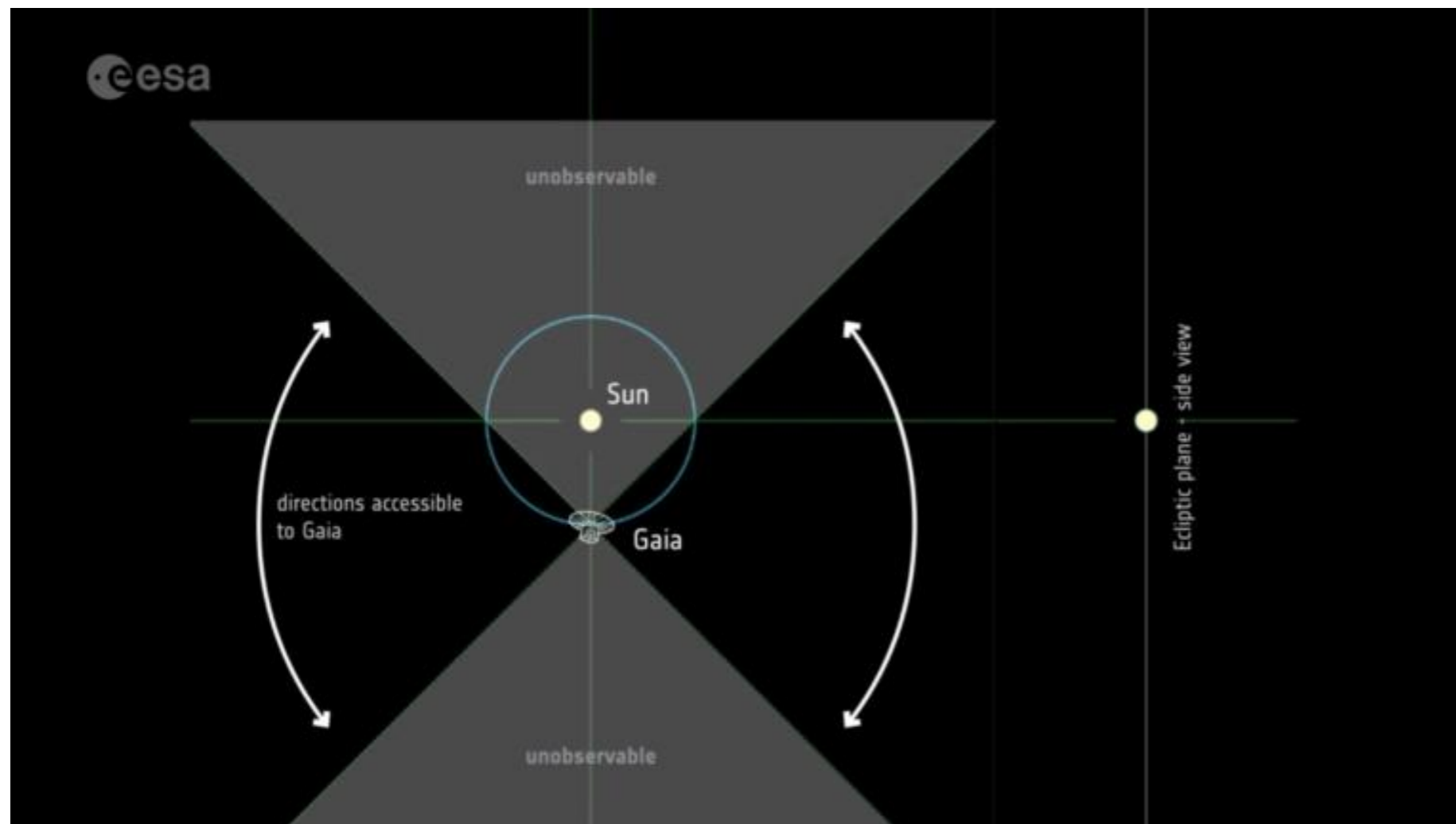
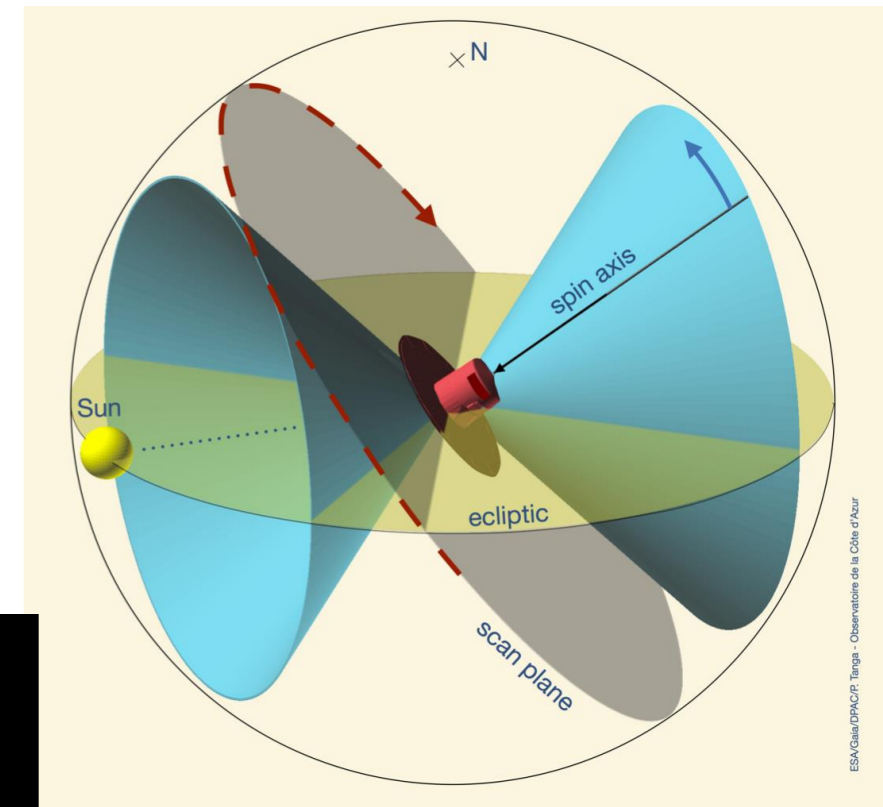
A teaser for SSOs





Gaia is observing SSOs

- Solar elongations 45° to 135° , with phase
- Limiting mag. $V \approx 20.7$
- 100,000 asteroid CCD obs. / day



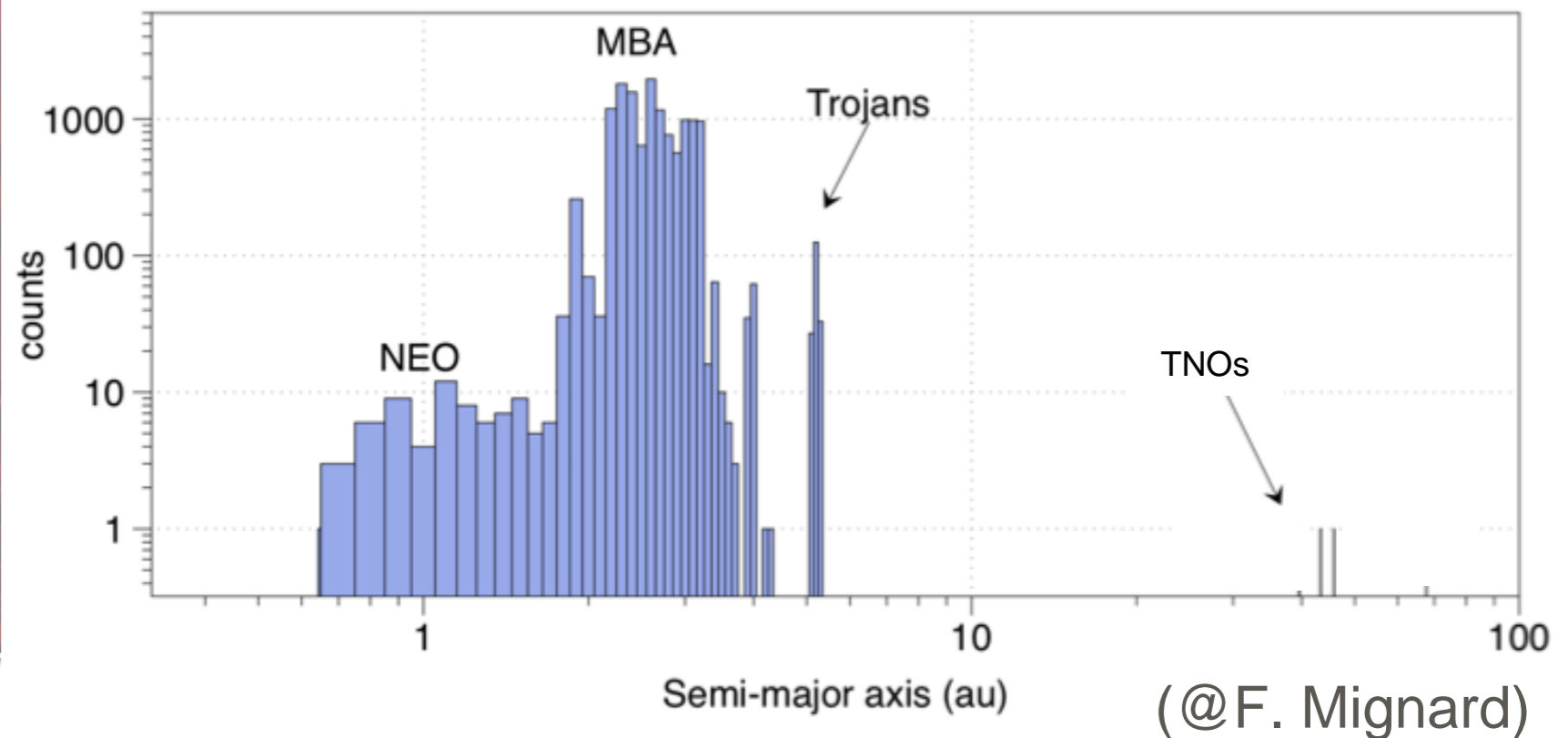
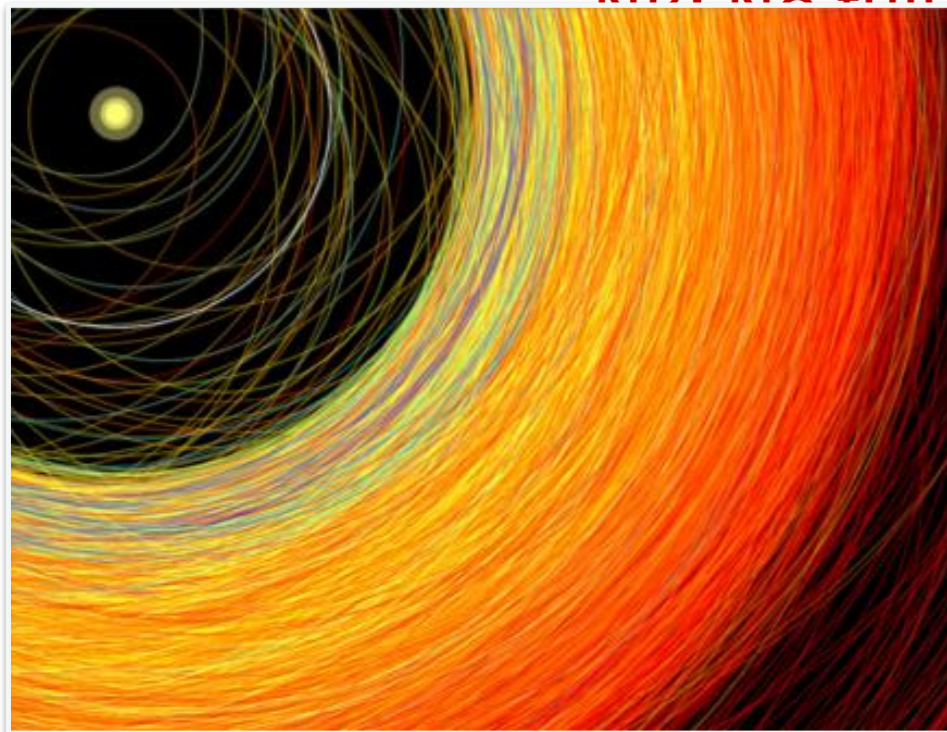


Gaia DR2 - Solar System

On the base of a *pre-selected* list of *known* objects (over $\approx 350,000$)
> 10 FOV transits over the 22 months of DR2, $V > 10.5$,

...

ND, NG, full



Objects

14 099 (*asteroids*)

Epoch astrometry

1 997 702 CCD positions

287 904 transits (52% : photometry)

Accuracy

0.3–8 mas (*along scan*)



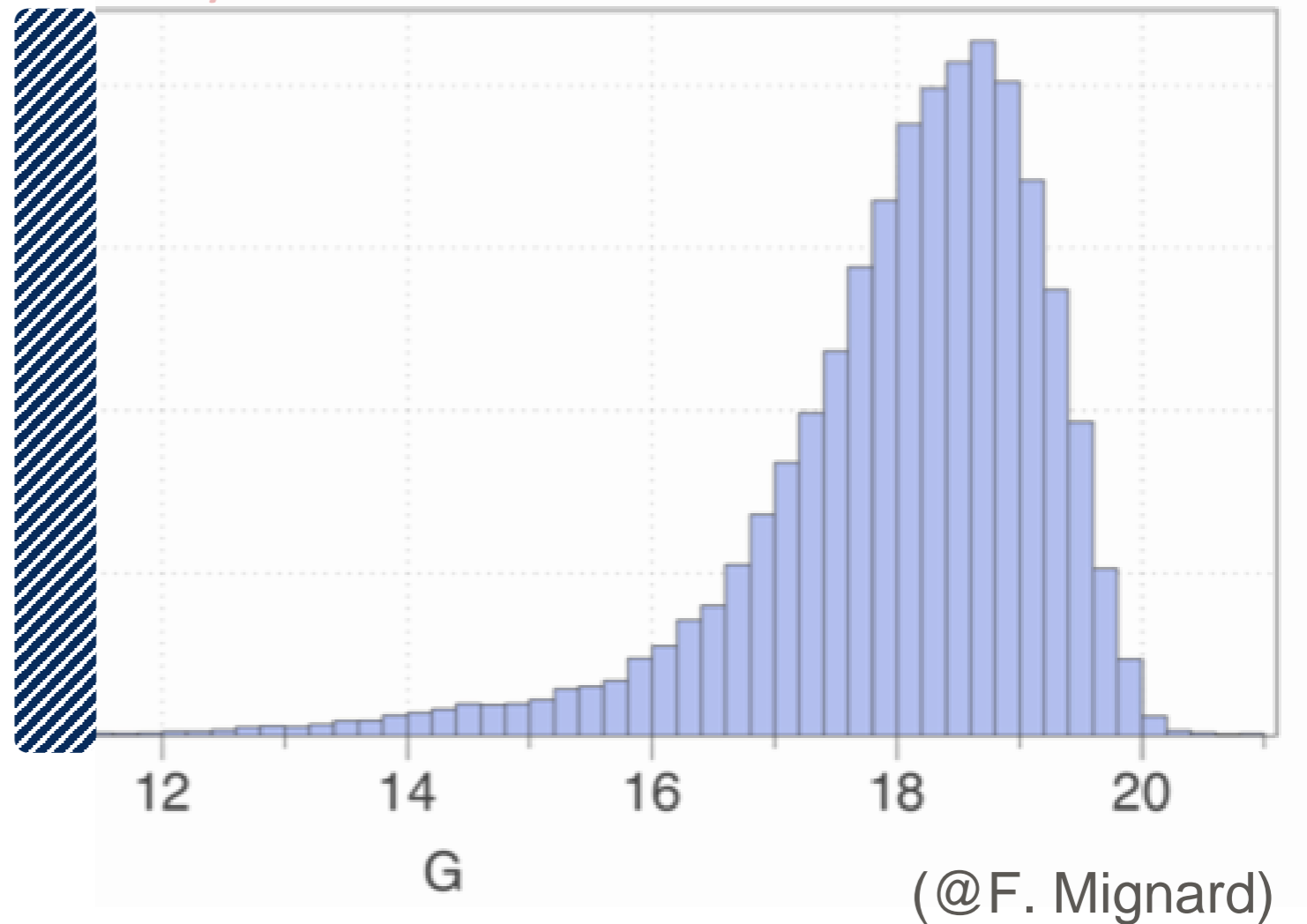
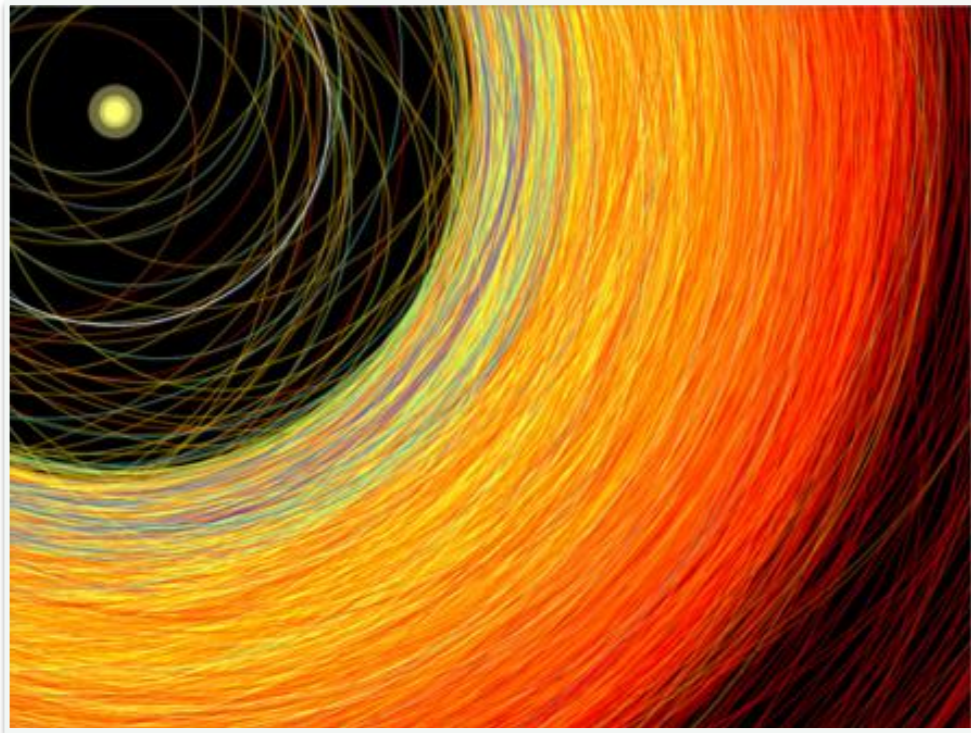
Gaia DR2 - Solar System

On the base of a *pre-selected* list of *known* objects

> 10 FOV transits over the 22 months of DR2, $V > 10.5$,

...

NB: No full calibration yet !!





(some) Results

- **Catalogue of stars DR1, DR2**
- **Direct observations**
- **Combination Ground-based + Gaia**



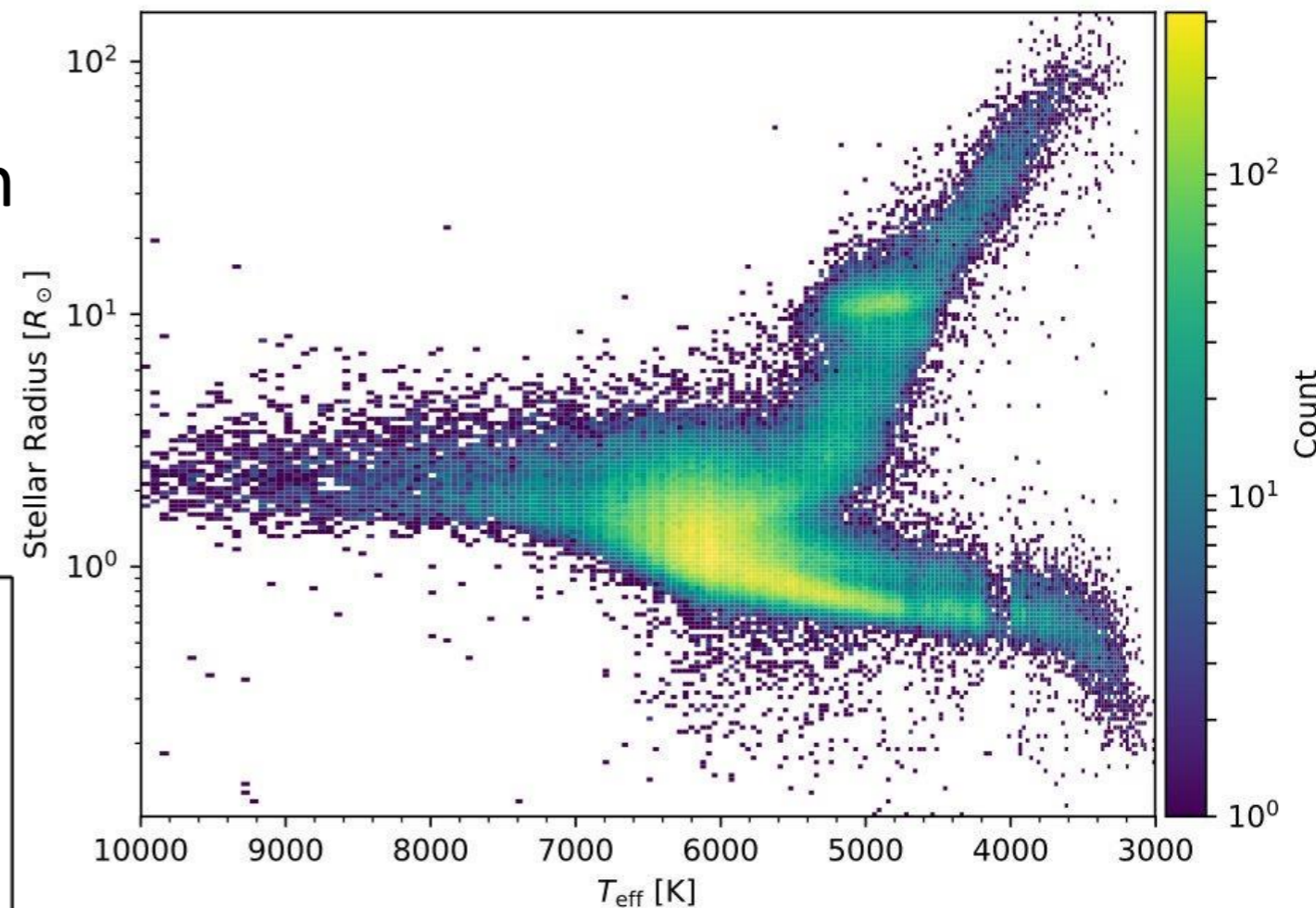
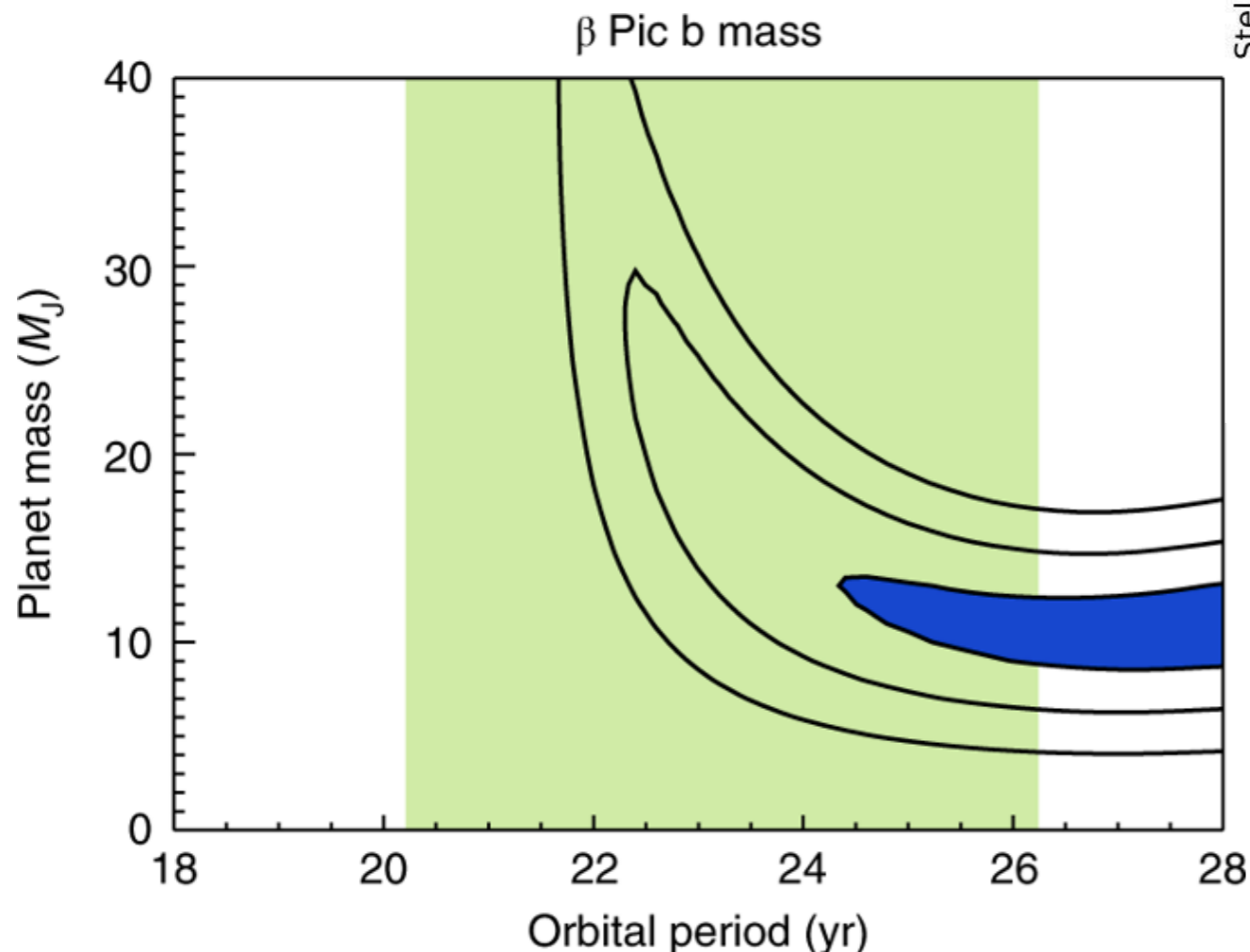
Exoplanets

(@A. Vallenari)

- No exoplanets yet, but... host stars

- Stellar radii + mass for exoplanet characterisation

(Berger 2018 ApJ)

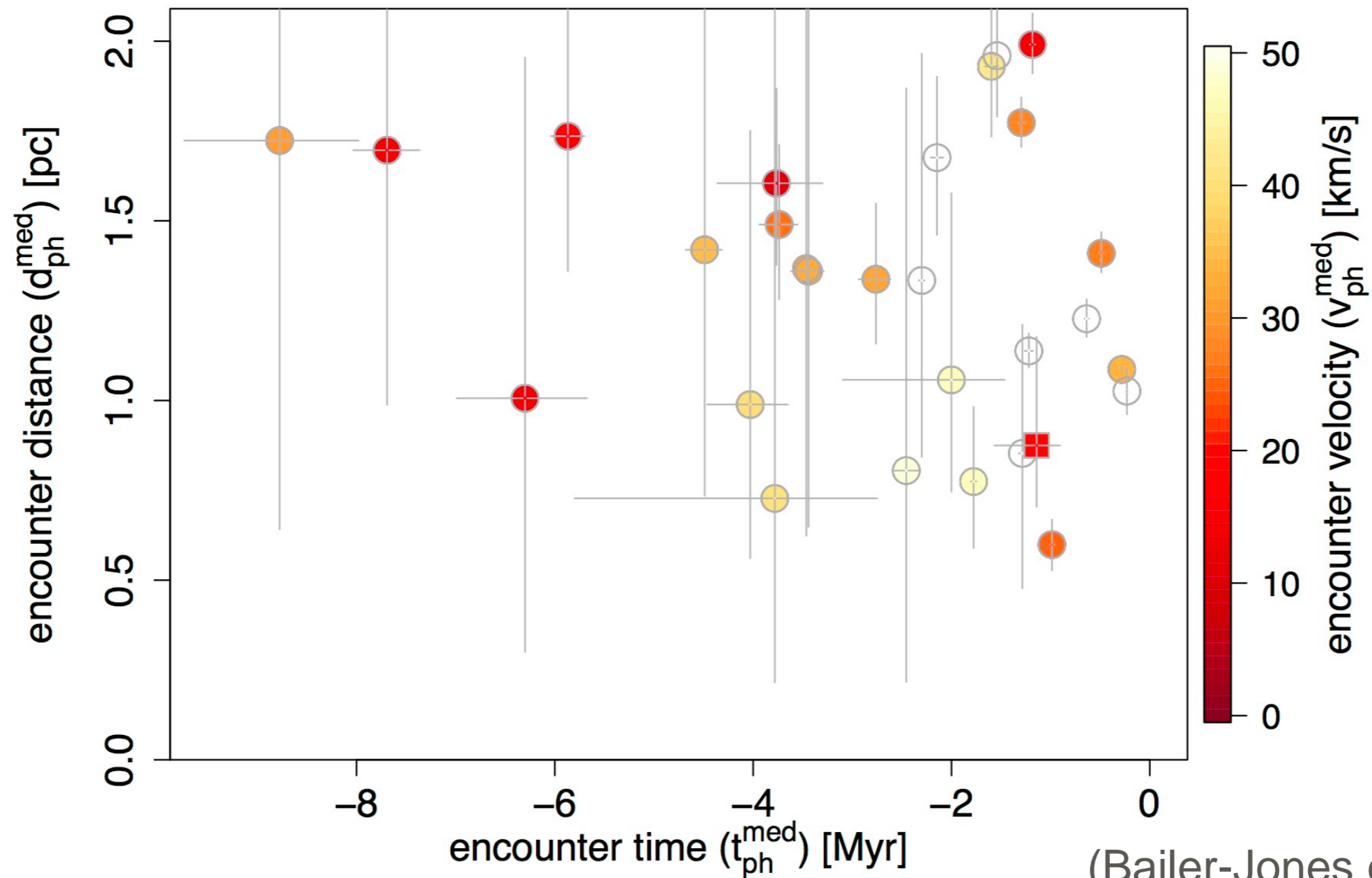


- Gaia + Hipparcos
Mass of planet β Pic b

(Snellen & Brown 2018 Nature)



'Oumuamua ISO's home



(Bailer-Jones et al. 2018 AJ)

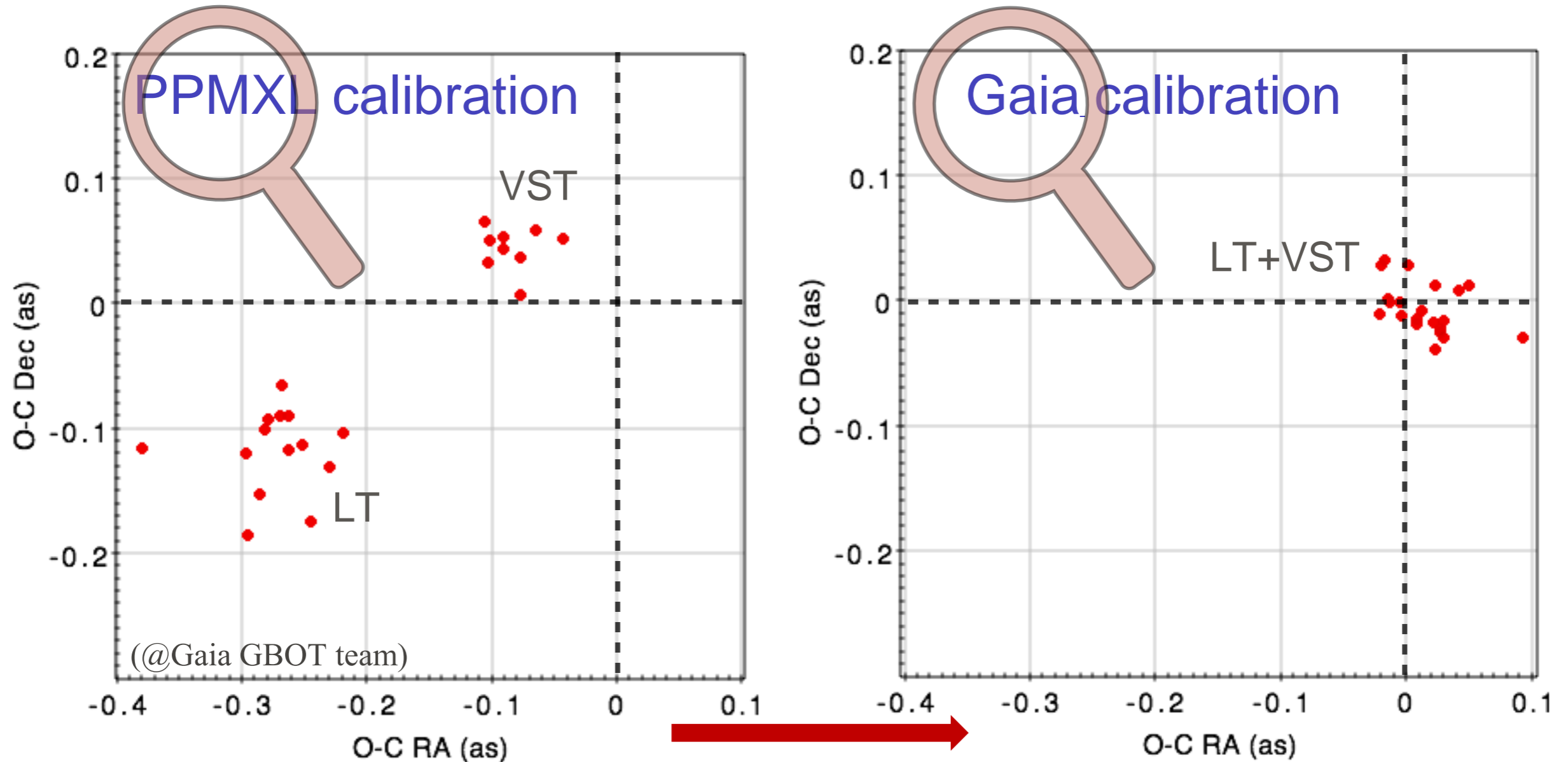
Other stellar encounters at origin of LPC comets from Oort cloud



Gaia DR2 catalogue - calibration

Asteroid (1132) Hollandia **ground-based observations** (GBOT)

Liverpool Telescope + VST (8 hours apart) & MPC ground-based data (~1900 positions)



Use of DR2 stellar catalogue
reduces errors



Gaia DR2 catalogue - calibration

Many programmes NEOs, TNOs, ...

Also for planetary satellites

(Wang et al. 2017 IAU's ; @Q.Peng Jinan univ.)

◇ DRI
◇ UCAC4

GDC

Geometric Distorsion Correction

Triton

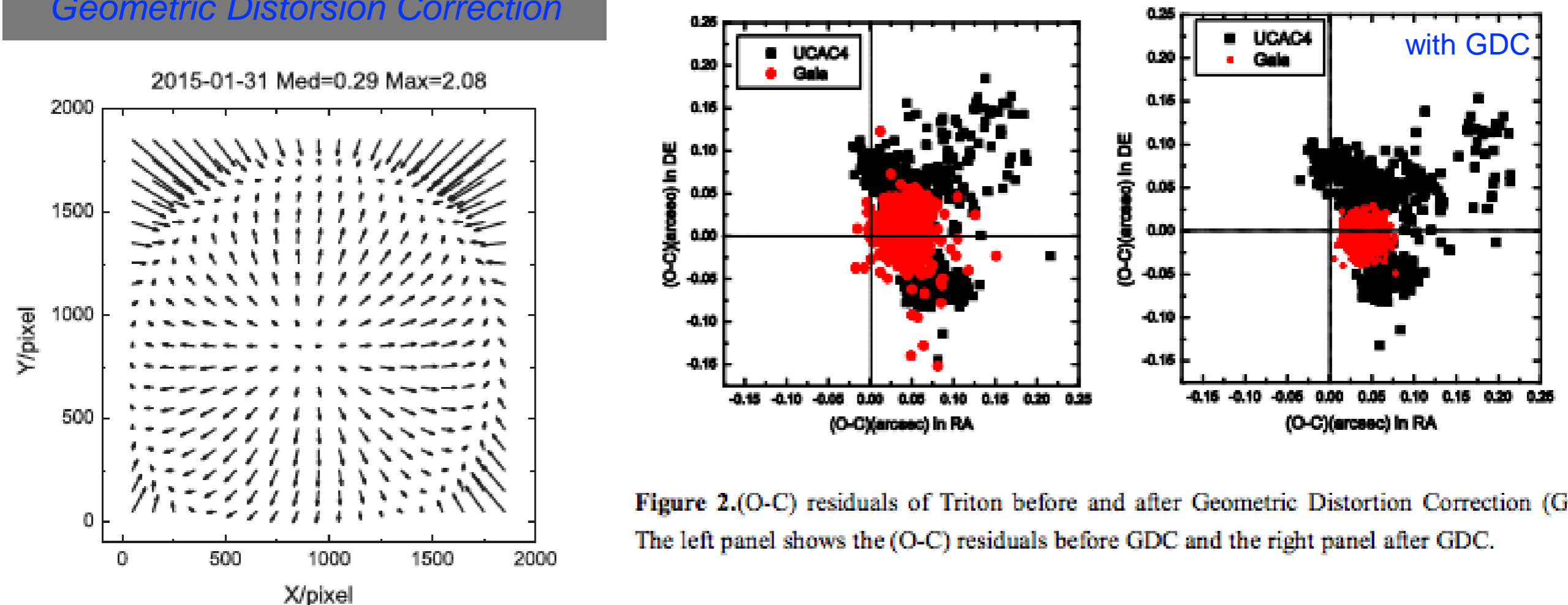
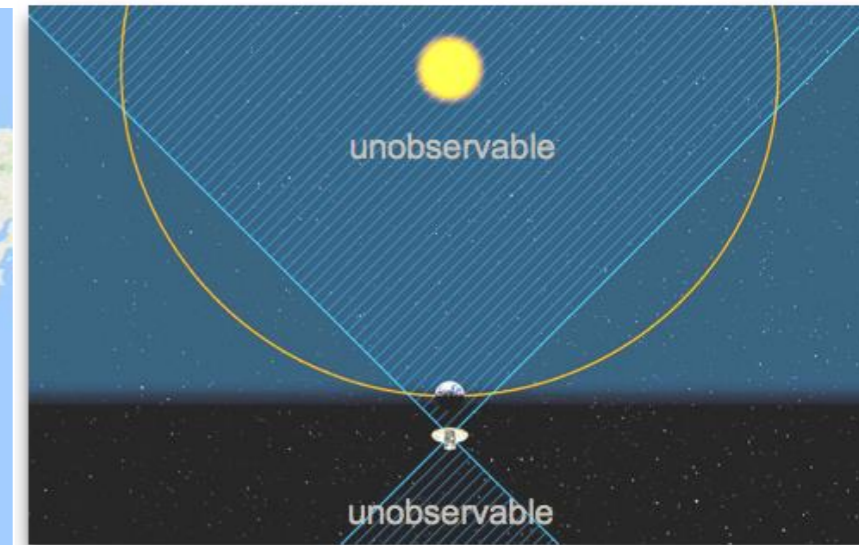


Figure 2.(O-C) residuals of Triton before and after Geometric Distortion Correction (GDC). The left panel shows the (O-C) residuals before GDC and the right panel after GDC.

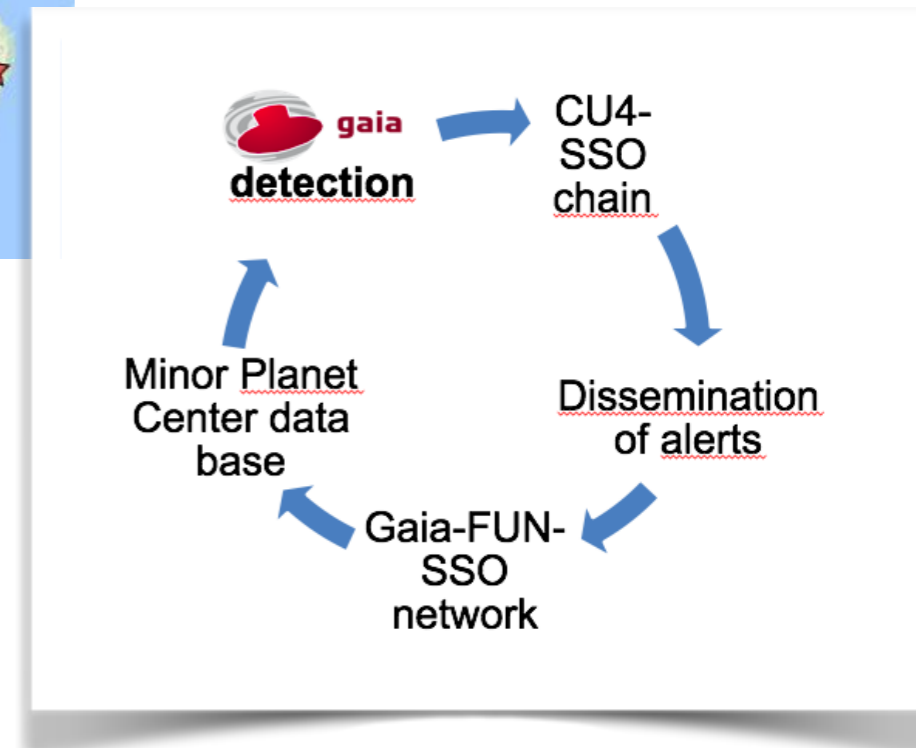


Alerts Gaia-FUN-SSO

(@W. Thuillot)



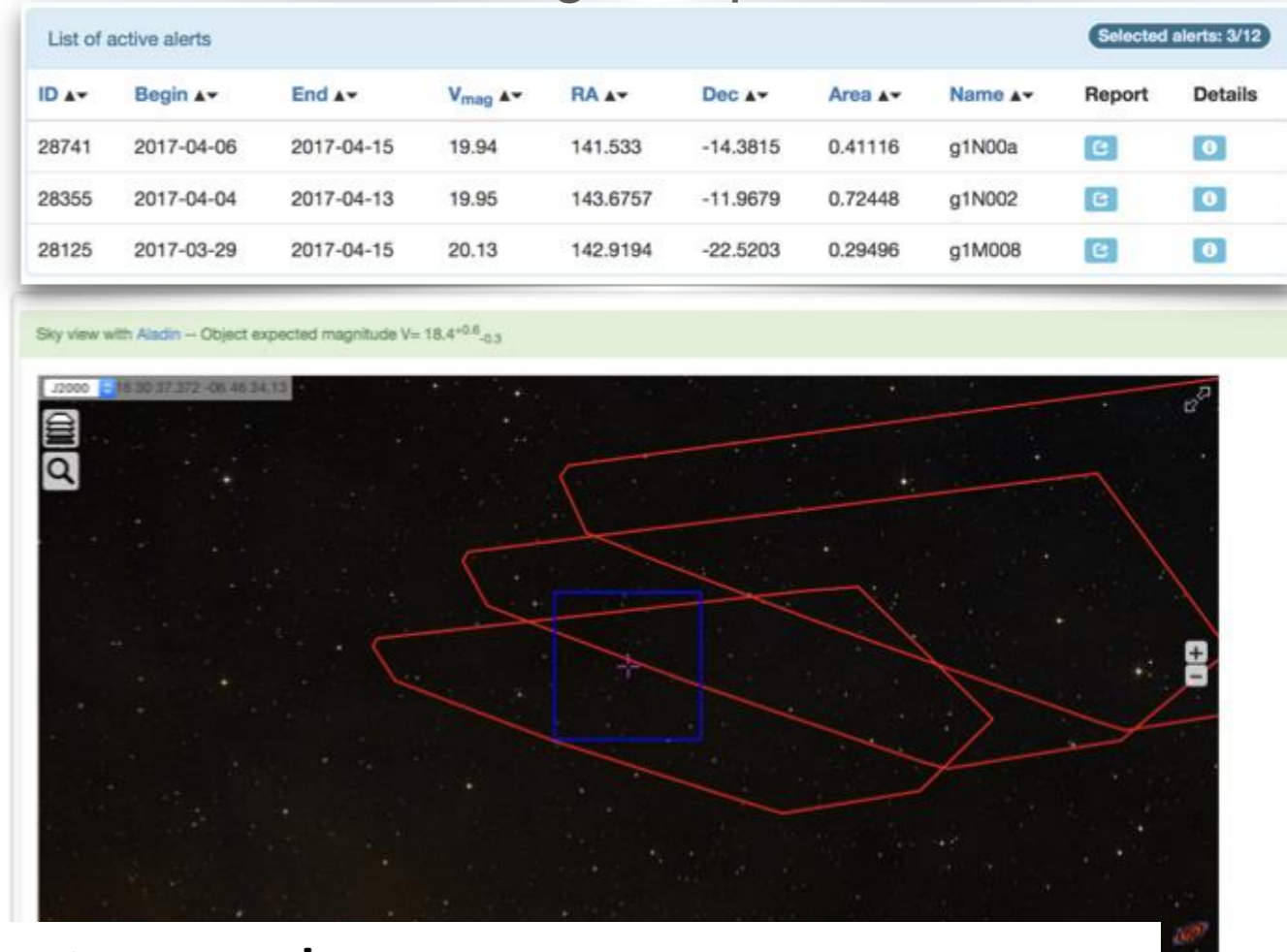
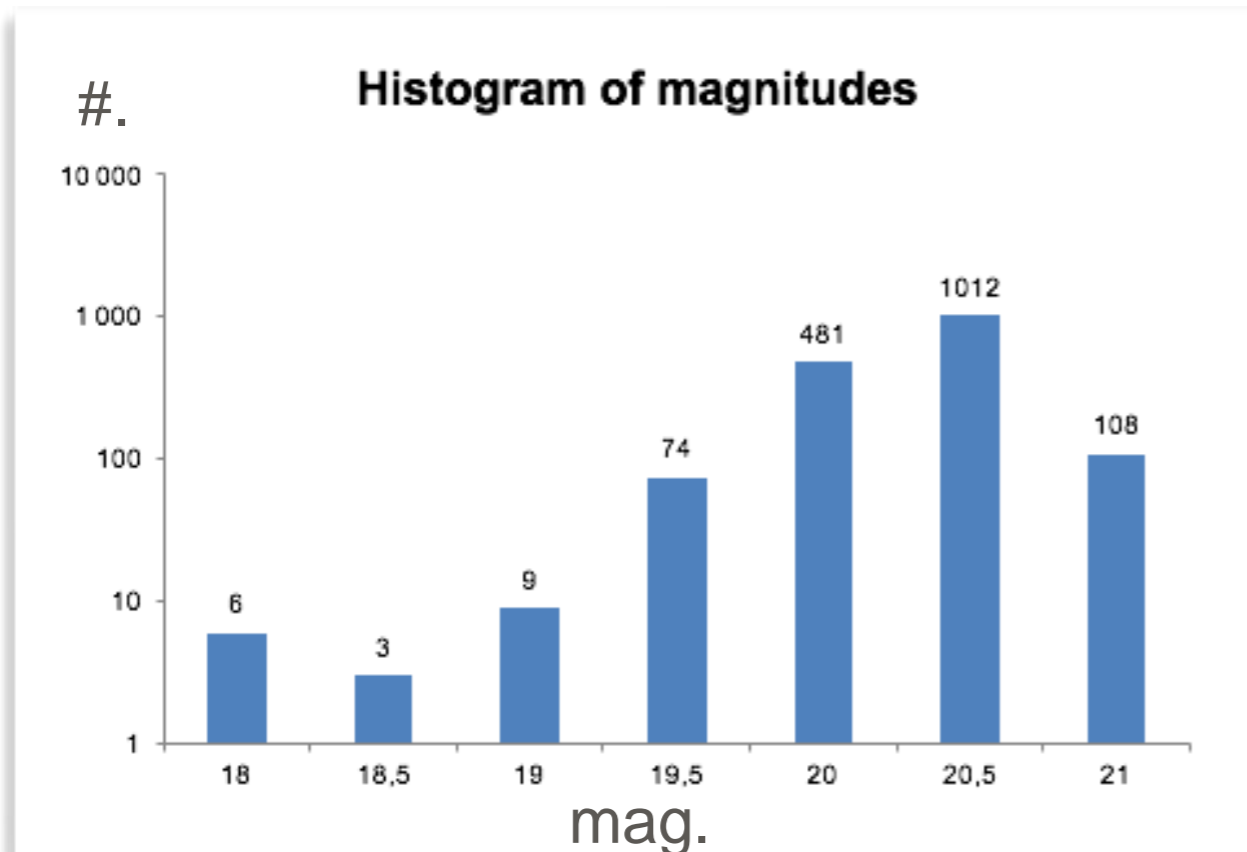
- Discovering new NEOs and SSOs
- Ground-based Follow-Up Network
 - Gaia scans at low elongations ($\approx 45^\circ$)
 - avoid losing newly discovered objects
 - short term chain with MPC





Alerts Gaia-FUN-SSO

alerts logs & ephemeris



- Ground-based Follow-Up Network
 - **i daily alerts posted!** $\approx 5/\text{day}$
 - faint, mag. 20.5 >T1m telescopes wanted
 - contact gaia-fun-sso@imcce.fr

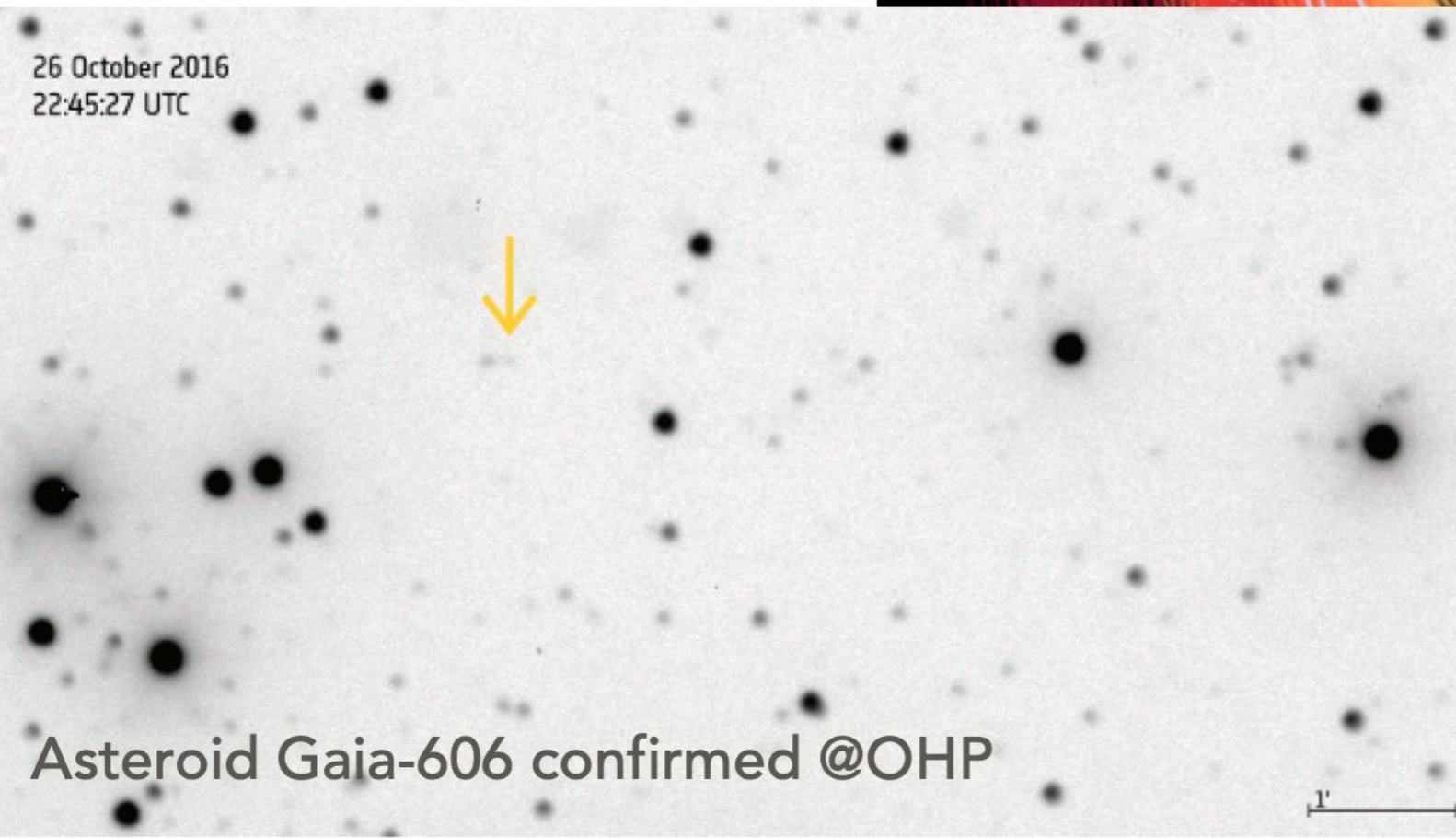
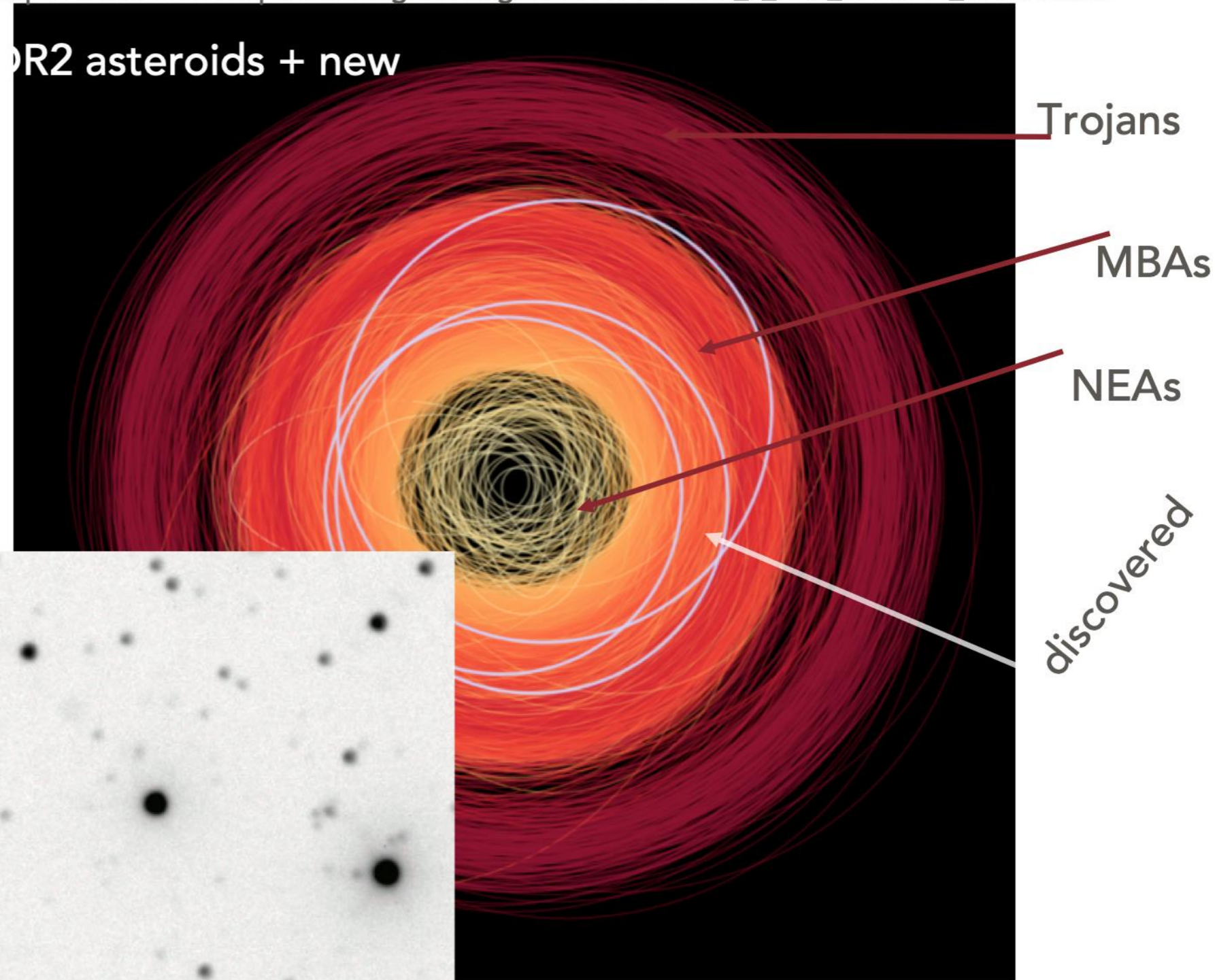


Alerts Gaia-FUN-SSO

http://www.esa.int/spaceinimages/Images/2019/04/Gaia_s_first_asteroid_discoveries

OR2 asteroids + new

- Discovery of asteroids
- MBAs, mag >20
- high inclination >15°



Asteroid Gaia-606 confirmed @OHP



Scientific Objectives

– briefly –

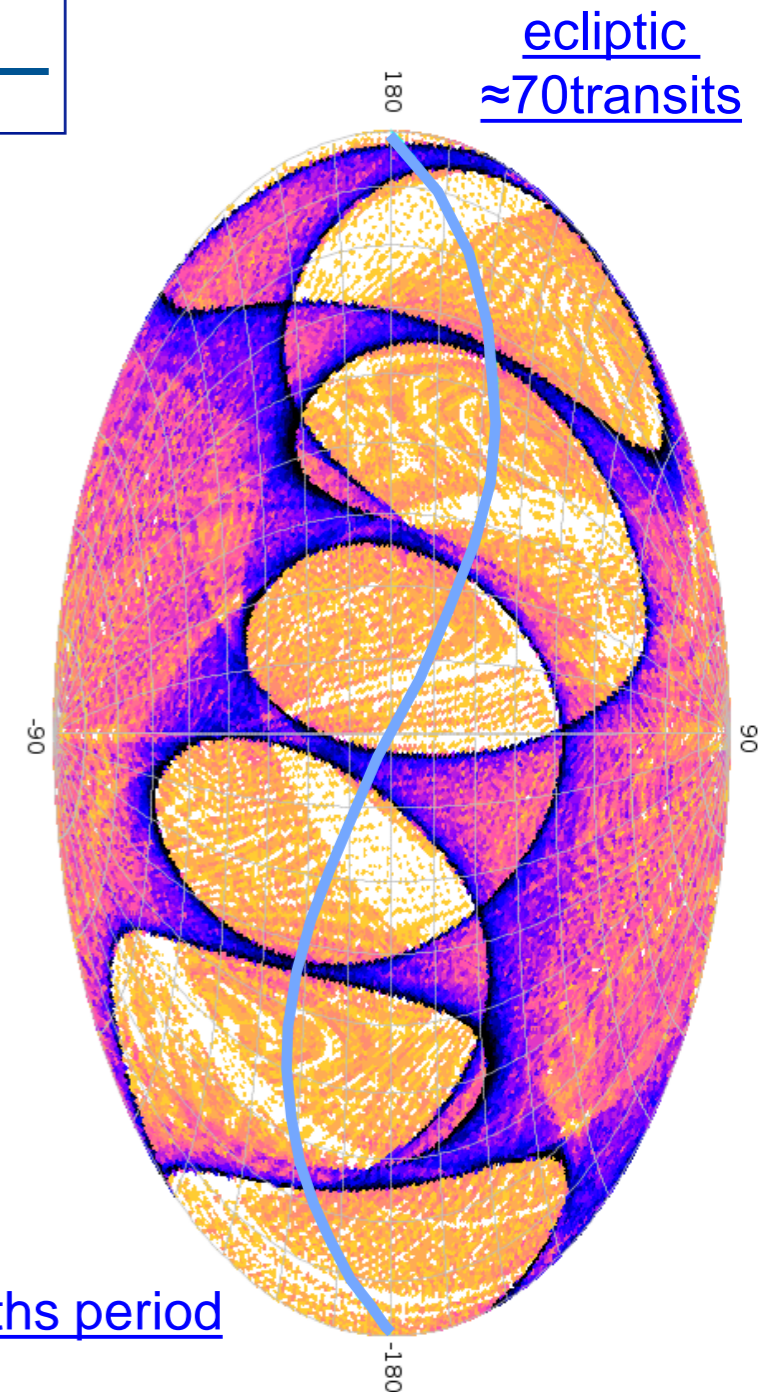
AF astr.

CCD

spectro-phot.

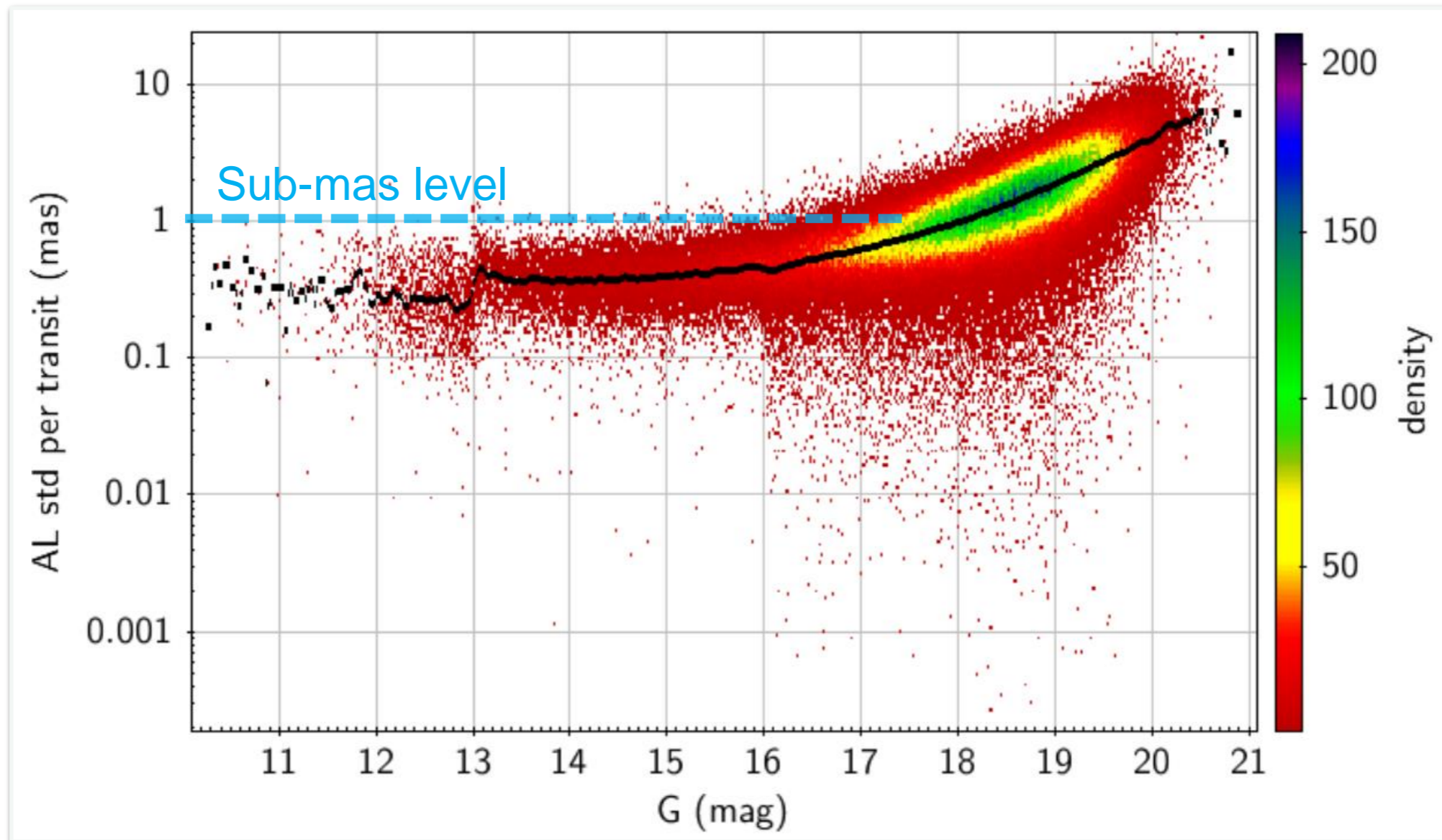
RVS

- **astrometry**
 - Ref. frames, discovering/alerts, orbits: dynamics, relativity, masses, non-gravi. forces, ...
- **‘imaging’**
 - size, shape, binarity, ...
- **G, RP/BP photometry**
 - lightcurves : spin, shape, colours-taxonomy,...
 - albedo, density
- *spectroscopy*
 - *calibration*
- not SNR limited, but telemetry limited





Gaia DR2 - SSO - transit astrometric performance

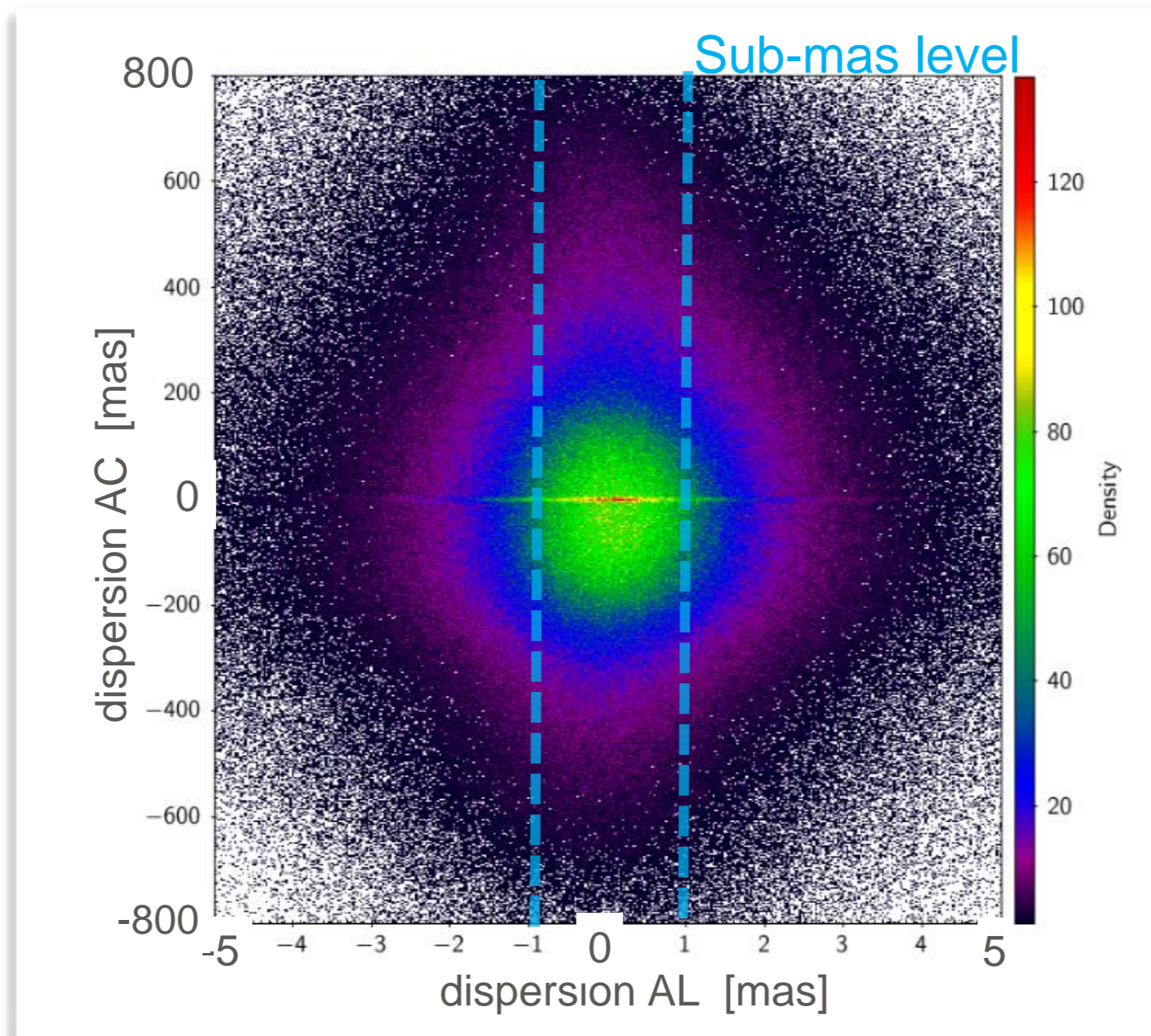


Residuals from the orbital fit of Gaia DR2 data only (along-scan direction)
(Spoto et al. 2018 A&A)

+ Improve past observations from stellar catalogue (GDR2 debiasing)



Gaia DR2 - SSO - transit astrometric performance



Residuals from the orbital fit of Gaia DR2 data only
AL + AC : note different scale!

+ Improve past observations from stellar catalogue (GDR2 debiasing)

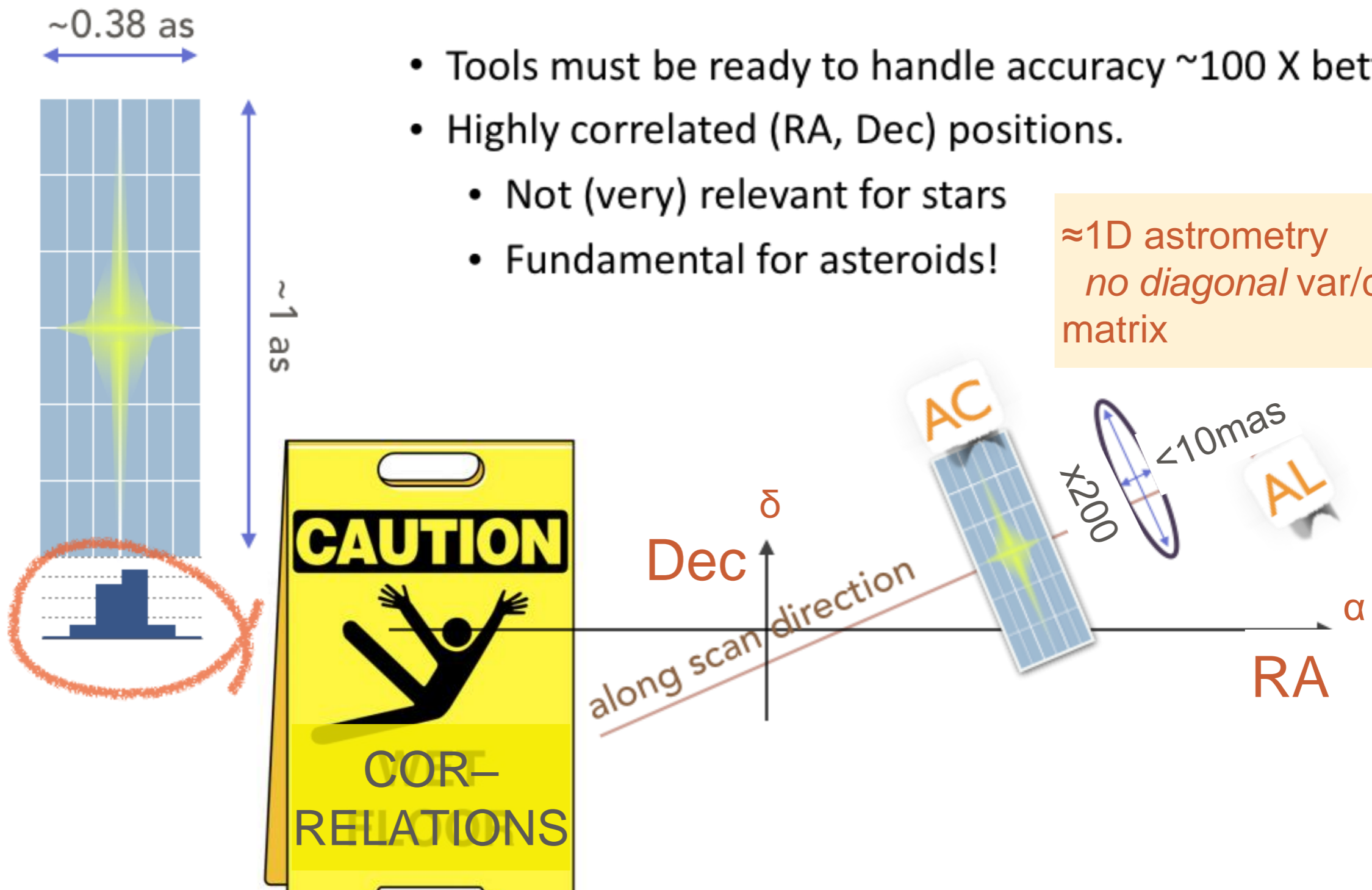


Gaia DR2 – SSO

CAUTION

- Tools must be ready to handle accuracy ~ 100 X better
- Highly correlated (RA, Dec) positions.
 - Not (very) relevant for stars
 - Fundamental for asteroids!

$\approx 1D$ astrometry
no diagonal var/covar matrix





Gaia and the asteroids

- **Before Gaia**

- 1.9×10^8 measurements in the archives of the Minor Planet Center
 - mostly CCD imaging + reduction catalogues (USNOB, UCAC, PPMXL, ...)
 - average accuracy: ≈ 400 mas
- ≈ 2000 radar ranging measurements
 - equivalent accuracy : 1-50 mas

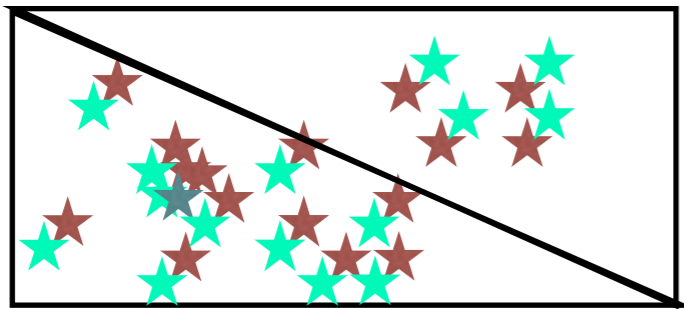
- **With Gaia Challenges from sub-mass accuracy**

- How motion / size impacts precision & accuracy?
- Can « 1D » measurements provide the expected orbit accuracy?
- Is our dynamical Solar System model accurate enough?
- How to combine Gaia and « pre-Gaia » astrometry?

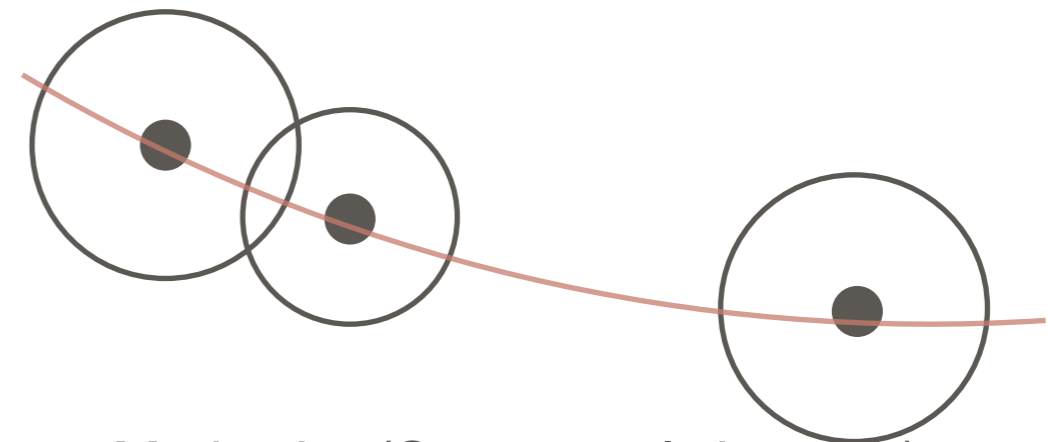


Orbit improvement : DR2 + other observations

- Combining DR2+past observations: a factor ≈ 2 (**only**) average improvement (DR2=snapshot)
- But: GB data affected by systematics (zonal/calibration) errors of the pre-Gaia catalogs, and large difference in astrometric precision
- **Debiasing required for past observations!**
 - if possible = re-reduction
 - simplest way = ad-hoc method



Method 1 : Chesley, Farnocchia
Eggl et al. *Icarus accepted*
all catalogs referred to a reference
• corrections computed on a grid
(F2015 scheme)



Method 2 (Spoto et al. *in prep.*) :
areas centered on archive asteroid observations
Reference: Gaia DR2
(DR2 scheme)

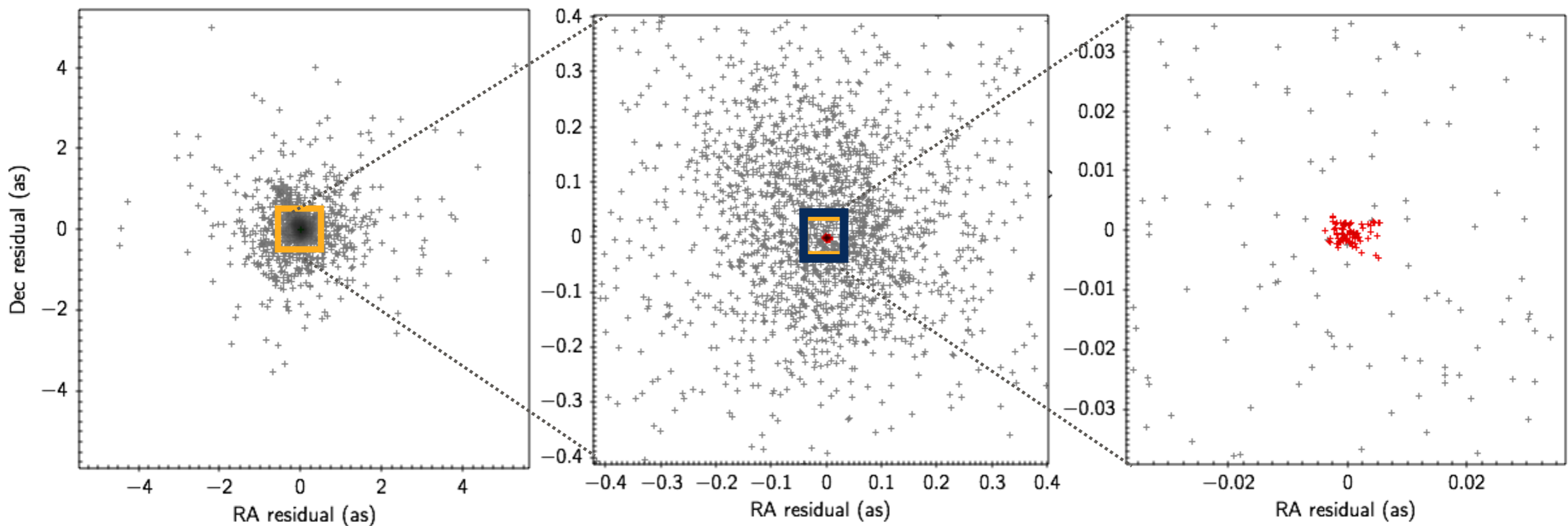
- **In all cases, precision and accuracy of past observations increased**



Asteroid (386) Siegena - residuals from orbital fit

Combining archive data (2776 obs.) to Gaia data

With GDR2 debiasing

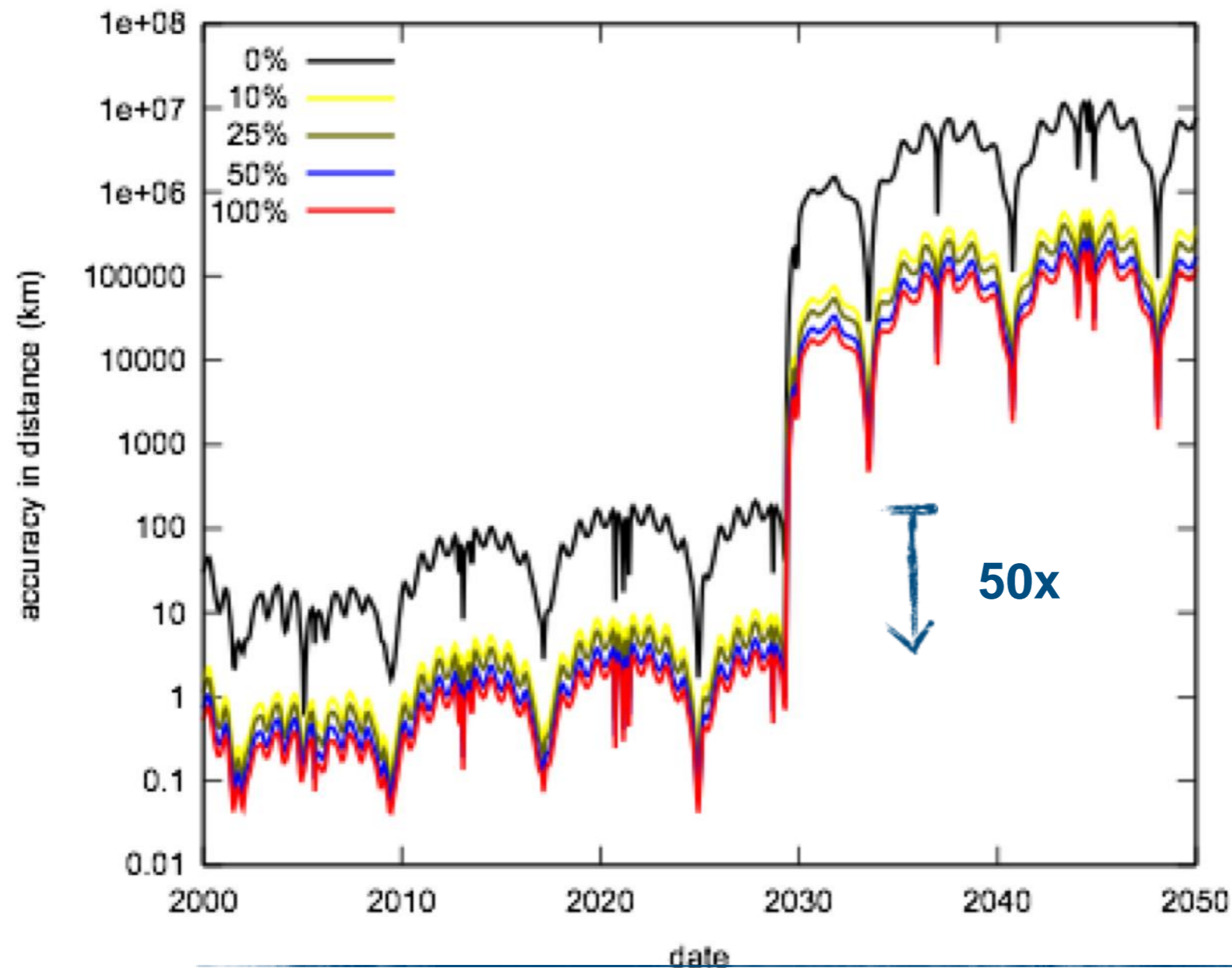


factor 100 X improvement!



Orbit propagation

- re-reduction of <50% data



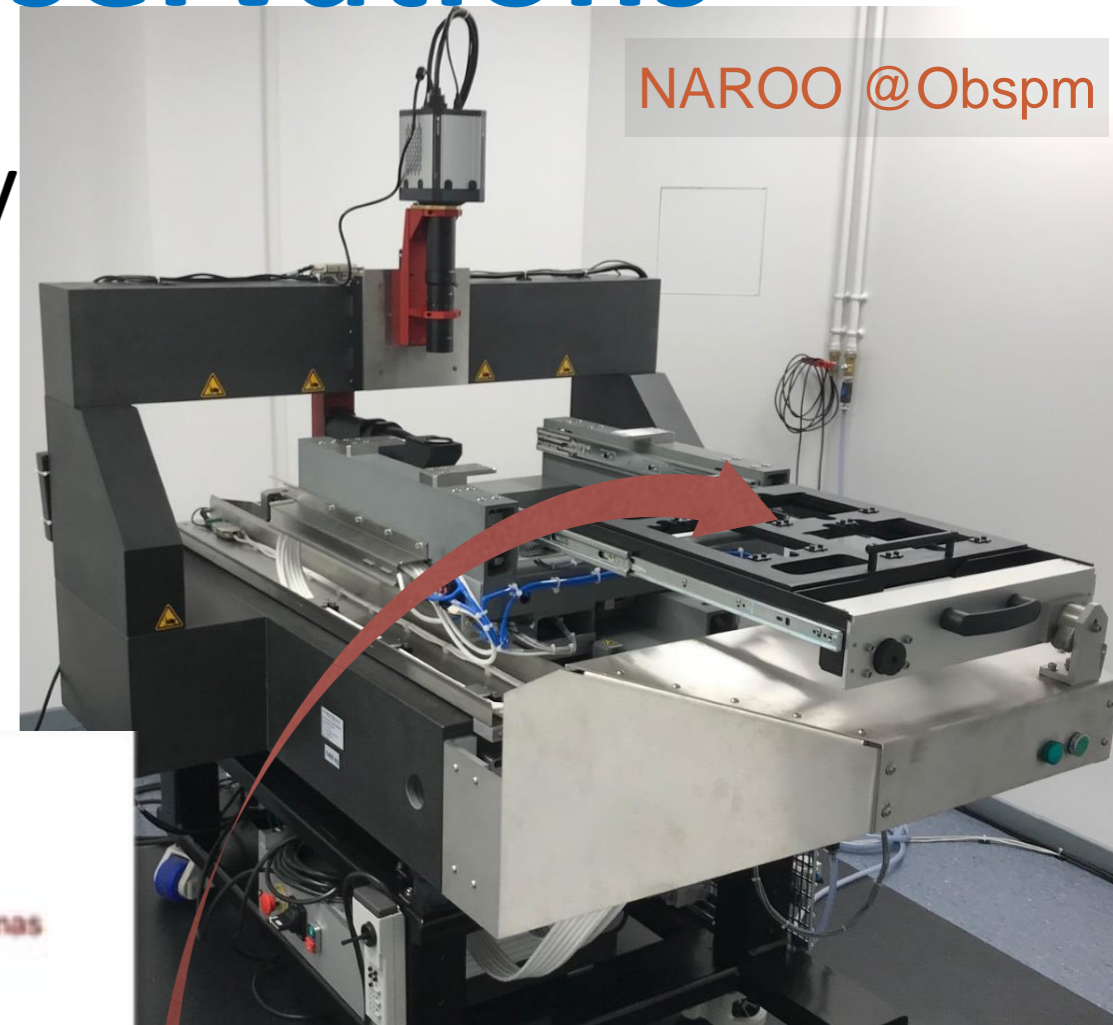
(Desmars et al. 2013)

Apophis: position uncertainty reduced by factor 50
with only 10% of MPC data re-reduced

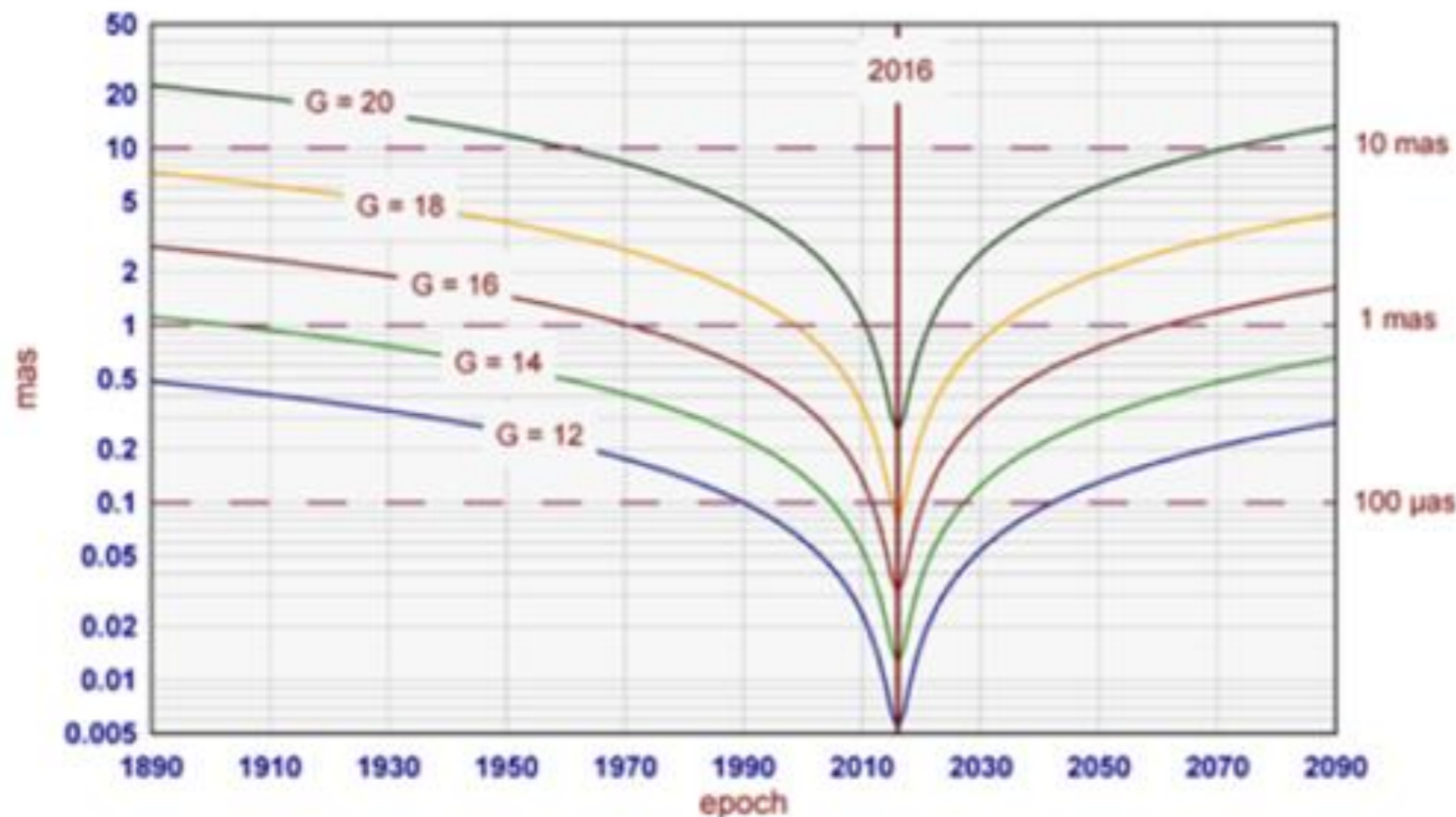


Revisiting past-observations

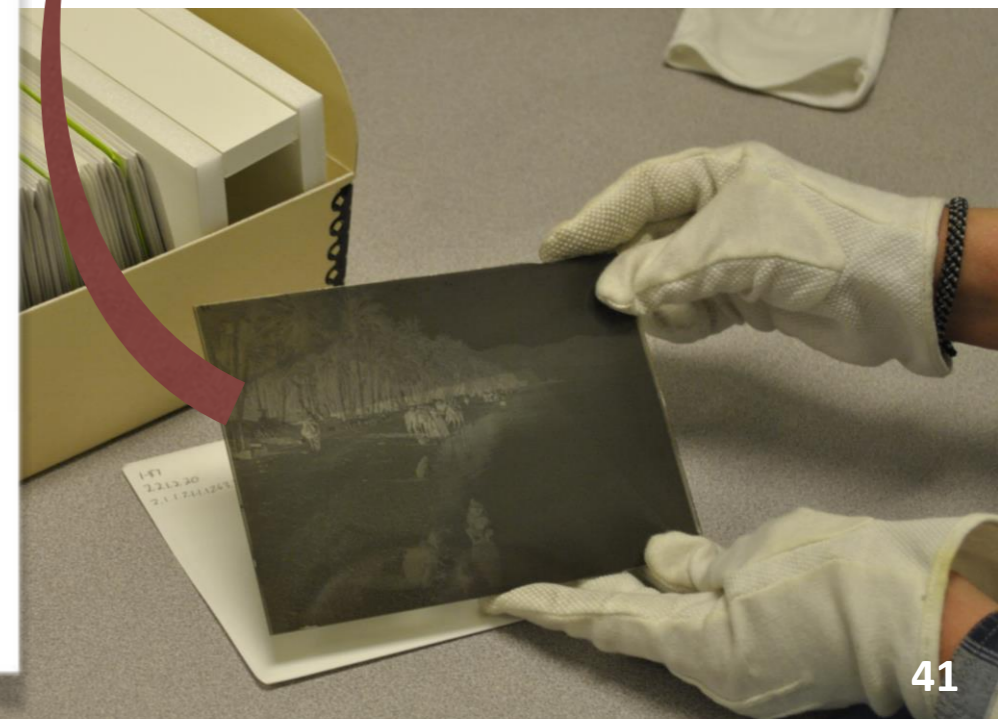
- Ad-hoc Debiasing MPC astrometry
- Recalibration
GBOT, OSSOS, Pan-Starrs, etc.
- Photographic plates: NAROO



NAROO @Obspm



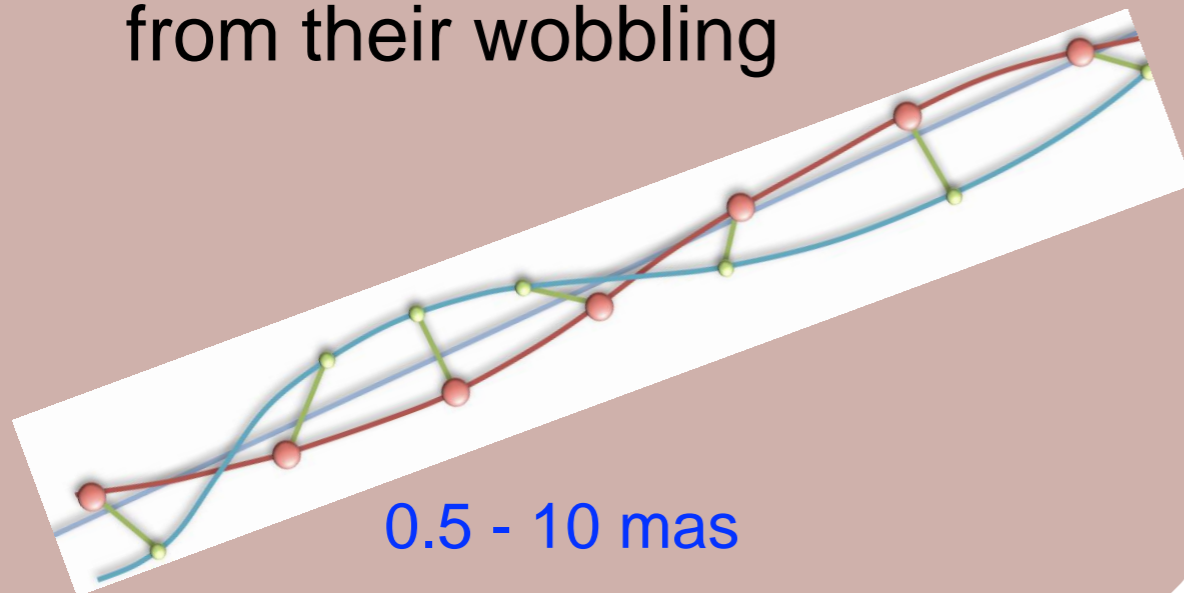
(© F. Mignard)



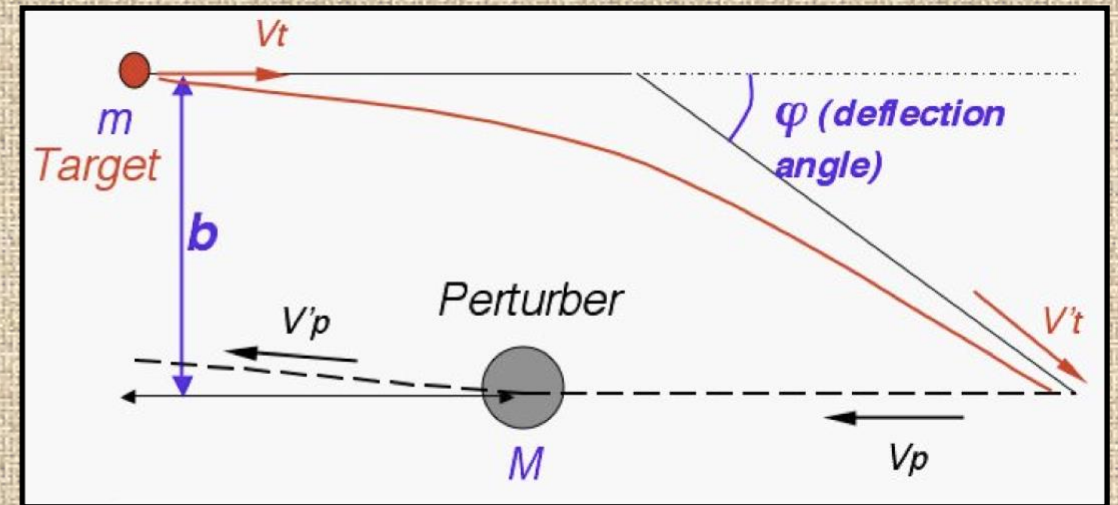


Some challenges for asteroid astrometry

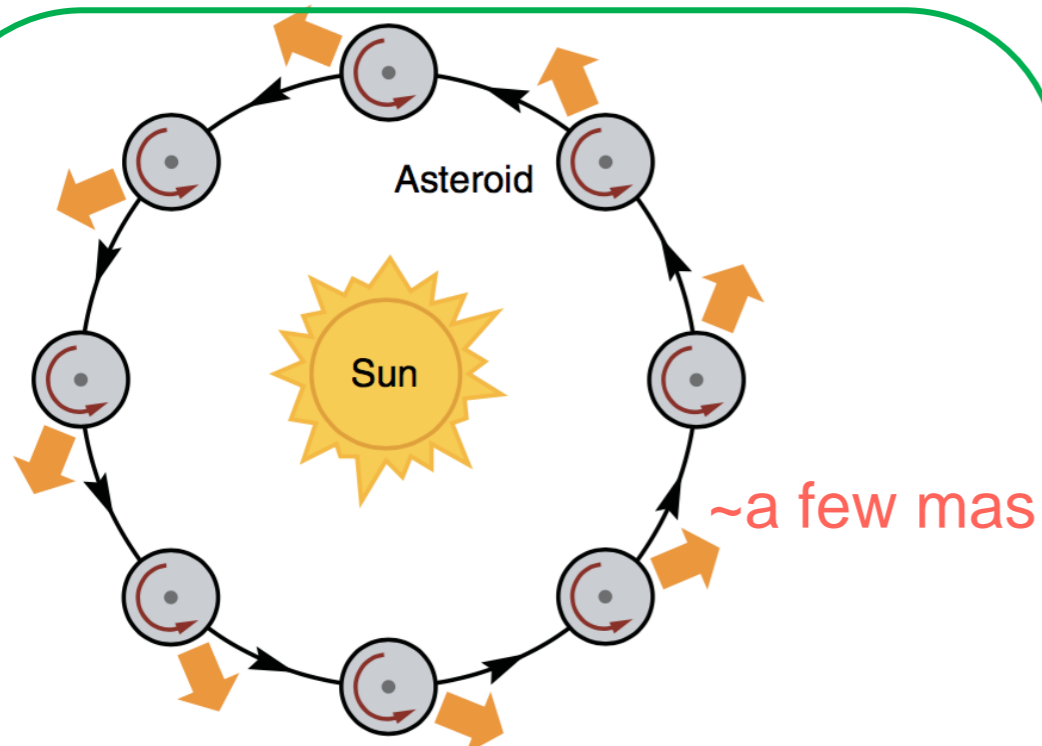
Discover asteroid satellites
from their wobbling



New / precise asteroid masses

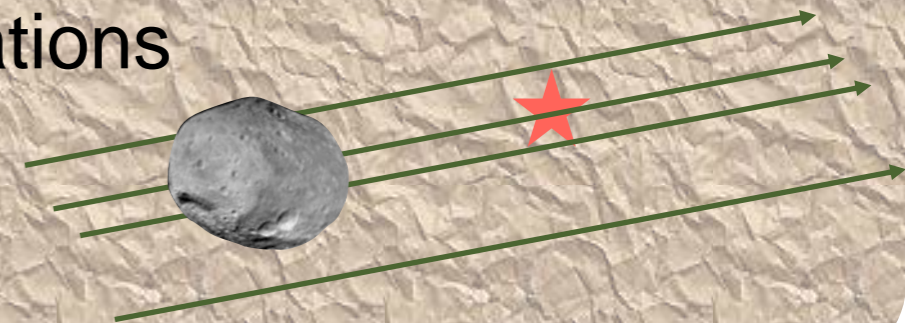


1000s encounters/year > 10 mas



Measure the orbital drift
due to Yarkovsky force

Improve predictions of stellar
occultations



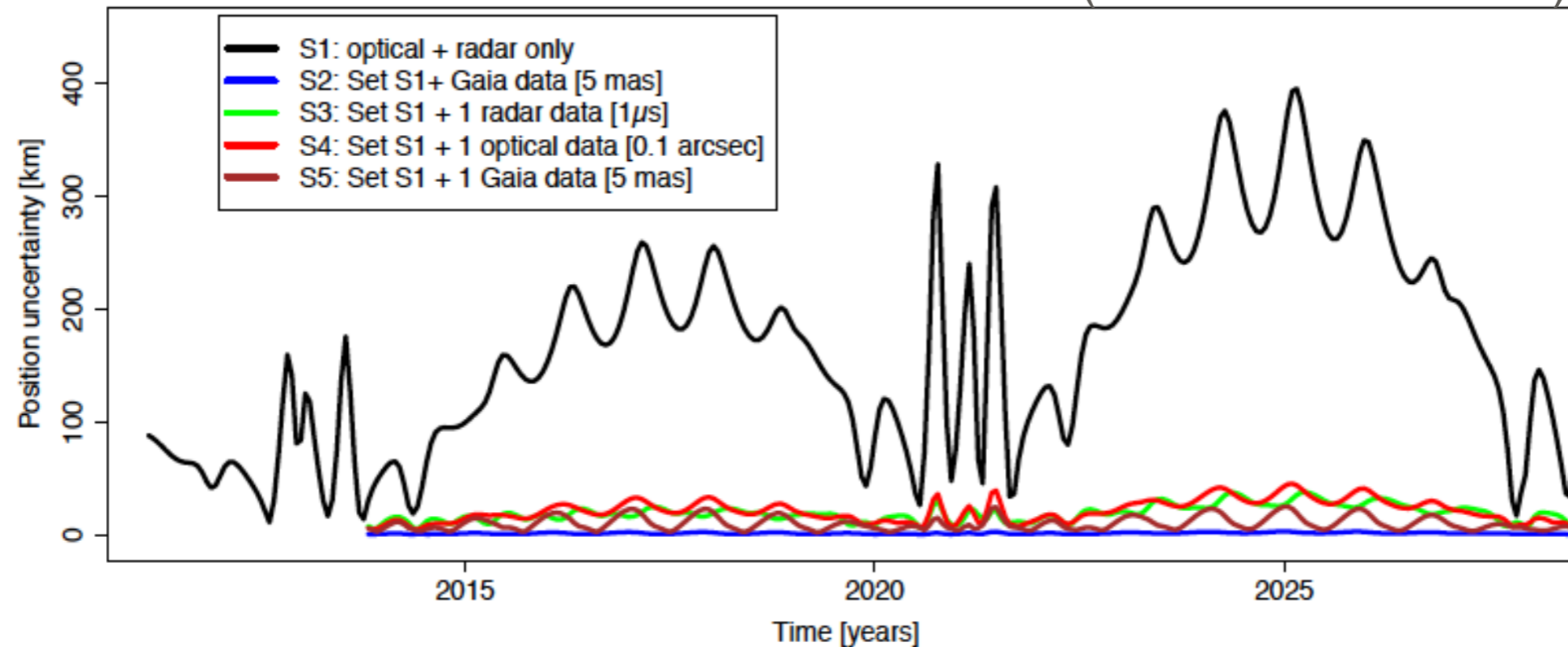
+ spectro-photometry, spin, taxonomy, sizes, ...
+ reference frames, GR testing, ...



NEO orbit propagation

- Better – longer prediction of orbits, and IP

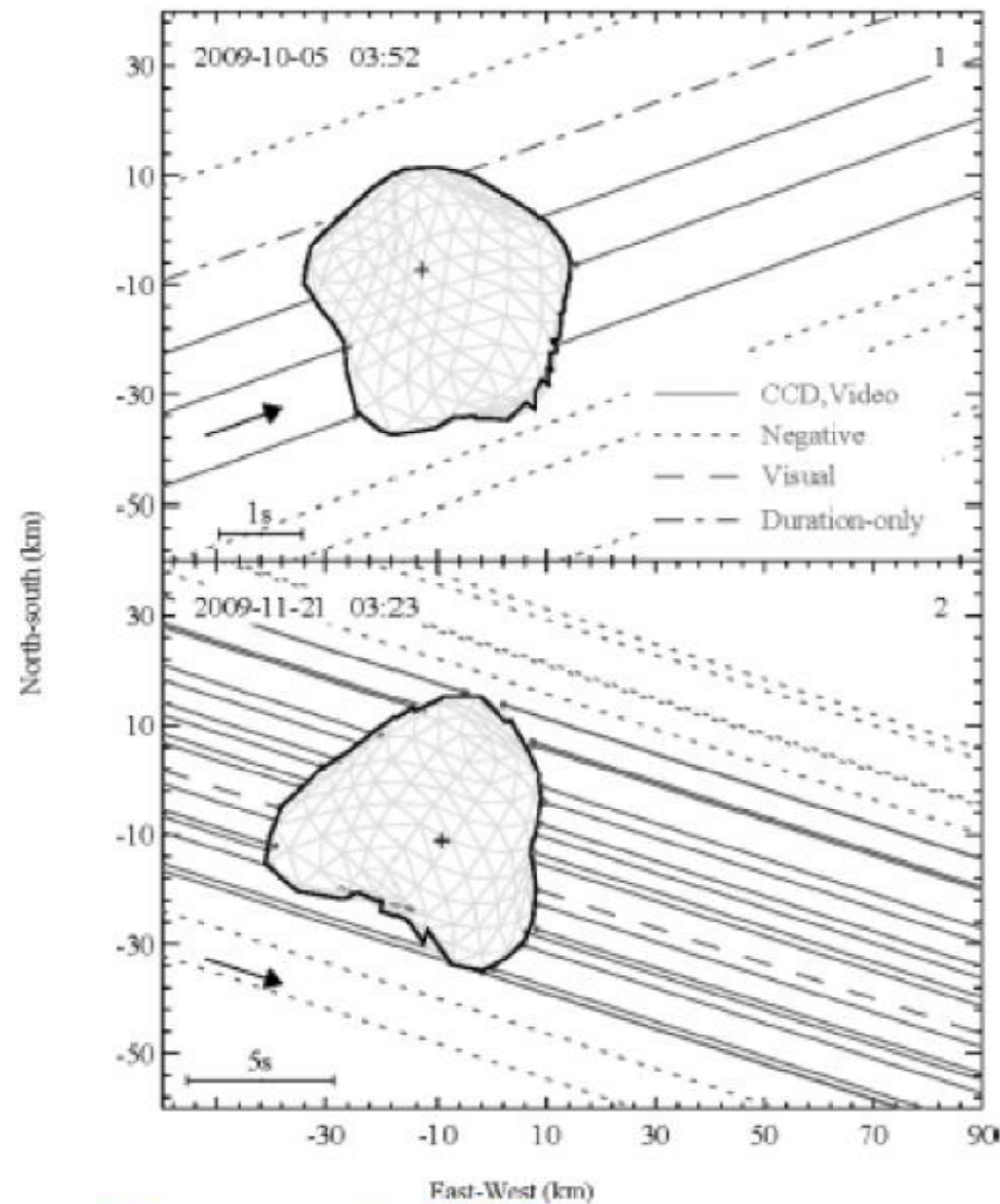
(Bancelin et al. 2013)



Apophis: position uncertainty reduced with Gaia at km level
similar to best radar observations



Stellar occultations



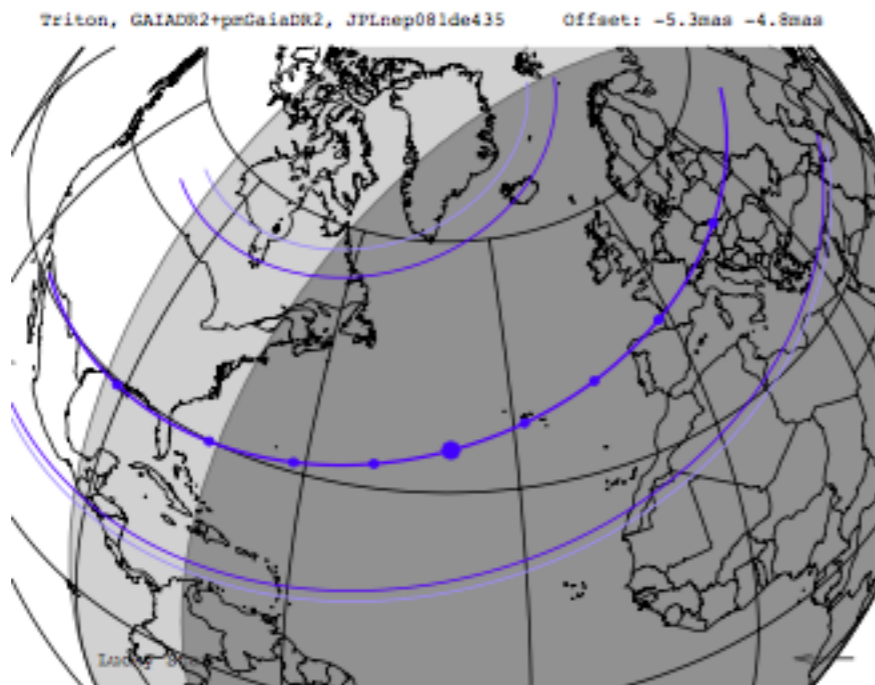
Tanga et al. 2015

Success!
better prediction => more chords
more chords => better science
Shape detailed and
direct size of small bodies

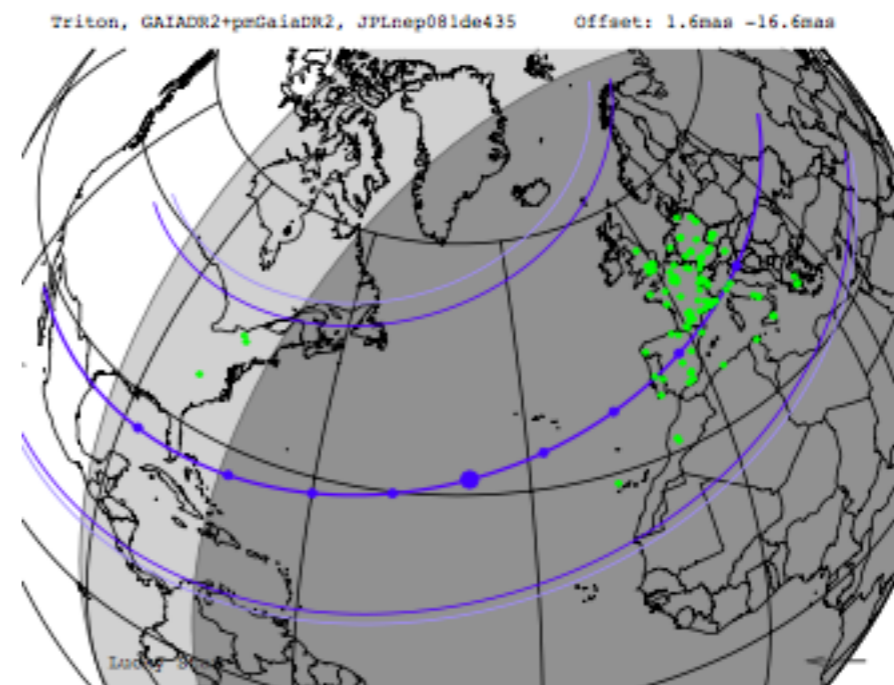
Stellar occultations with Gaia DR2

Triton occultation – 5 Oct. 2017

DR1 in Sep.



DR2 in Oct.



(@B. Sicardi et al.)

use of 119 preliminary DR2 stars

better positions of both Triton & occulted star

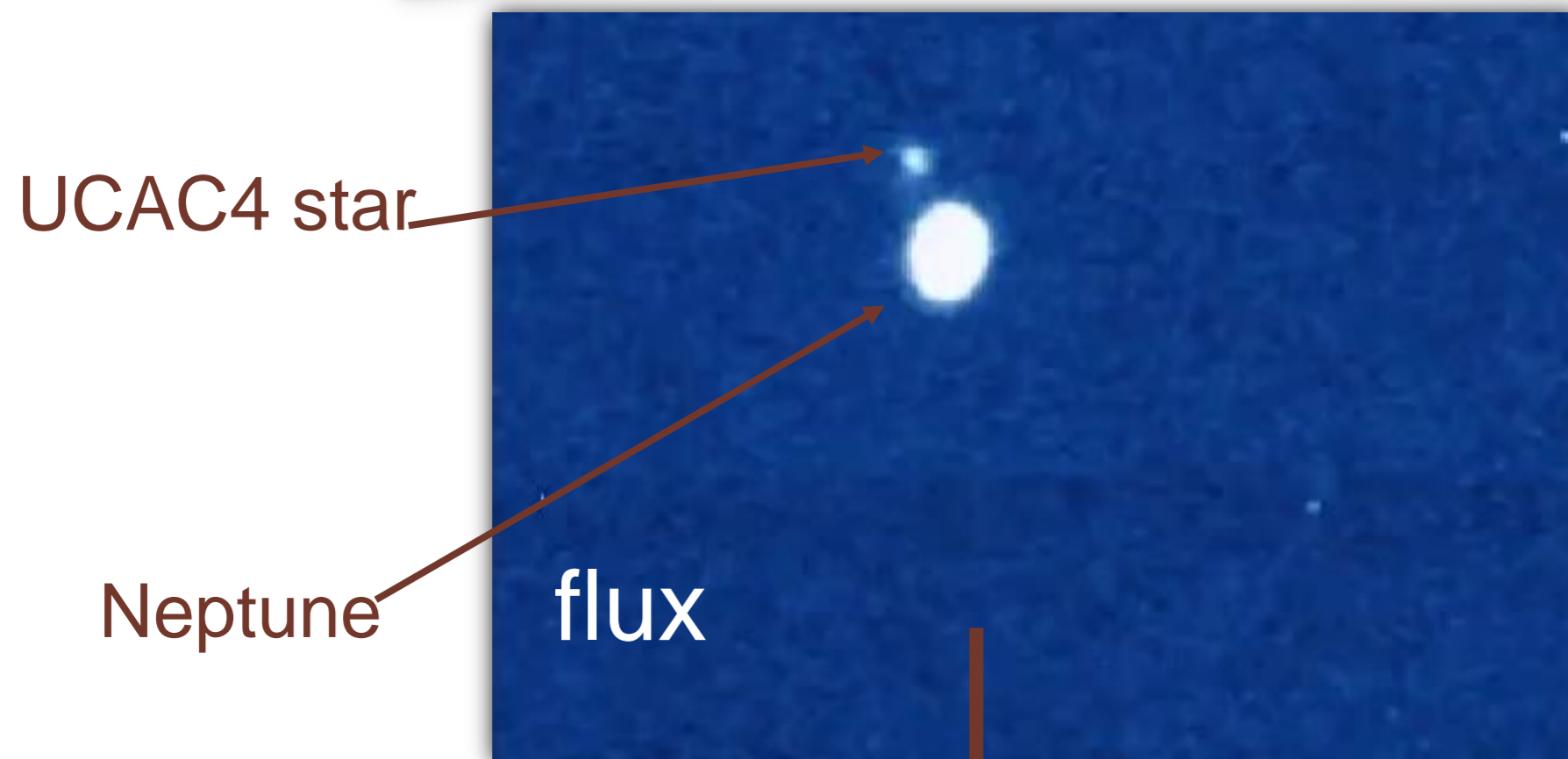
https://www.cosmos.esa.int/web/gaia/news_20170930

<http://sci.esa.int/gaia/60015-central-flash-during-triton-occultation/>

Stellar occultations with Gaia DR2

Triton occultation – 5 Oct. 2017

precision on ground
 $\approx 7\text{s}$ in time and $\approx 40\text{km}$ on path



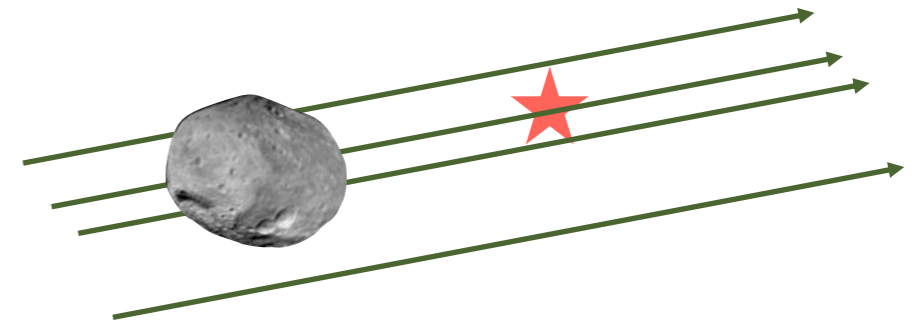
(© B. Sicardi et al.)

Central flash!! from satellite atmosphere



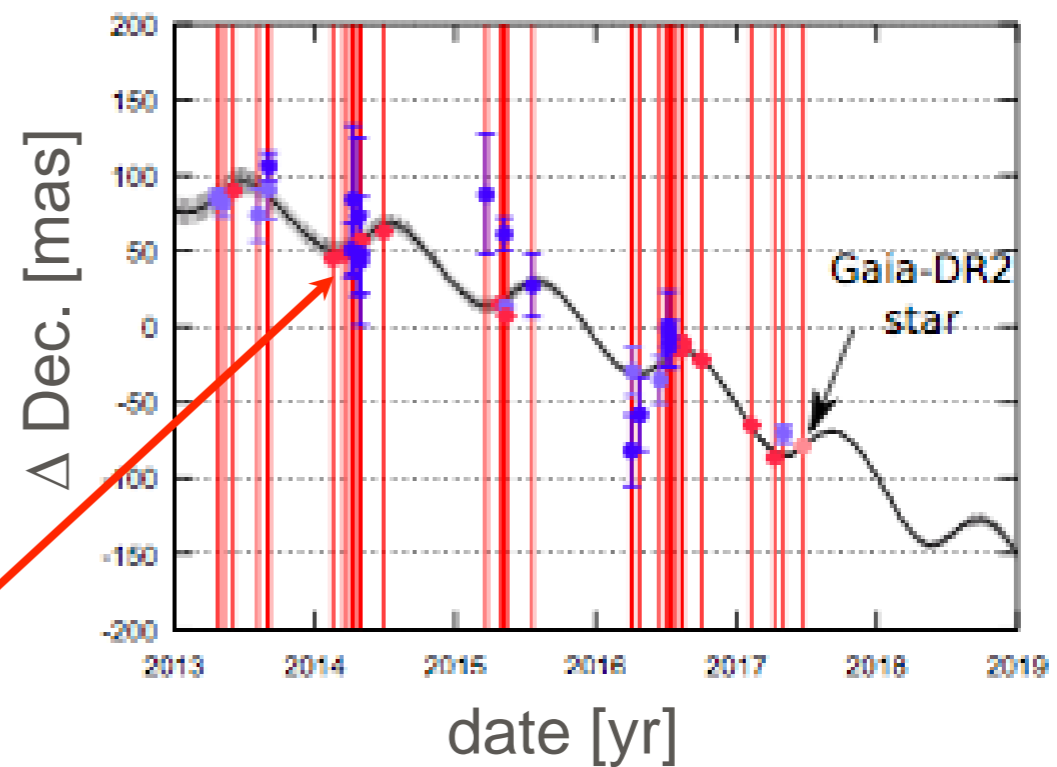
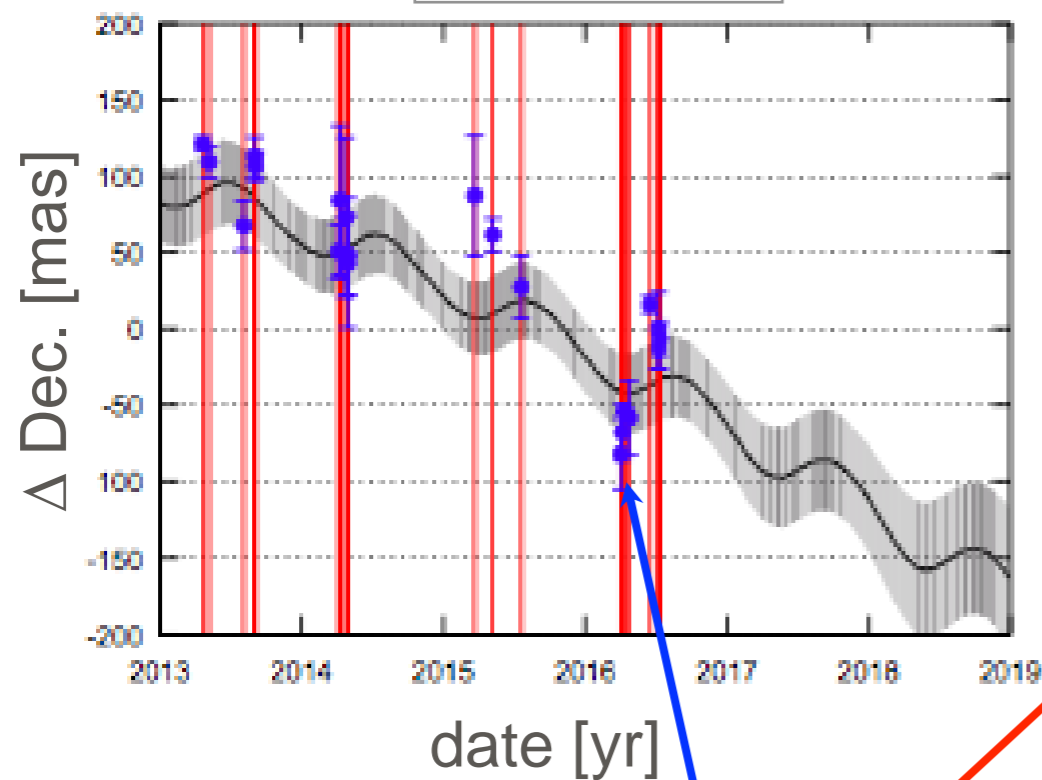
The occultation astrometry with Gaia

- DR2 successfully exploited
 - Triton (October 2017)
 - 2014 MU69 ($V=27.5$, 50 km KBO, ~ 1.4 mas)!!



UCAC 4

Gaia DR2



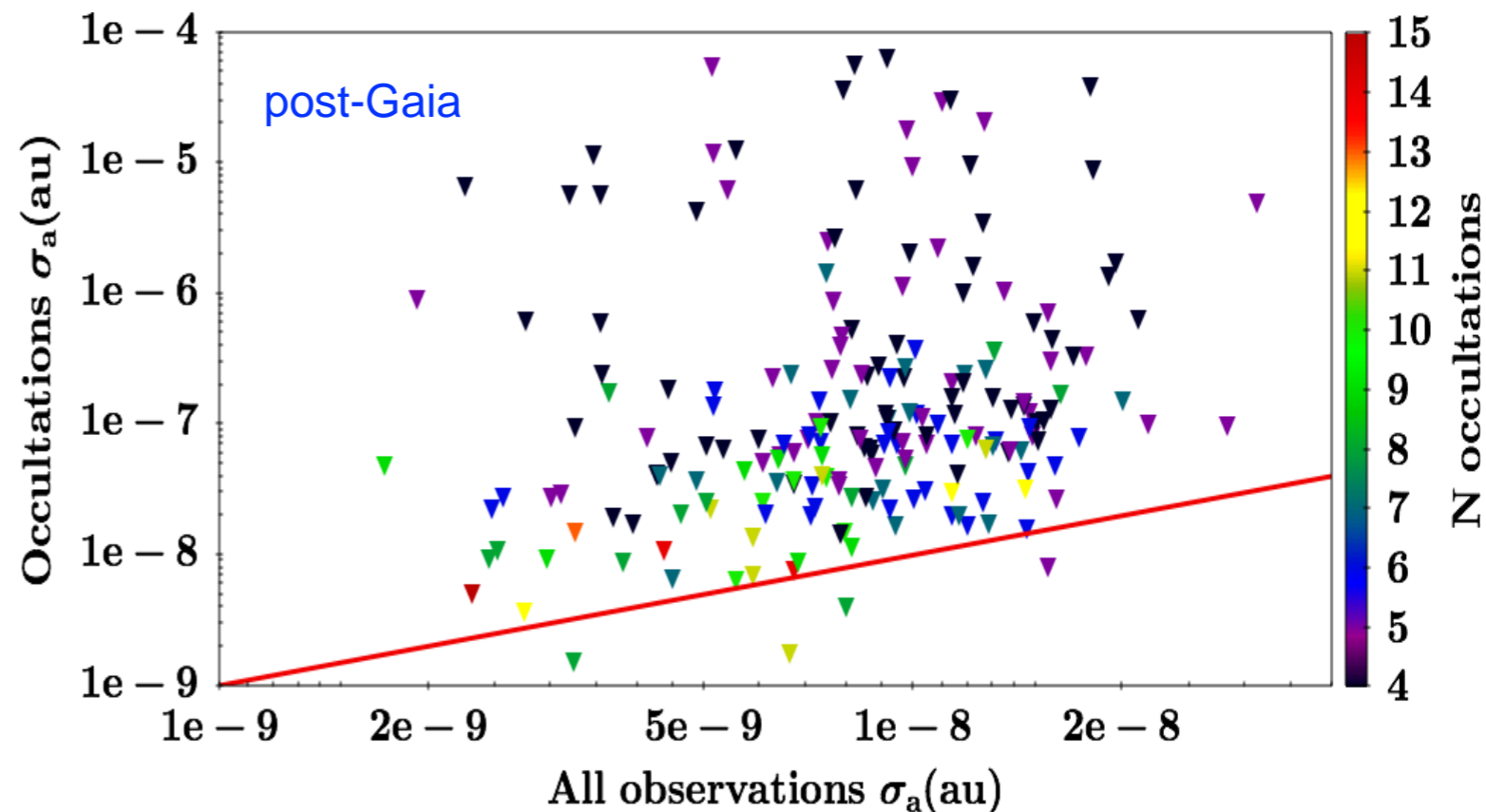
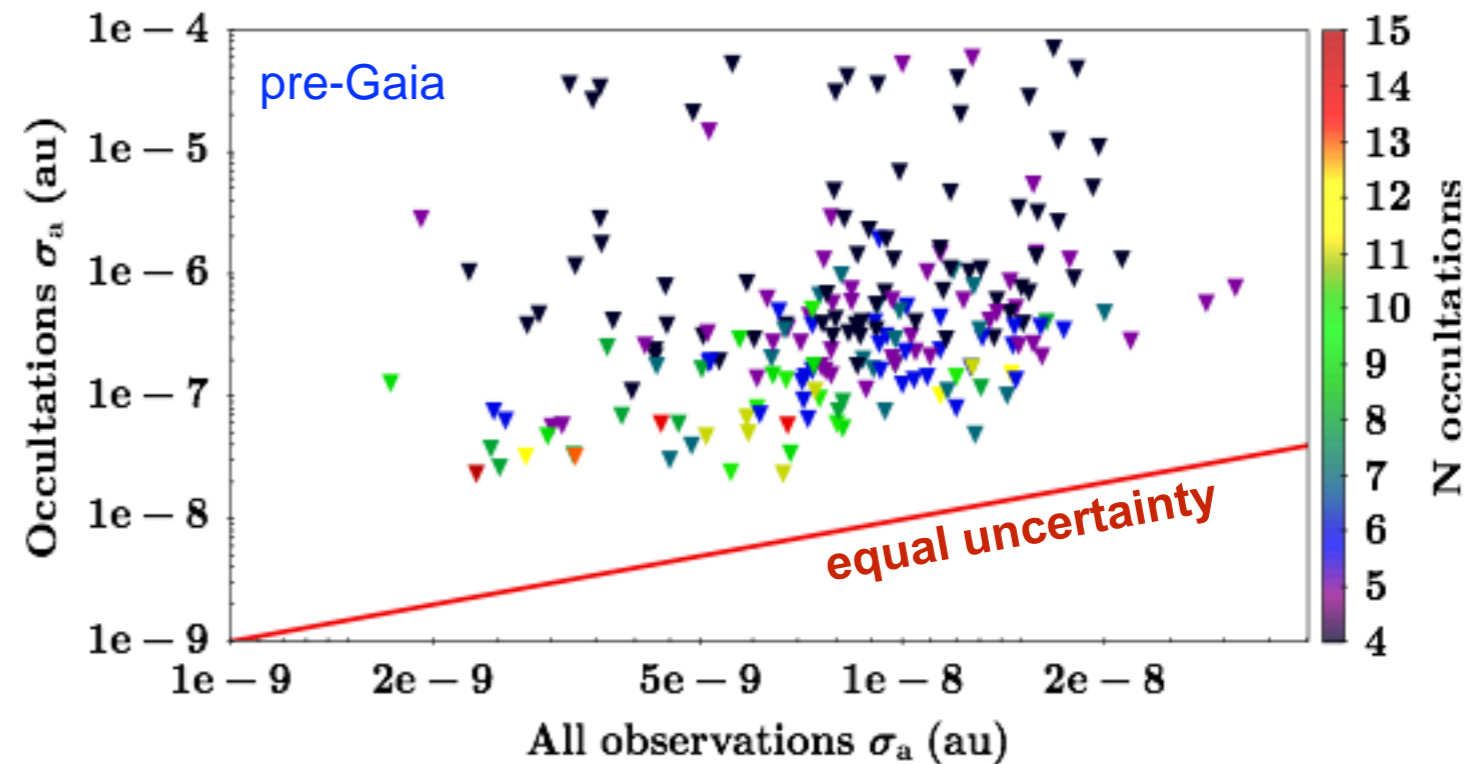
red: SSO = DR star position

blue: classic astrometry

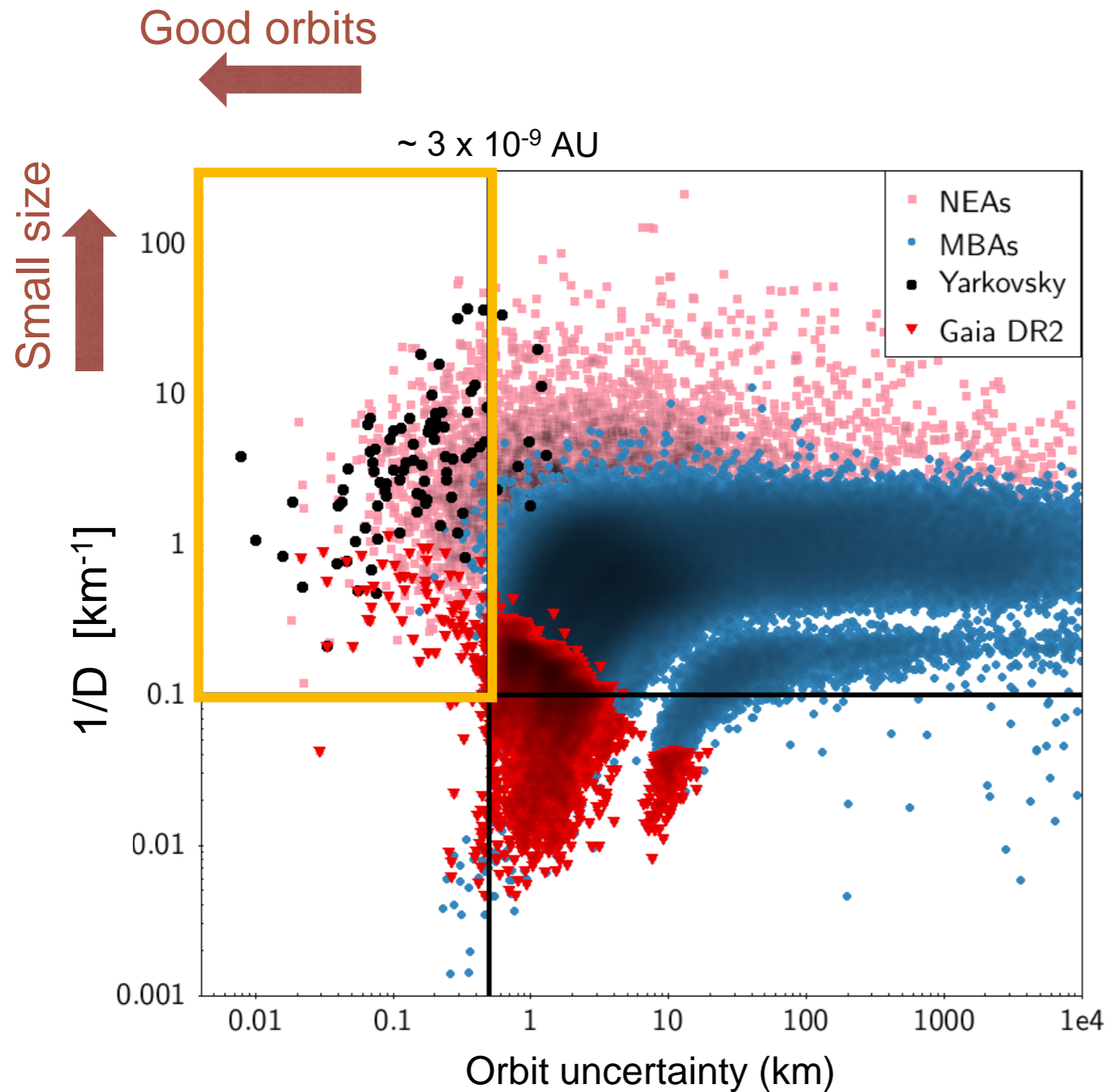


The occultation astrometry with Gaia

- Occultations: very accurate asteroid position at the level of the star astrometry (mas - μ as)
- Comparison orbit precision
 - Few (past) occultations/object vs.
 - all (1000s) MPC data
- ...beyond the duration of Gaia!
- Future of astrometry?
- $\approx 10,000$ objects down to 20km



Yarkovsky effect



Pre-Gaia orbit quality (MPC data)

In DR2:

- 3 NEOs with measured Yarkovsky
- 5 NEOs with marginal detection
- about 20 good MB candidates

(@F. Spoto)

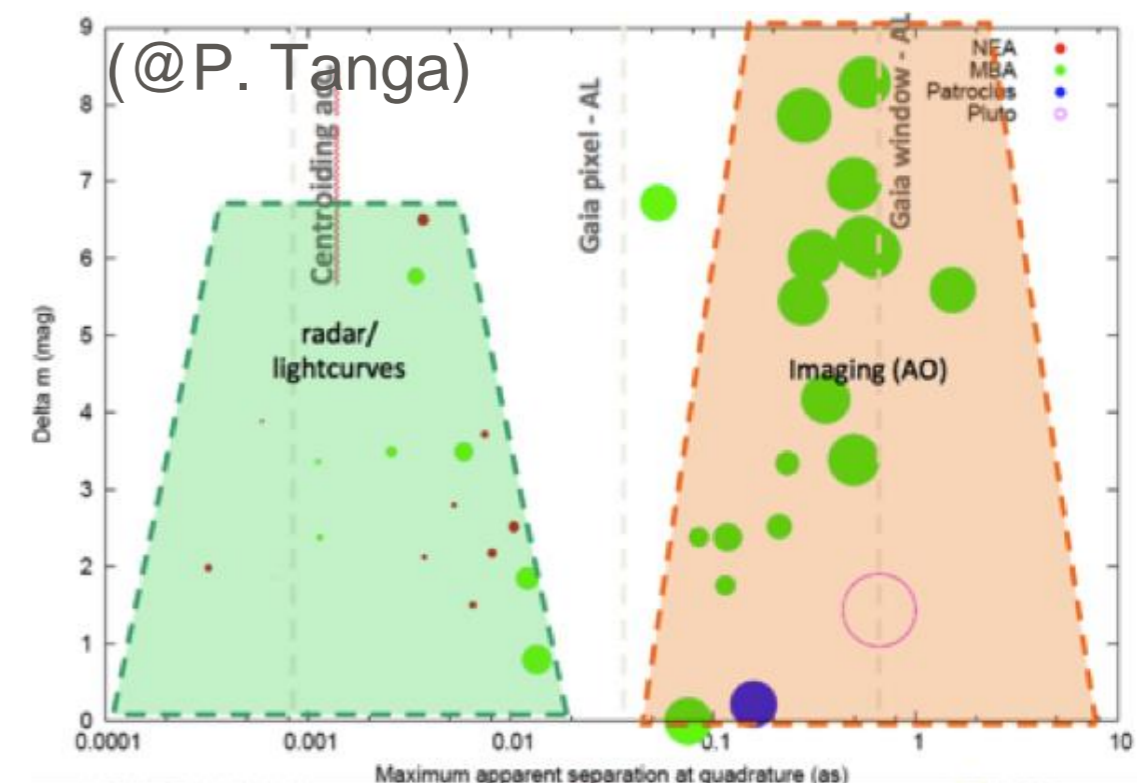
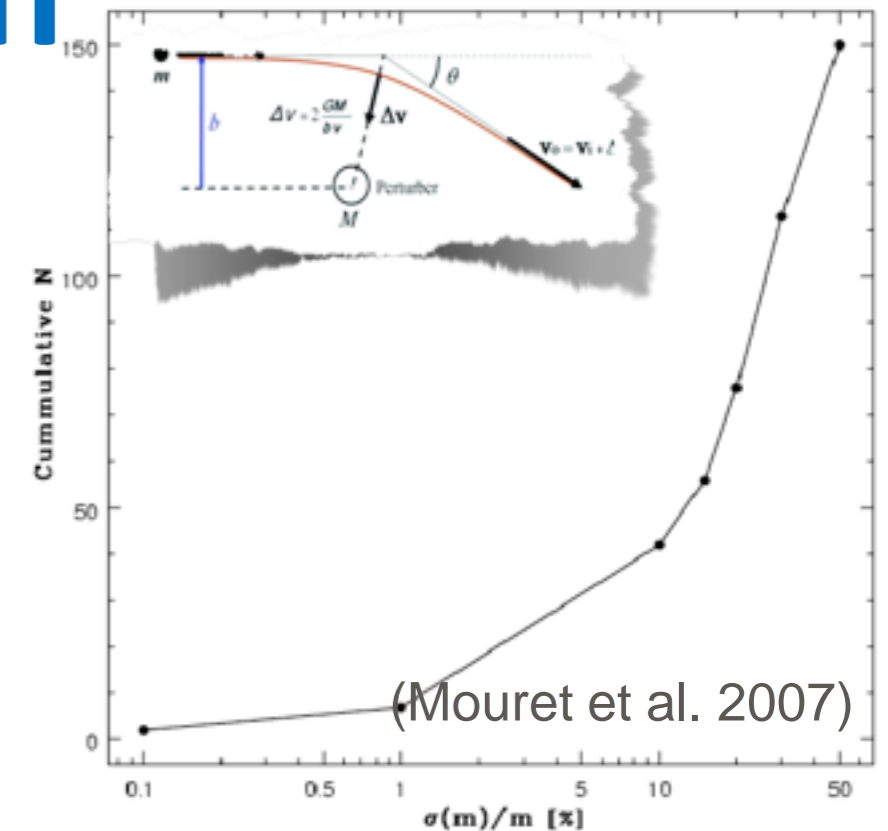


Mass determination

- Close encounters
 - ≈ 100 mass determination
 - from Gaia data alone

(@Ivantsov et al. 2017 IAU 330)
- Binary asteroids
 - resolved binaries
 - astrometric binaries
 - MCMC inv. Code

(@Kovalenko et al. 2016)
- With shape and size/albedo/taxo
 - bulk density
 - macro-porosity

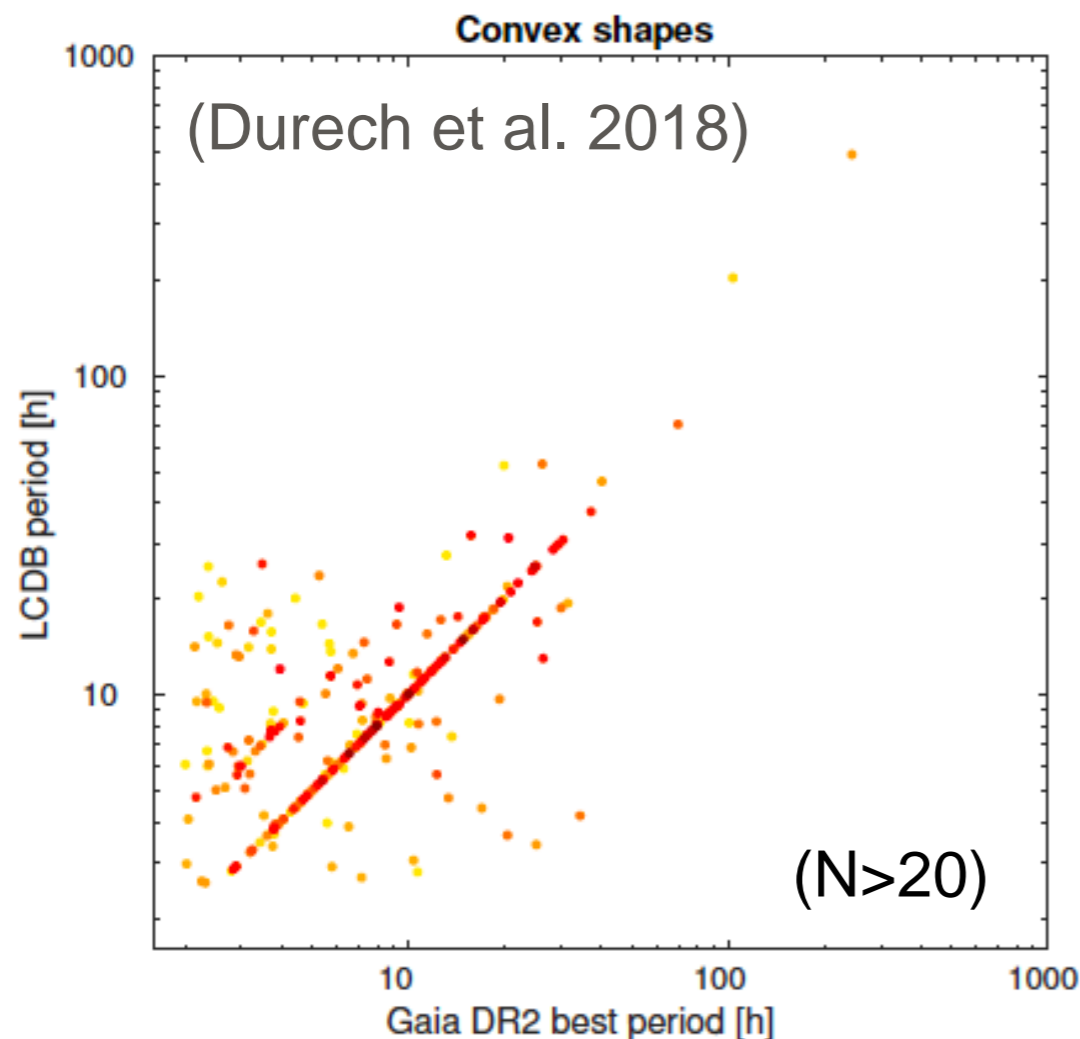




Photometry – lightcurve

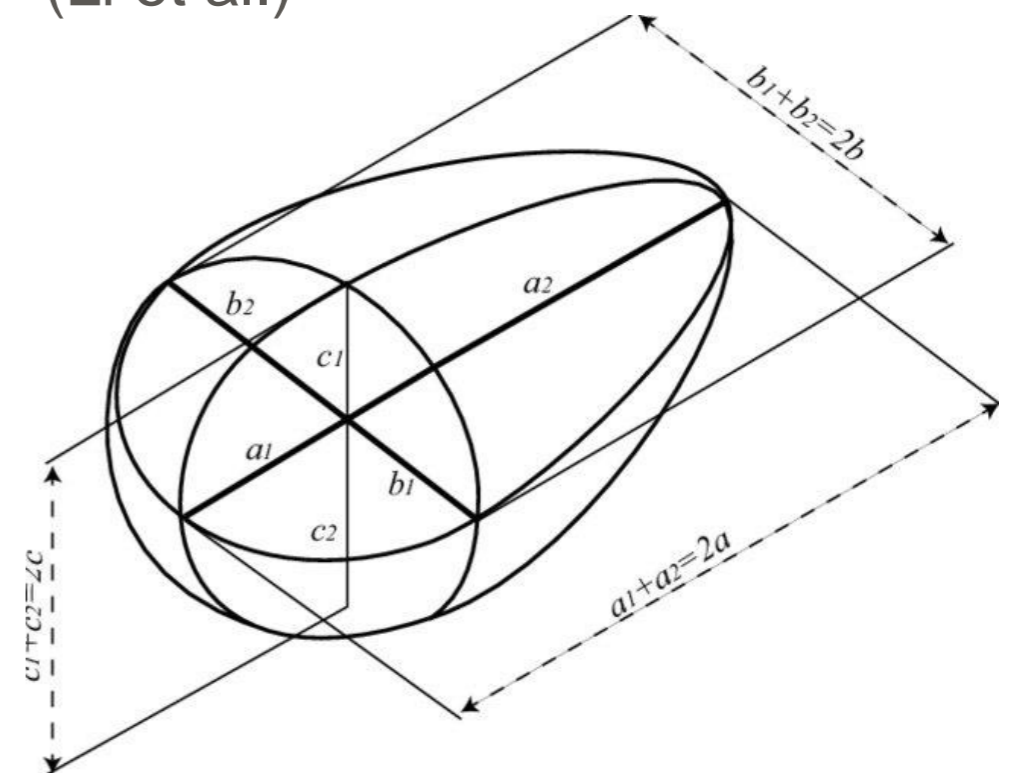
- sparse data – no full LC, yet over several years at milli-mag
- derive spin (period, direction), shapes
- valuable with $N \geq 30$ transits
- Also phase relation

topographic



Cellinoids

(Li et al.)

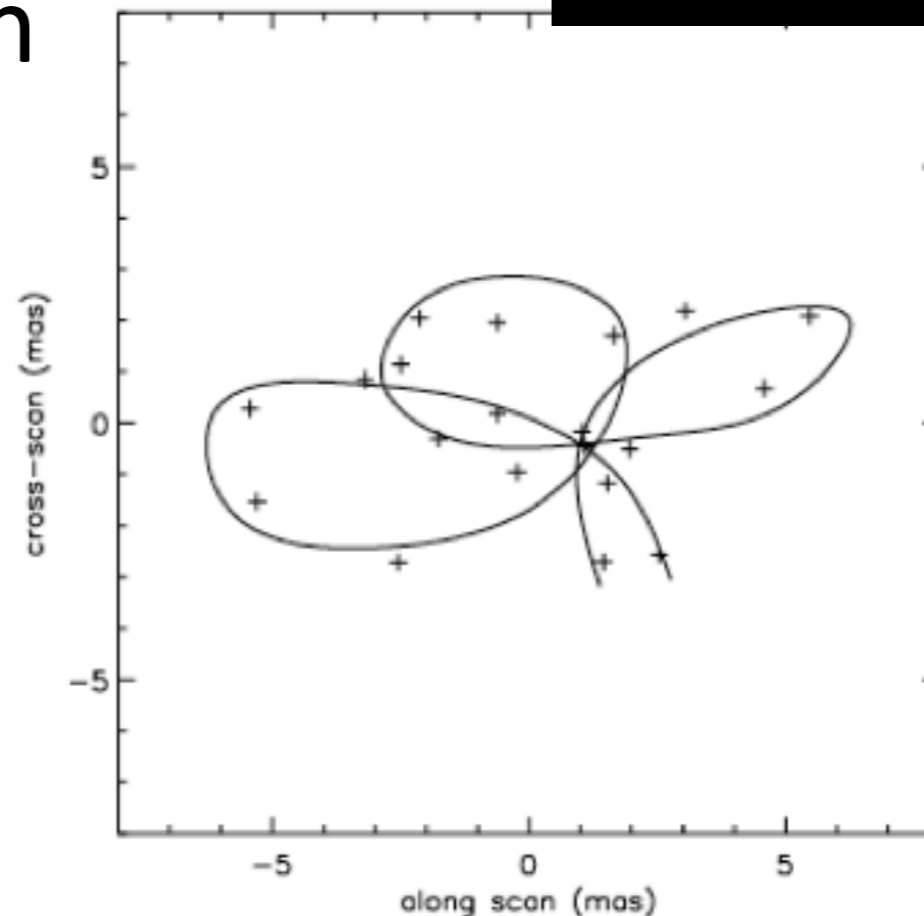
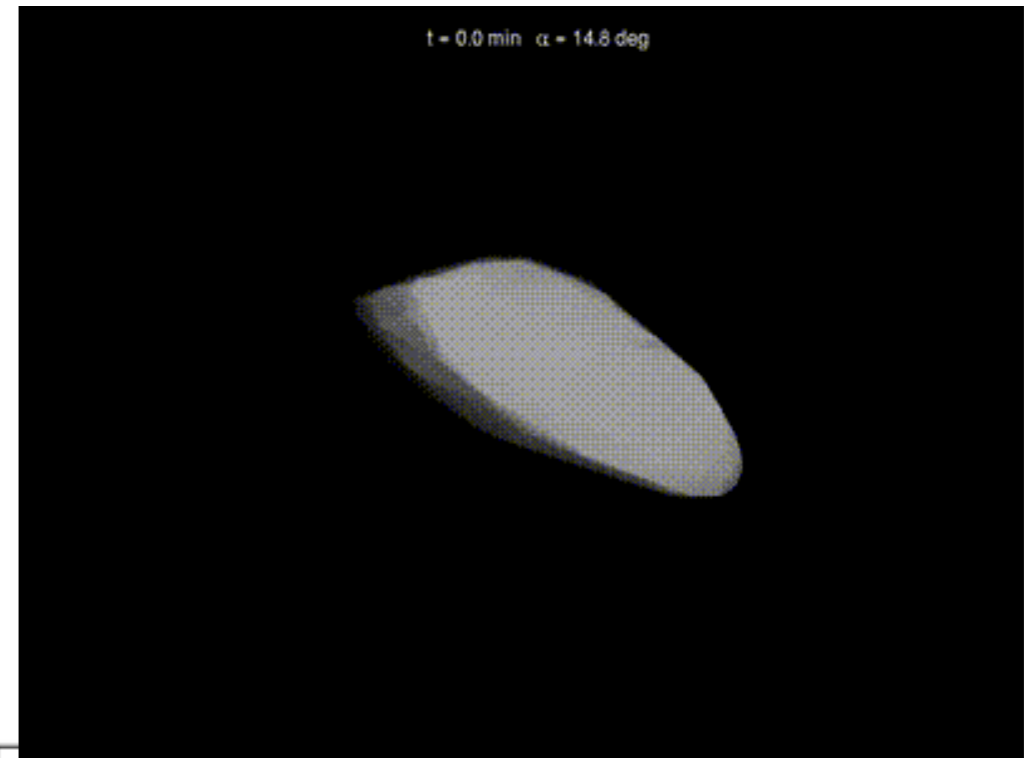


(combine Hipparcos+Gaia)



Photocentre offset

- Due to phase effect during rotation
 - + propto size
 - + shape
 - + light-diffusion

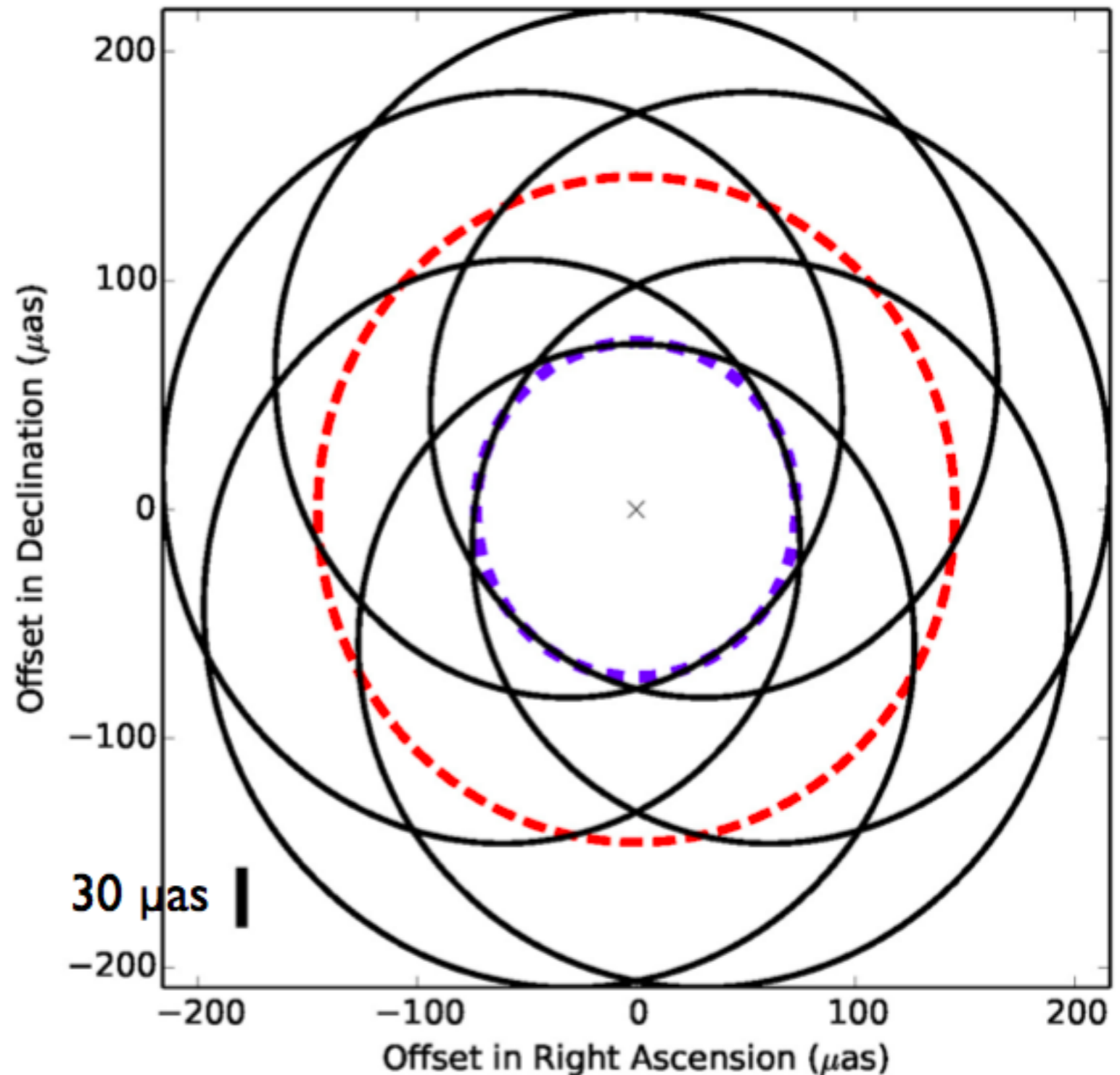


(Kaasalainen et al. 2004)



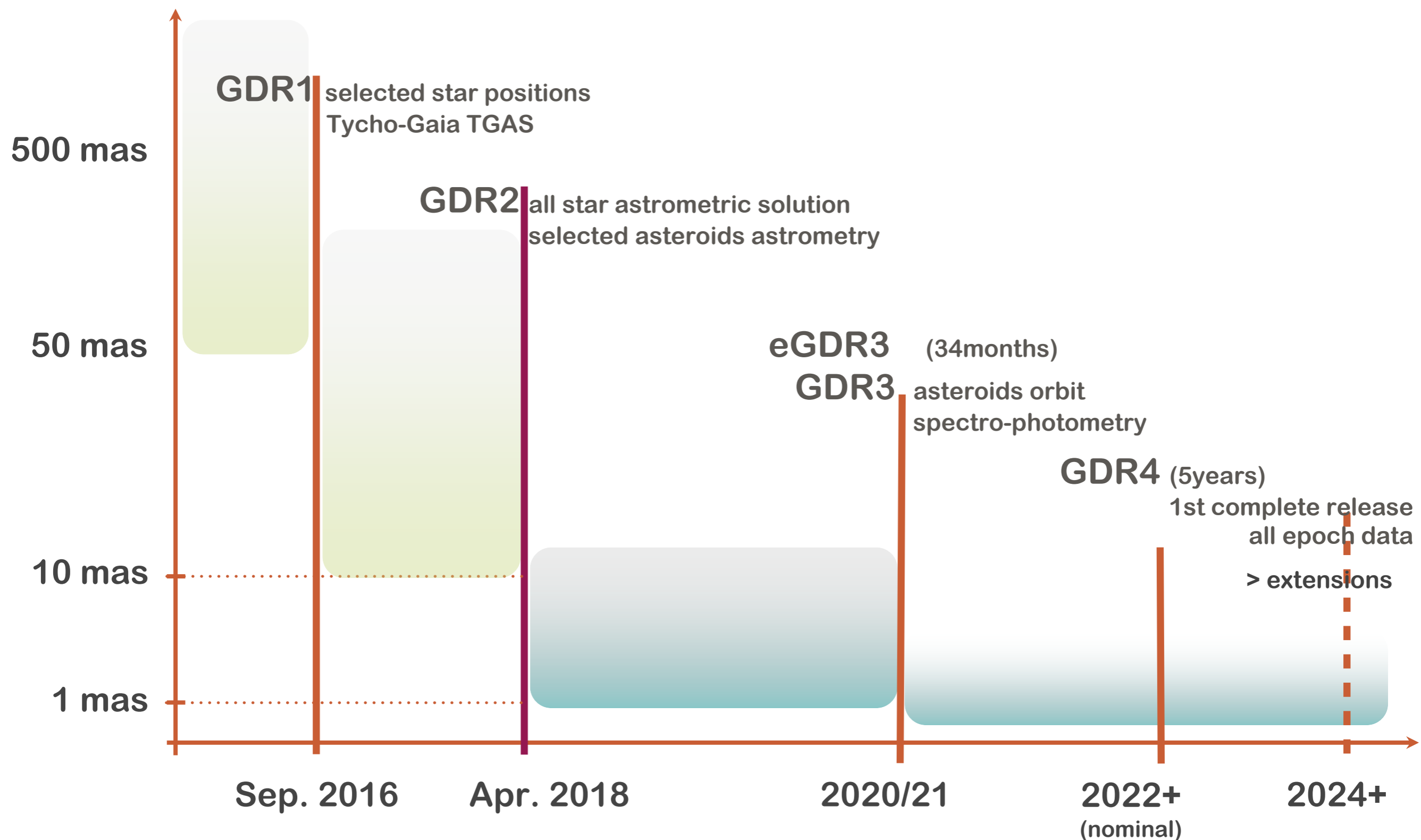
Exoplanets

- ≈ 1000 's of exoplanets expected
- circumbinary planets
- 10% of tight binaries host one





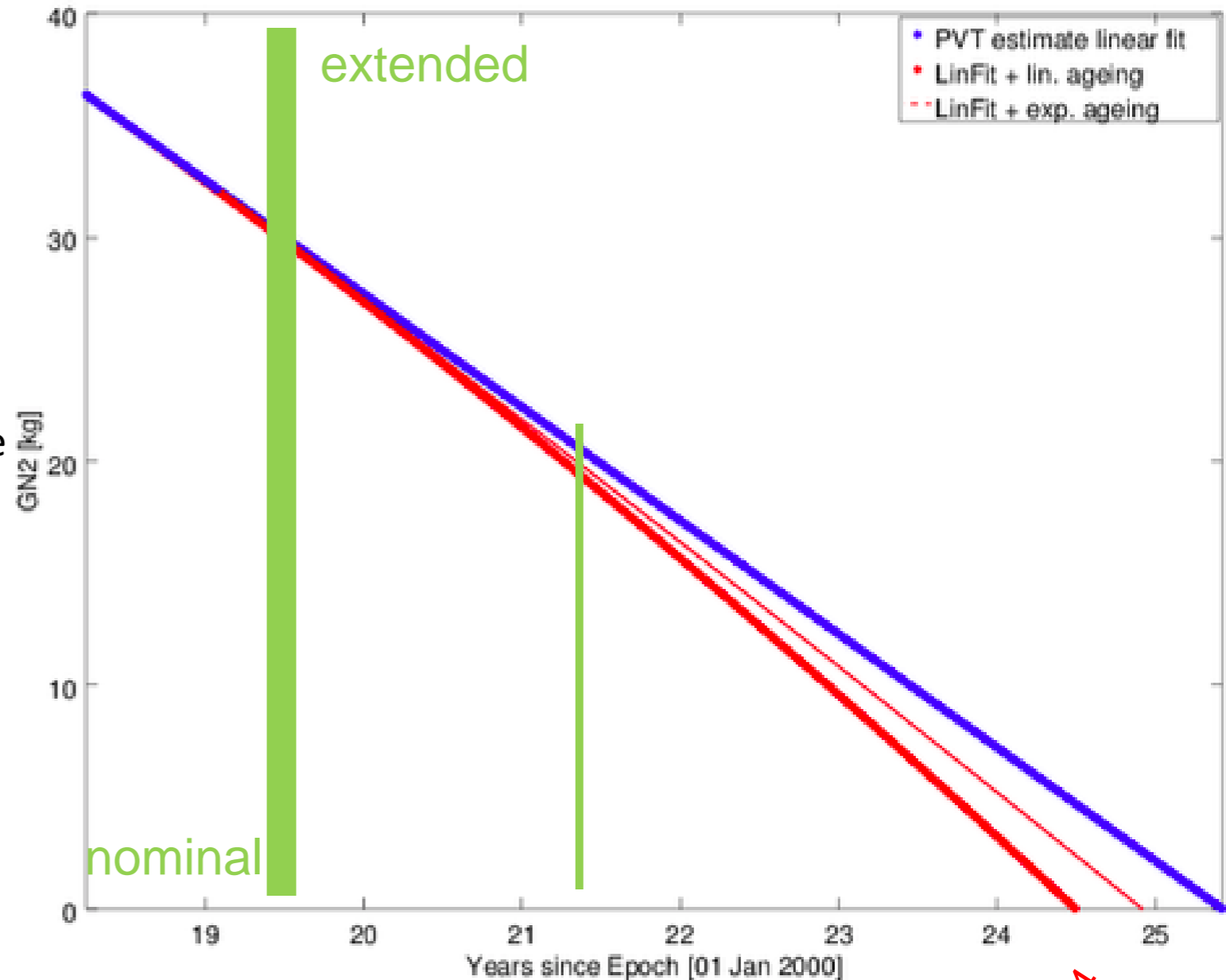
Timeline





Future - Extension

- Extension
-> 10 years expected
(depending cold-gas tank)
- More photons
faint-end objects performance
- Good for dynamics
 - SSOs
 - Exoplanets
 - Binaries
 - ...



- Need of DPAC operations continuation!..

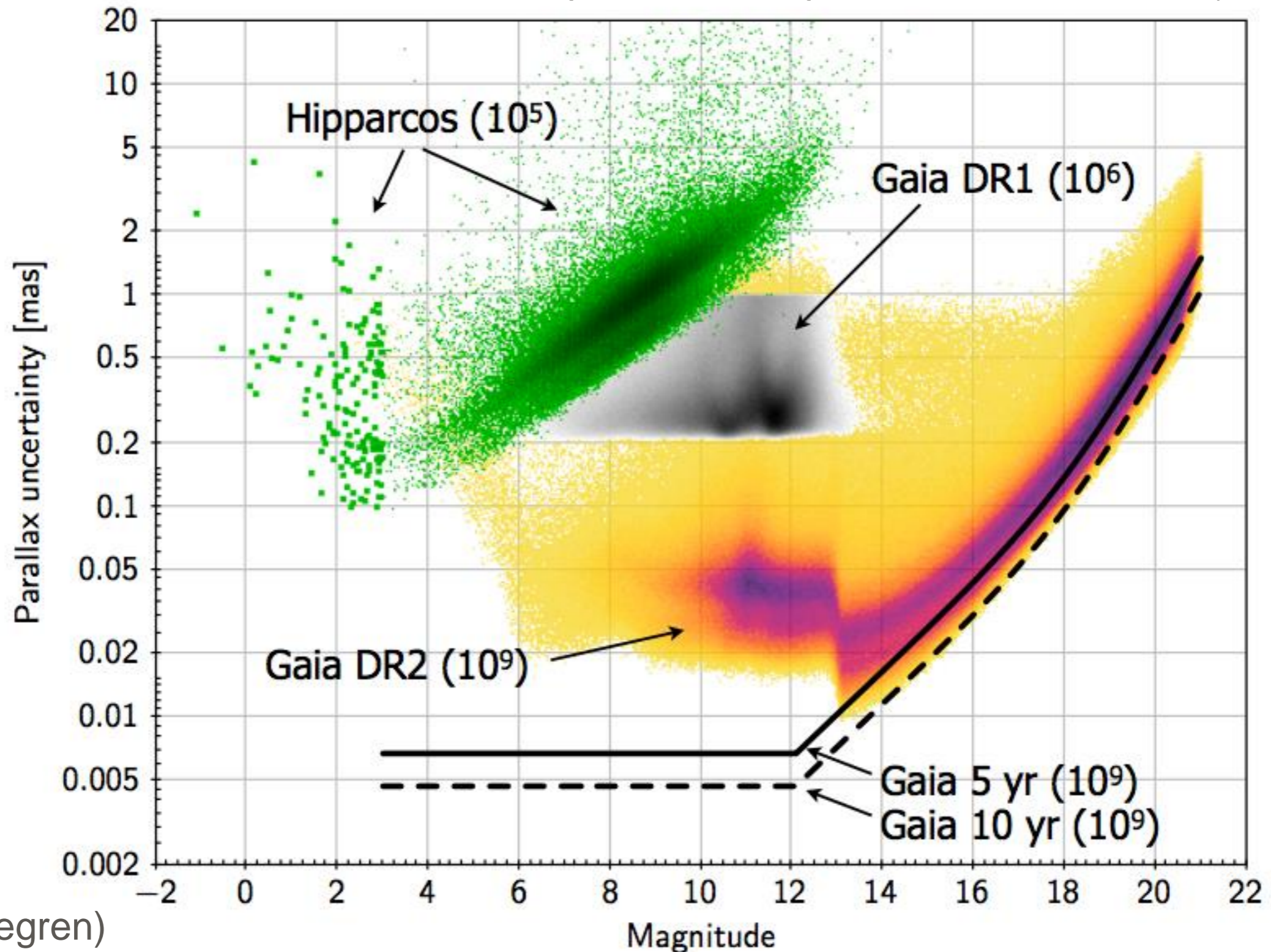
11/2024
±6months



Future – Extension

Parallax

Proper motion improvement even better ($\propto 3/2$)

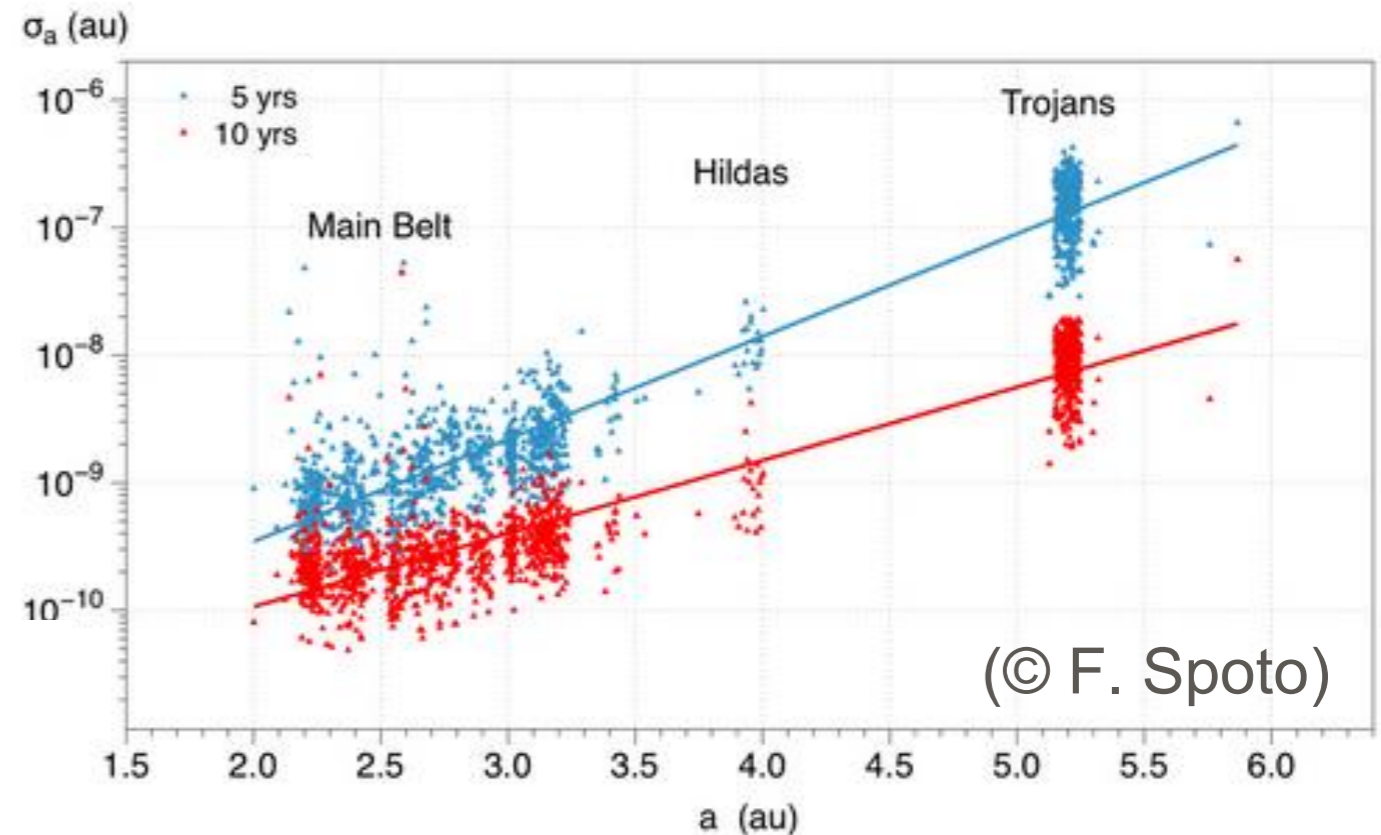
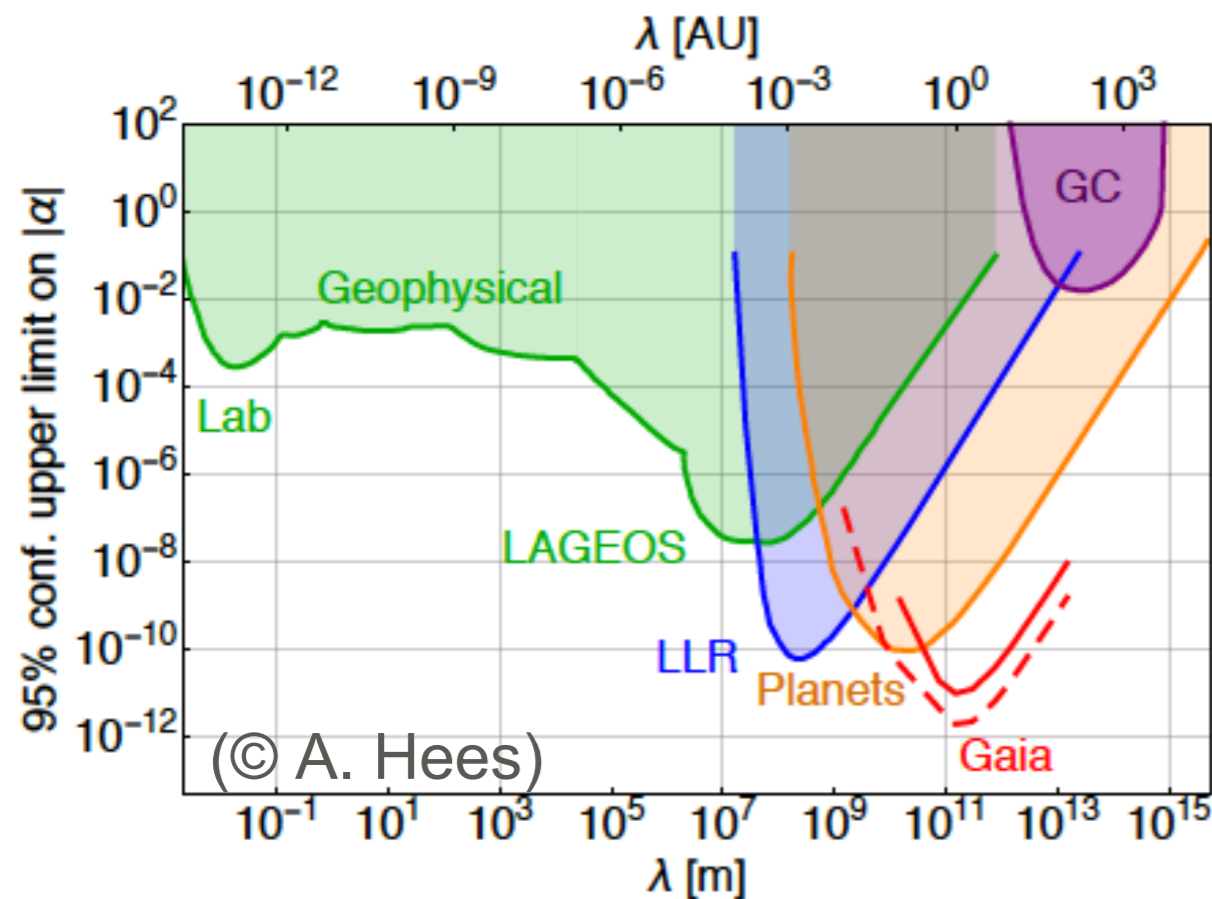


(@L. Lindegren)

Future - Extension



Dynamics of SSOs



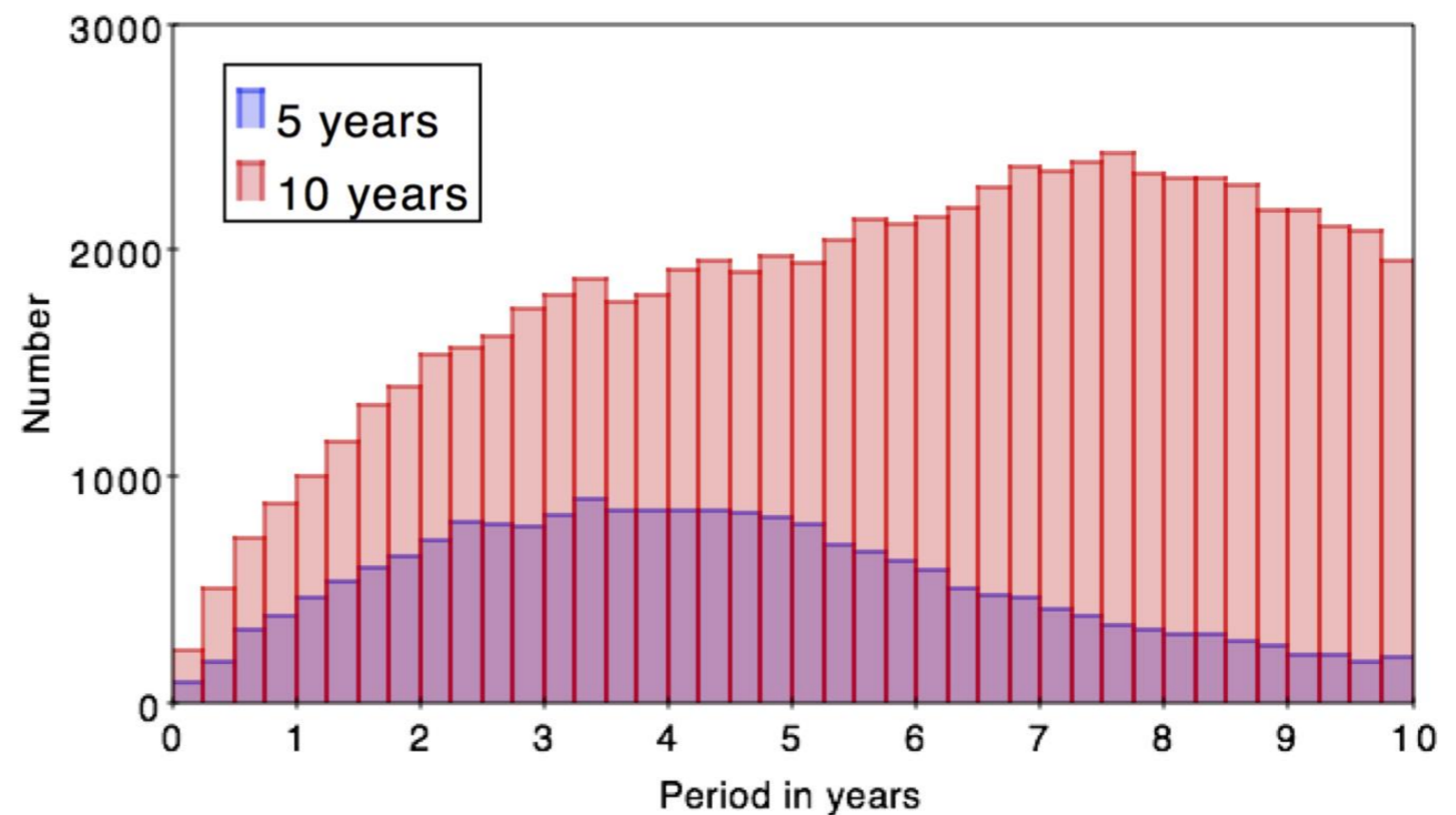
- SSOs orbits improvement
- Masses from encounters
- Tests of GR and gravitation



Future - Extension

- Gaia's strength is Neptune-Jupiter mass planets around stars
- Mission extension reveals population of giant planets above several AU distances from the parent star
 - ▶ giant planets before migration, systems with giant 'guarding' habitable zone
- Exoplanets research gains enormously from the improved parallaxes helping to describe the host star

Exoplanets



(@A. Brown)



Conclusions

- Gaia DR2 stars for occultations ; for catalogue astrometric corrections
- Gaia DR2 asteroid data: the first sample
already useful to test some dynamical effects
 - general orbit improvement, stellar occultations, Yarkovsky, ...
- Combination with other observations
- Past => accurate debiasing required (or re-reduction)
- Future ground-based spectro surveys, LSST, space astrometry in IR...
- Then DR3 &+, not only astrometry,
not only asteroids !...





Gaia web & archive

<http://gea.esac.esa.int/archive/>

→ EUROPEAN SPACE AGENCY  ABOUT ESAC 

SIGN IN 

gaia archive



HOME SEARCH STATISTICS VISUALISATION DOCUMENTATION HELP

Welcome to the Gaia Archive

Gaia is an ambitious mission to chart a three-dimensional map of our Galaxy, the Milky Way, in the process revealing the composition, formation and evolution of the Galaxy. Gaia will provide unprecedented positional and radial velocity measurements with the accuracies needed to produce a stereoscopic and kinematic census of about one billion stars in our Galaxy and throughout the Local Group. This amounts to about 1 per cent of the Galactic stellar population.



Top Features



Citation

How to cite and acknowledge Gaia.



Search

Query for Gaia sources using an ADQL (Astronomical Data Query Language) interface in an asynchronous mode (UWS).



Download

Direct download of Gaia data files.



Help

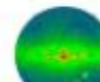
For questions, suggestions or problem reports, contact the Helpdesk.



Documents



Gaia Mission



Statistics





Partners




Gaia web & archive

<http://gea.esac.esa.int/archive/>

→ EUROPEAN SPACE AGENCY 

ABOUT ESAC 

SIGN IN 


gaia archive

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Documents



Gaia Mission

Position File

☒ Name
☐ Equatorial

Target in ☒ Circle ☐ Box

Name for All Radius 5 arc sec

Search in:

Extra conditions

Display columns

<input checked="" type="checkbox"/> solution_id	<input checked="" type="checkbox"/> source_id	<input type="checkbox"/> observation_id	<input checked="" type="checkbox"/> number_mp	<input checked="" type="checkbox"/> epoch
<input type="checkbox"/> epoch_err	<input type="checkbox"/> epoch_utc	<input checked="" type="checkbox"/> ra	<input checked="" type="checkbox"/> dec	<input type="checkbox"/> ra_error_systematic
<input type="checkbox"/> dec_error_systematic	<input checked="" type="checkbox"/> ra_dec_correlation_systematic	<input checked="" type="checkbox"/> ra_error_random	<input checked="" type="checkbox"/> dec_error_random	<input type="checkbox"/> ra_dec_correlation_random
<input checked="" type="checkbox"/> g_mag	<input type="checkbox"/> g_flux	<input type="checkbox"/> g_flux_error	<input checked="" type="checkbox"/> x_gaia	<input checked="" type="checkbox"/> y_gaia
<input checked="" type="checkbox"/> z_gaia	<input checked="" type="checkbox"/> vx_gaia	<input checked="" type="checkbox"/> vy_gaia	<input checked="" type="checkbox"/> vz_gaia	<input checked="" type="checkbox"/> position_angle_scan
<input type="checkbox"/> level_of_confidence				

☐ Select All / None

Max. number of results: 500

Reset Form

Show Query

Submit Query

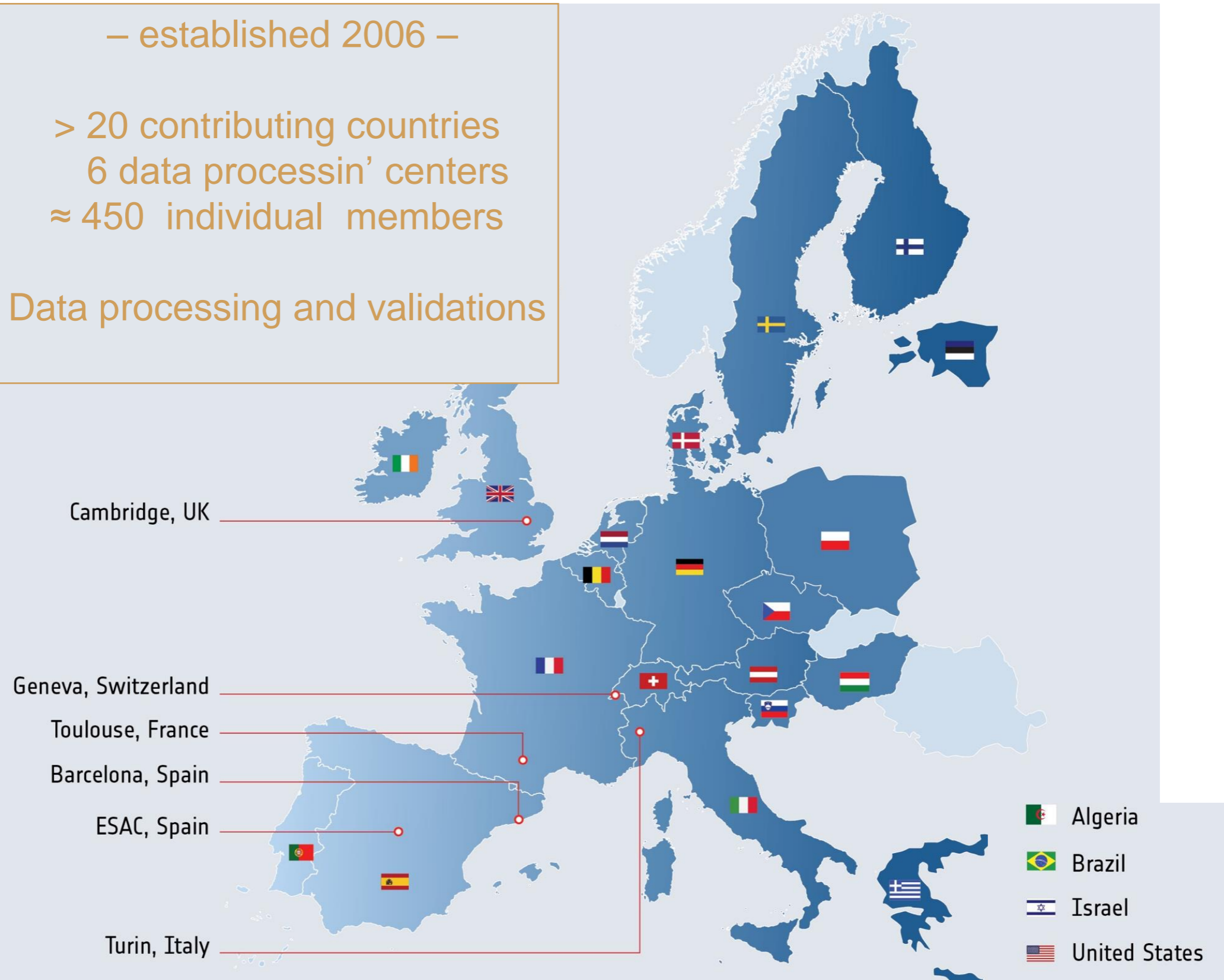


DPAC consortium

— established 2006 —

> 20 contributing countries
6 data processin' centers
≈ 450 individual members

Data processing and validations





Thanks to DPAC



‘At the back of your paper’: You too, acknowledge DPAC !



Citation

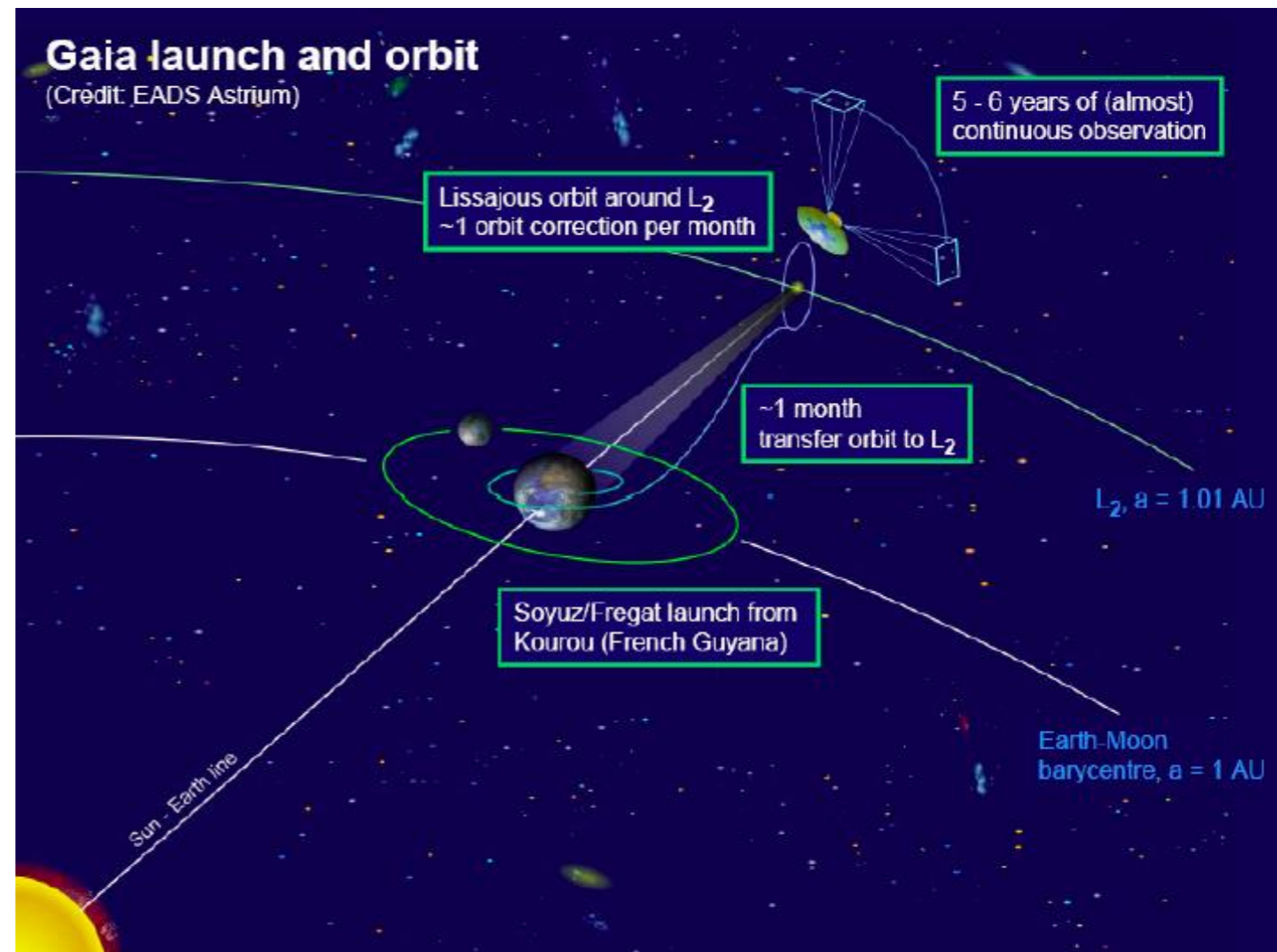
How to cite and acknowledge Gaia.

“This work has made use of data from the European Space Agency (ESA) mission Gaia (<https://www.cosmos.esa.int/gaia>), ...”



Operations

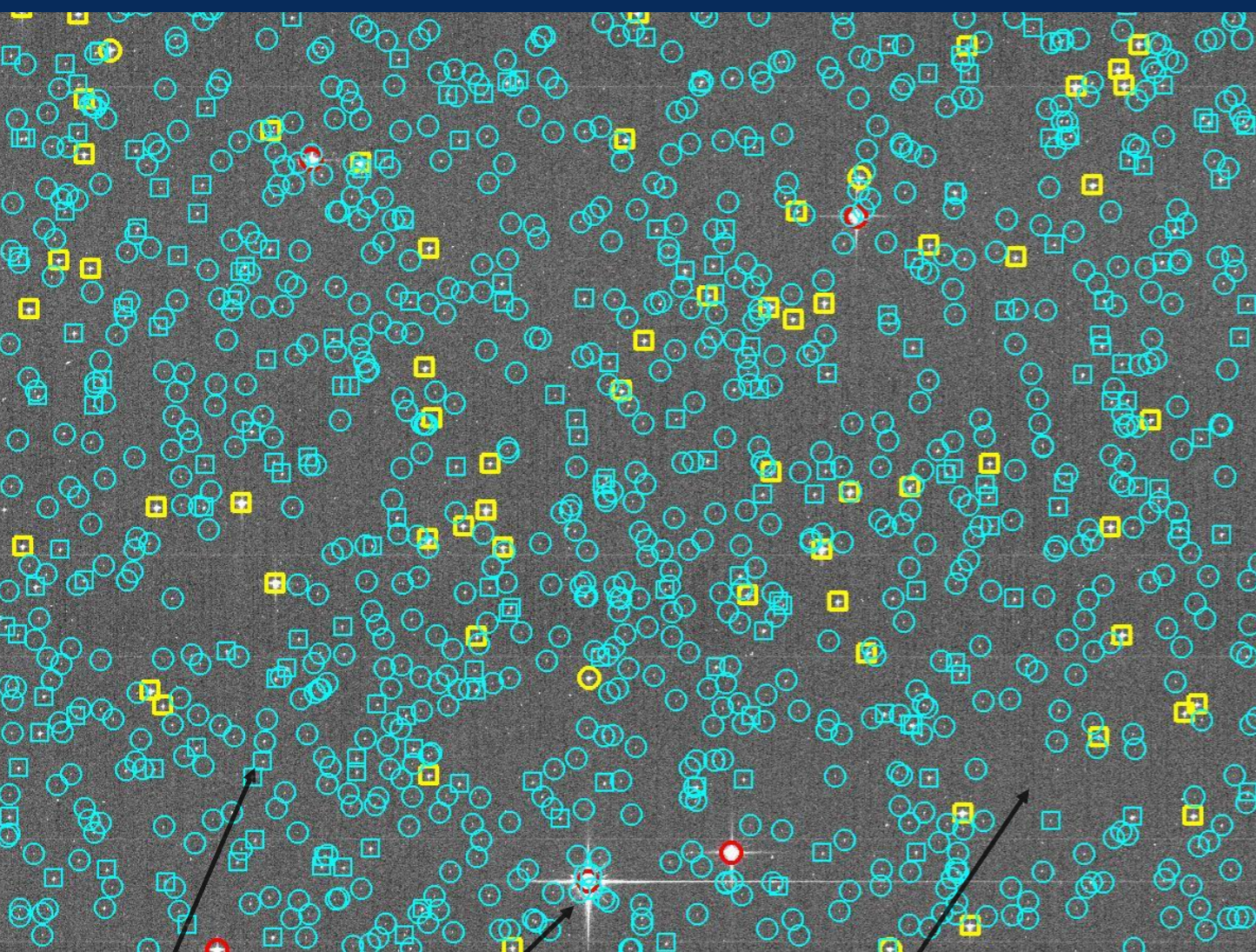
- Send to L2
- Science since mid-2014
now 56months
- Everything works fine
 - stray light - did impact RVS
 - contamination – OK, managed
 - radiation damage – factor 6 below expectations
 - micro-meteorites & micro-clanks





Focal plane - Payloads

Automatic source detection

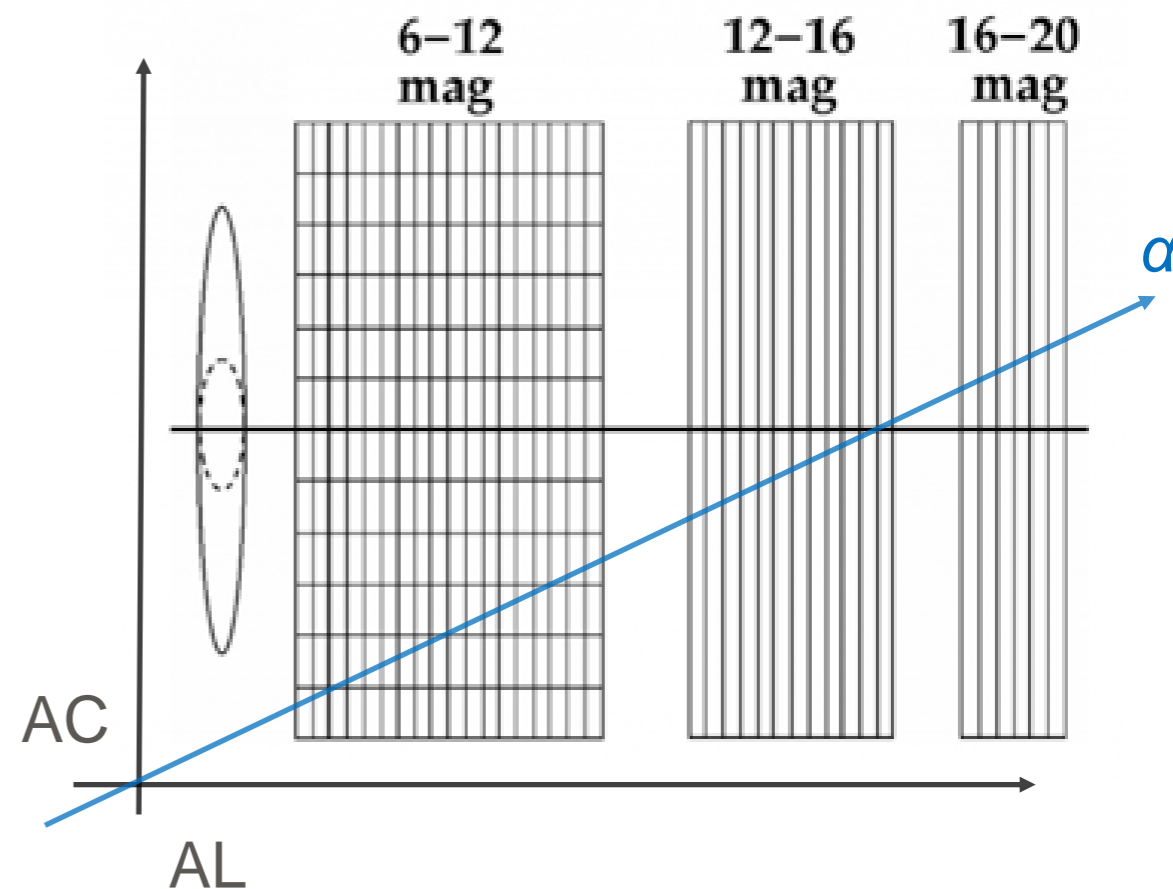


deal with confusion
window sent to ground

pixel not sent

Windows and patches

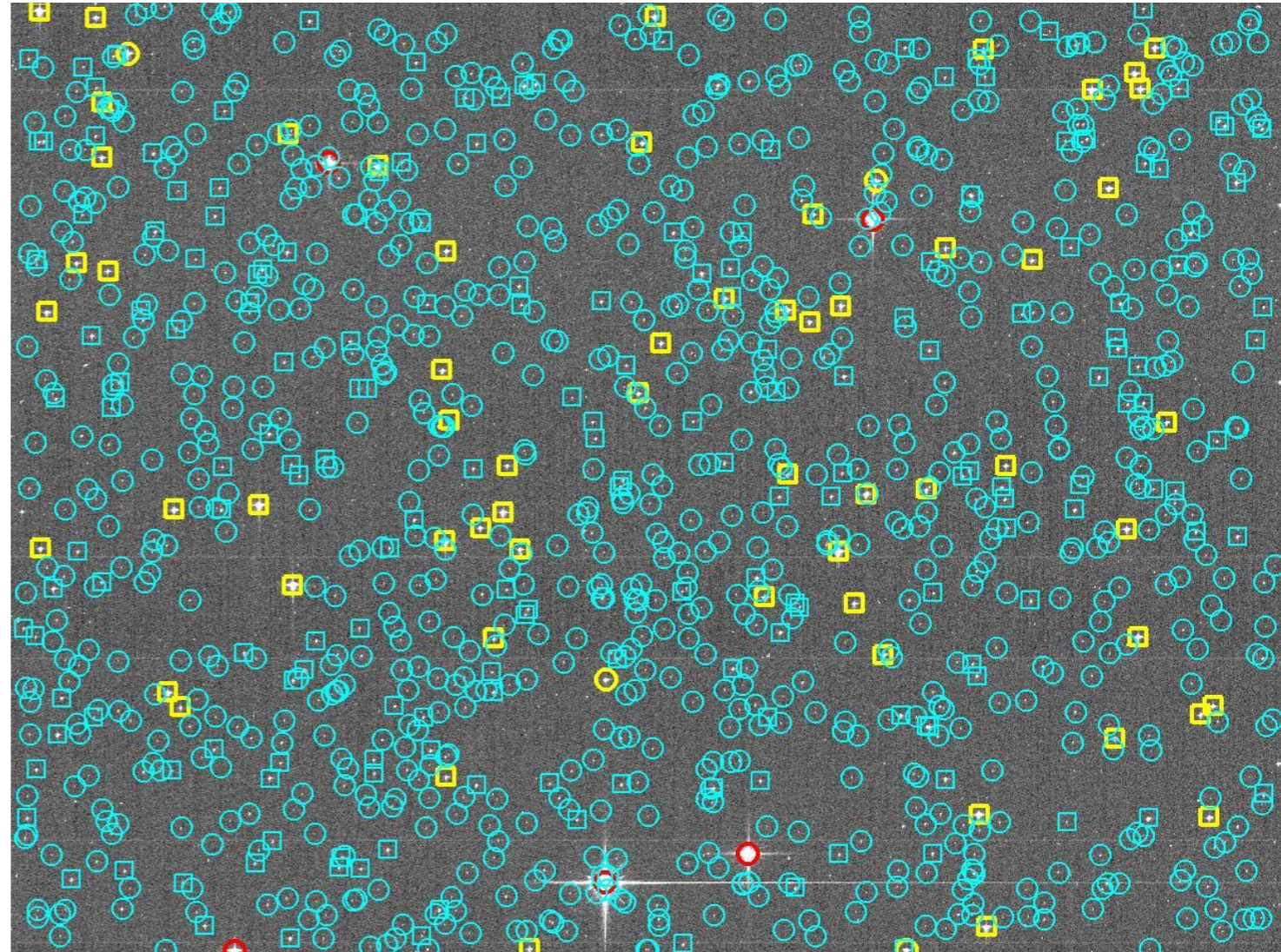
- no data compression
not all Gpix. sent to ground
- adapted to stars motion and TDI
not adapted to fast moving objects
- pixels binning \Rightarrow 1D signal
highly correlated (α, δ)





Automatic detection

- Sky mapper detection based on PSF pick not SNR limited
- Patch around source transmitted to ground based on scanning
- No large source ($\approx 0.9''$)
- No blurred source
- TDI mode not adapted for fast moving objects



Summary - Future



- Much more about DR2 <https://www.cosmos.esa.int/web/gaia/data-release-2>
- Gaia archive <http://gea.esac.esa.int/archive/>
- Recent IAU Div.A days session https://www.iau.org/science/scientific_bodies/divisions/A
- Conf. ESLAB#53 The Gaia Universe 8-12 April 2019
<https://www.cosmos.esa.int/web/the-gaia-universe/home>

- Future GDRs
 - Incremental, improvement
 - More data! and solutions
 - Astrophys. parameters
 - Exoplanets
 - Asteroids, satellites, comets
 - All epoch and transit data

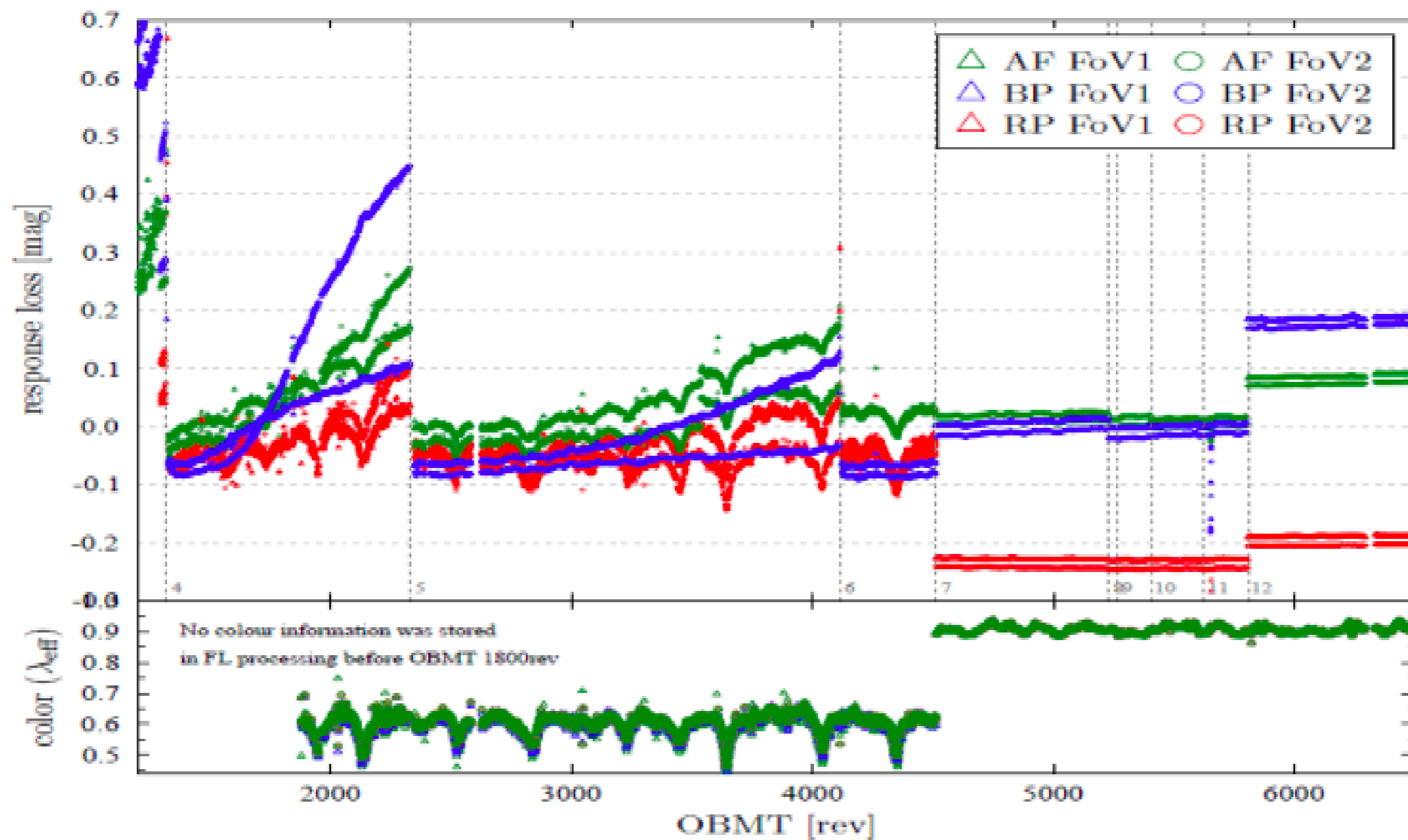
Astrometry	●	●	●	●	●
Spectro-photometry	●	●	●	●	●
Spectroscopy	●	●	●	●	●
Variability	●	●	●	●	●
Binary solutions	●	●	●	●	●
Astrophysical parameters	●	●	●	●	●



- Ground-based spectroscopic surveys, LSST
 - deeper Northern survey - Continuation to faint end
 - high accuracy epoch astrometry + photometry
 - 37 billion stars+galaxies, 10 years , 15Pb
- Space astrometry in IR



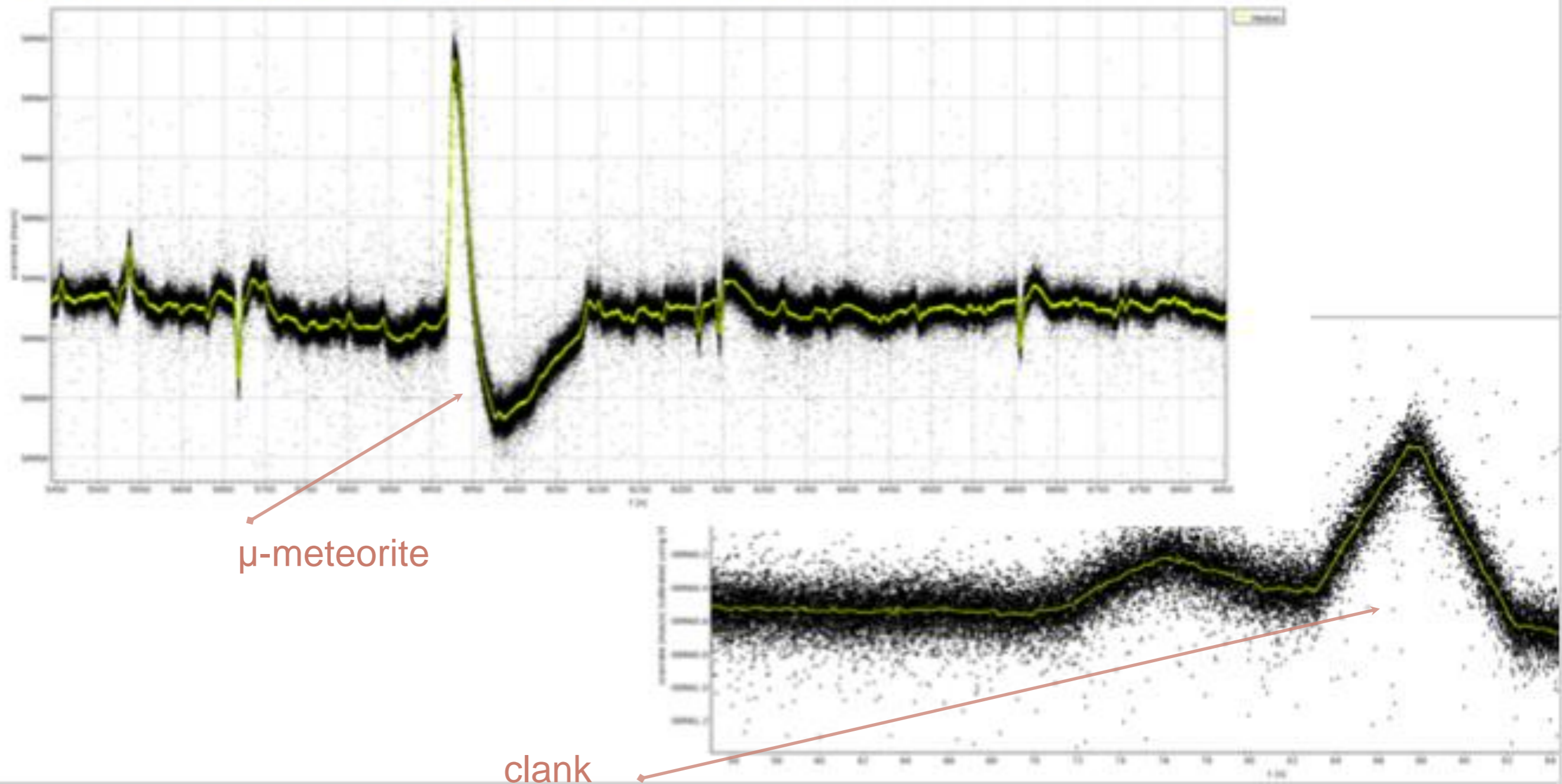
Contamination





Fine tuning of scan

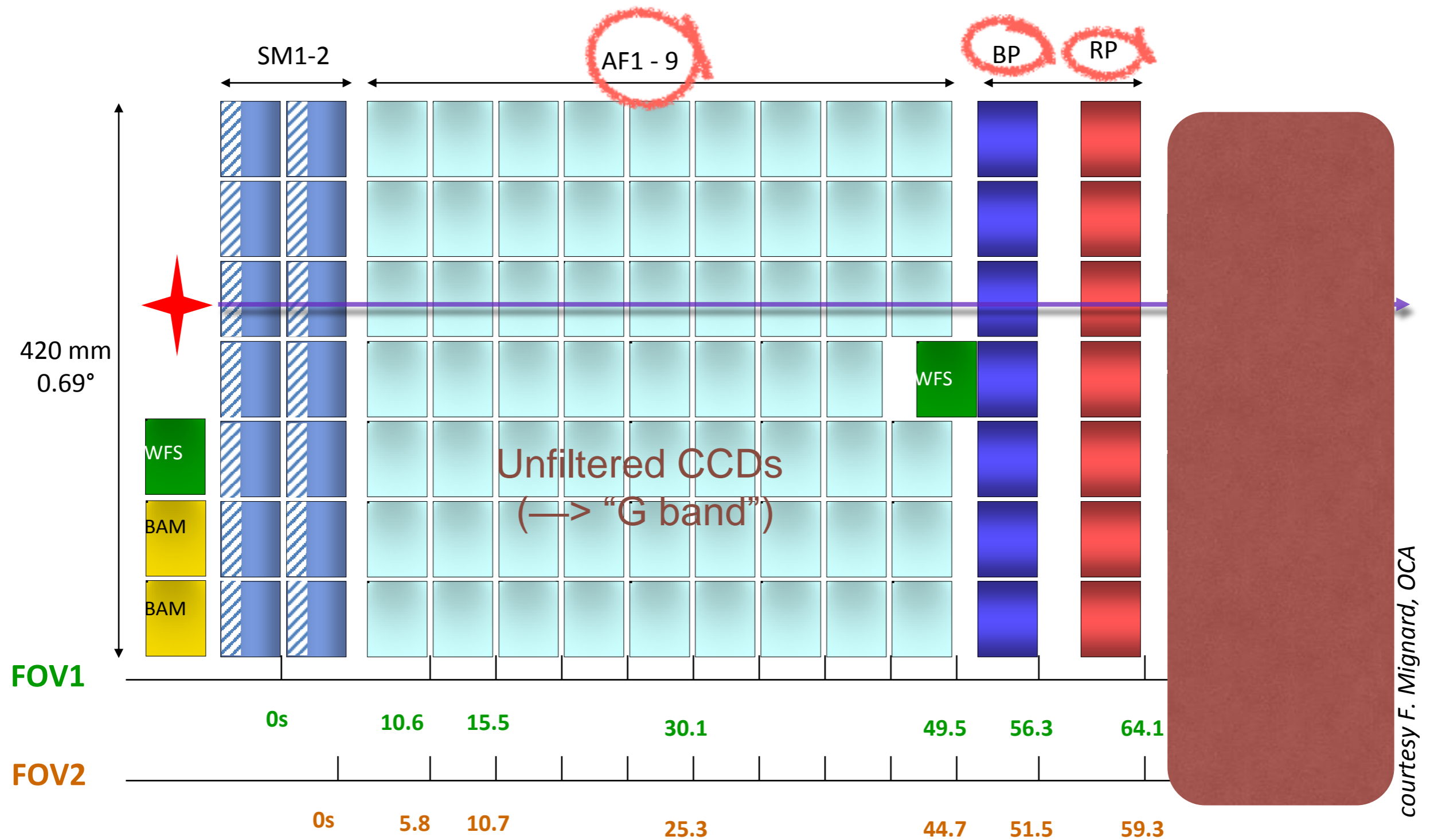
Micro-meteoroids and micro-clanks



Probably common on s/c, but Gaia is the first one to be able to measure it!



An observation = a transit = 10 positions + spectro-photometry





First Yarkovsky measurement by Gaia: (2062) Aten

- ground-based observations from 1955 to 2017
- good pre-Gaia orbital fit with Yarkovsky measurement

	SNR_{A2}	da/dt $10^{-4} au/Myr$
69 Gaia + 959 optical + 7 radar	14.63	-4.90 ± 0.34
959 optical + 7 radar	8.64	-5.98 ± 0.68

Spoto et al. 2018

- The combination of Gaia + other data significantly changes the value and improves SNR
- A coherent debasing is fundamental: only the *DR2 scheme* correctly combines Gaia + other observations (—> no rejection of DR2 astrometry)
- Note: $\sigma_a = 1 \times 10^{-10}$ au !!!

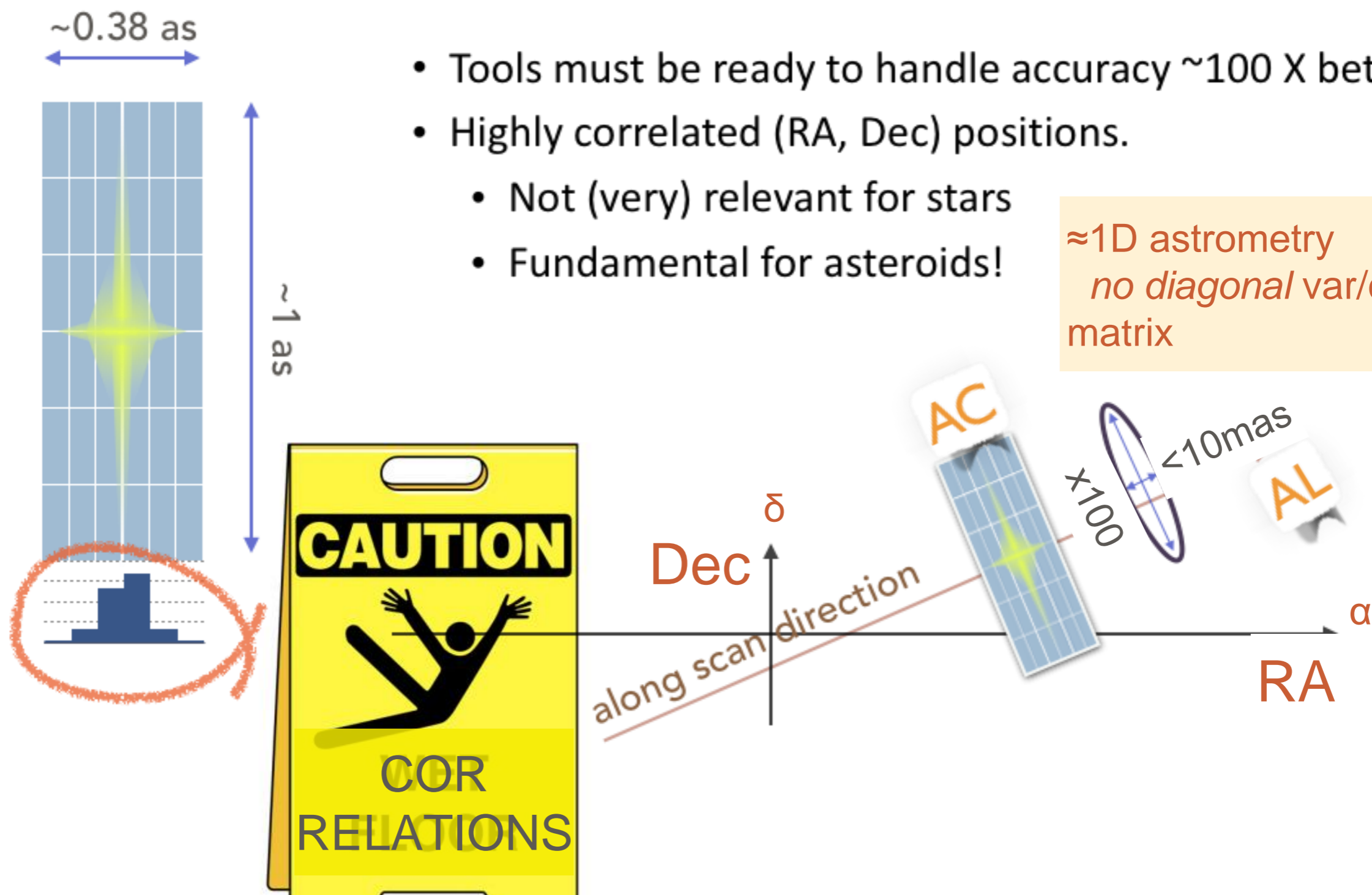


Gaia DR2 – SSO

CAUTION

- Tools must be ready to handle accuracy ~ 100 X better
- Highly correlated (RA, Dec) positions.
 - Not (very) relevant for stars
 - Fundamental for asteroids!

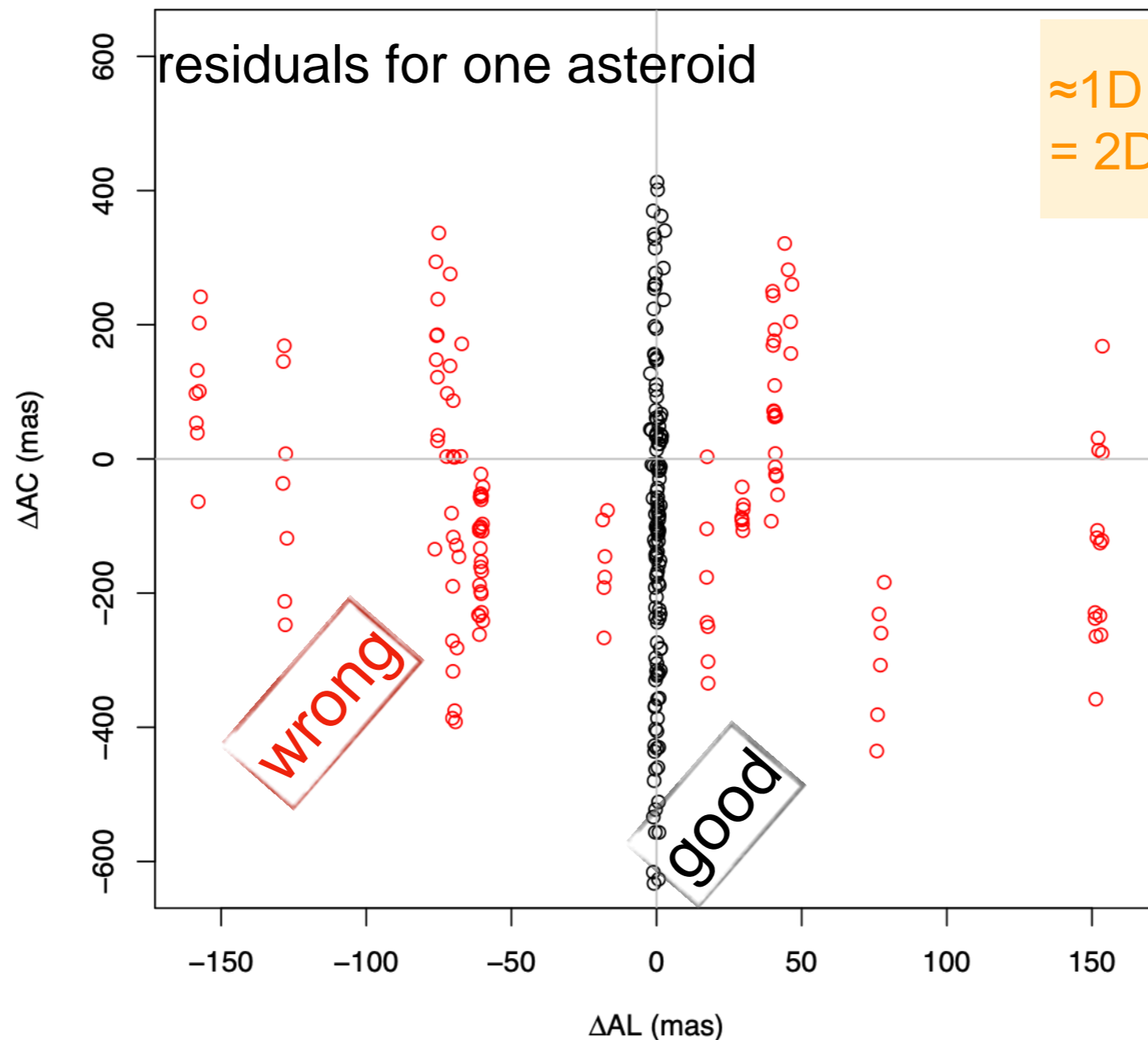
$\approx 1D$ astrometry
no diagonal var/covar matrix



Gaia DR2 – SSO

CAUTION

- o W = diagonal
- o W = var/covar



$\approx 1D$ astrometry in (AL, AC)
= 2D var/covar matrix (α , δ)





Conclusions

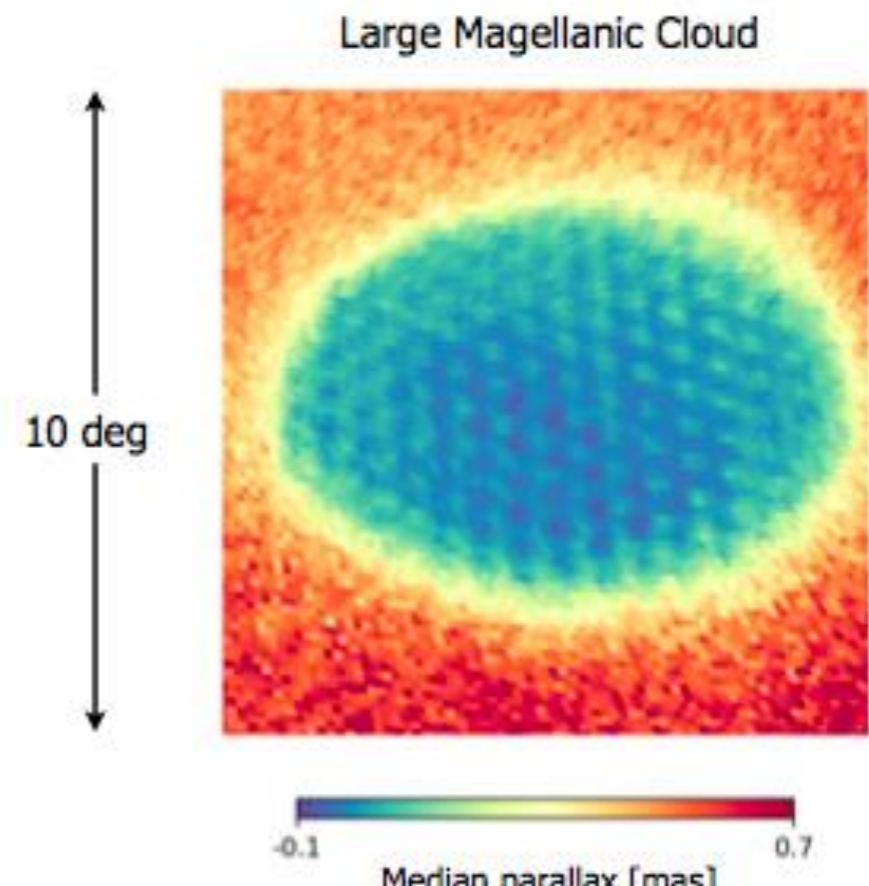
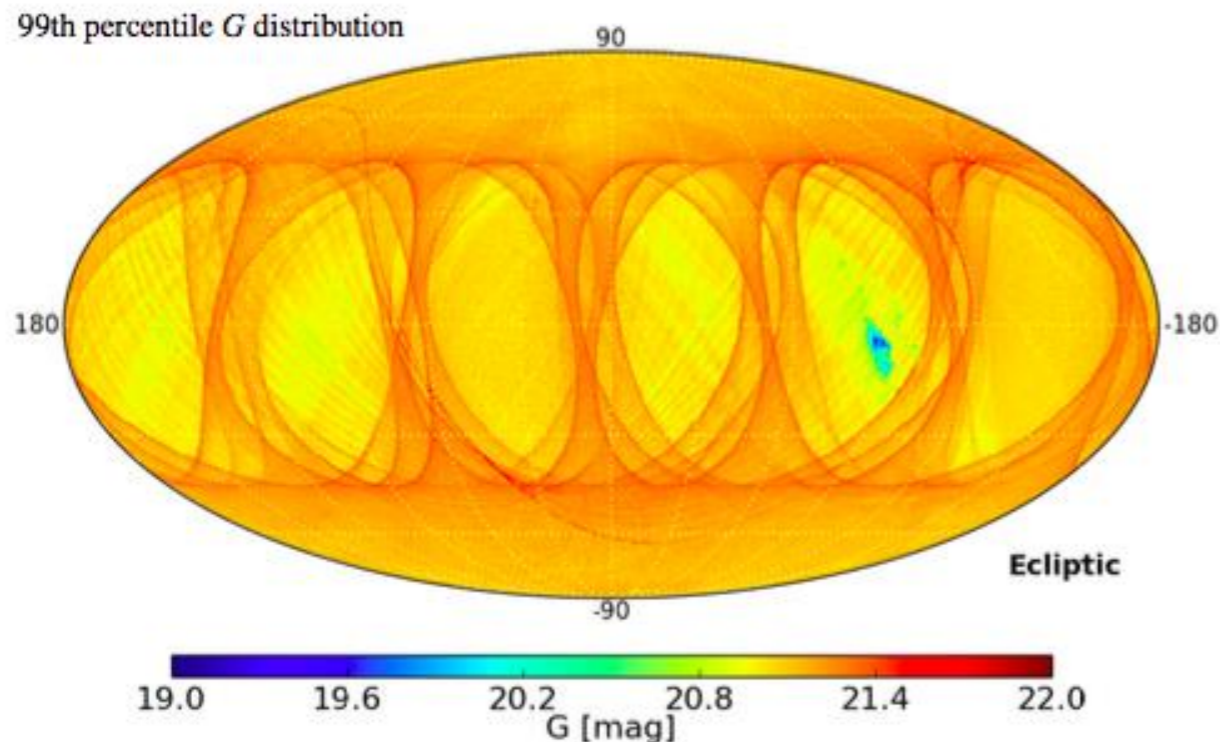
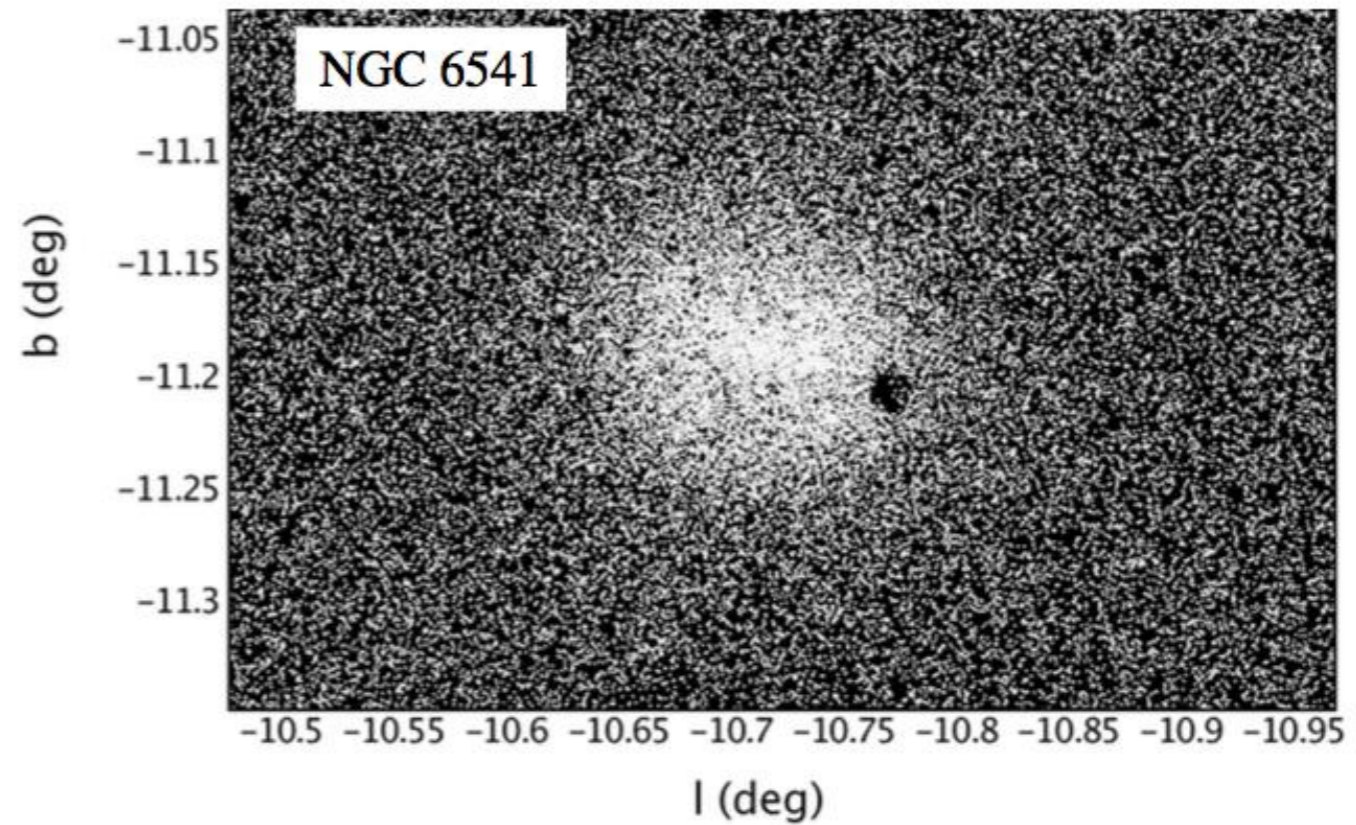
- Gaia DR2 asteroid data: the first sample, already useful to test subtle dynamical effects
 - Yarkovsky determination
 - general orbit improvement
 - application to stellar occultations
 - Combination with other observations: it starts to work...
 - but weighting of the data is critical
 - accurate debiasing is required
 - a new method successfully implemented and tested
- Yarkovsky detection *in the Main Belt* is getting closer...
- Other dynamical effects to come!
- Then, not only astrometry, not only asteroids !...



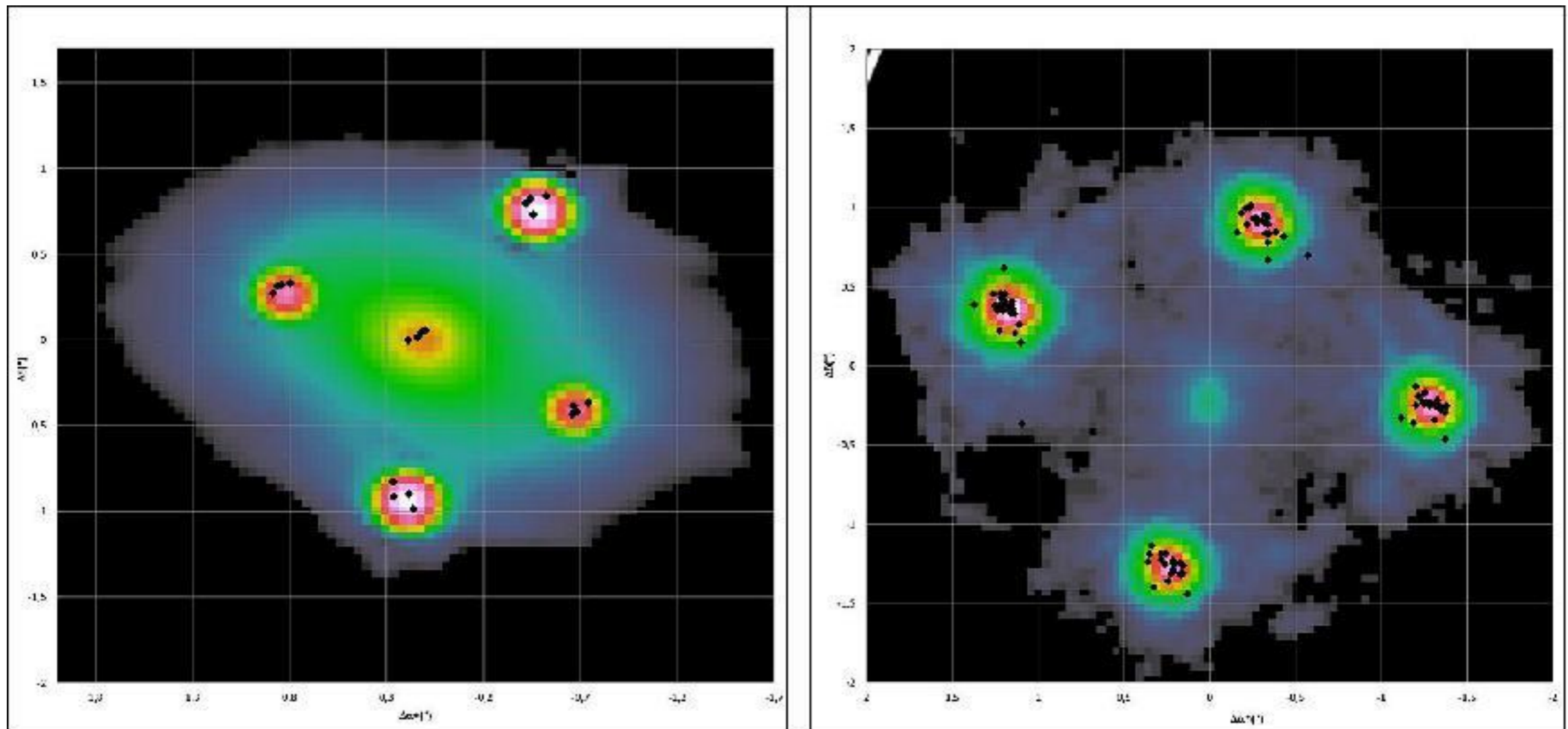


DR2 - Shortcomings

- Some caveats
 - bright stars calibration
 - crowded regions,
 - fast moving stars
 - scanning law systematics
 - Varying magnitude limit



Gravitational lenses multi-imaged QSO

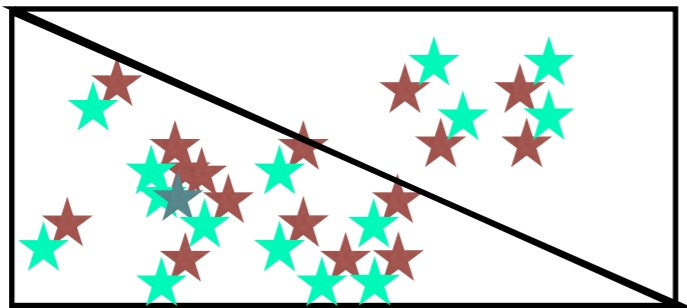


Einstein Cross (left) and HE0435-1223 (right) (@C. Ducourant)
with Gaia astrometric positions placed over HST images
and galaxy visible at centre

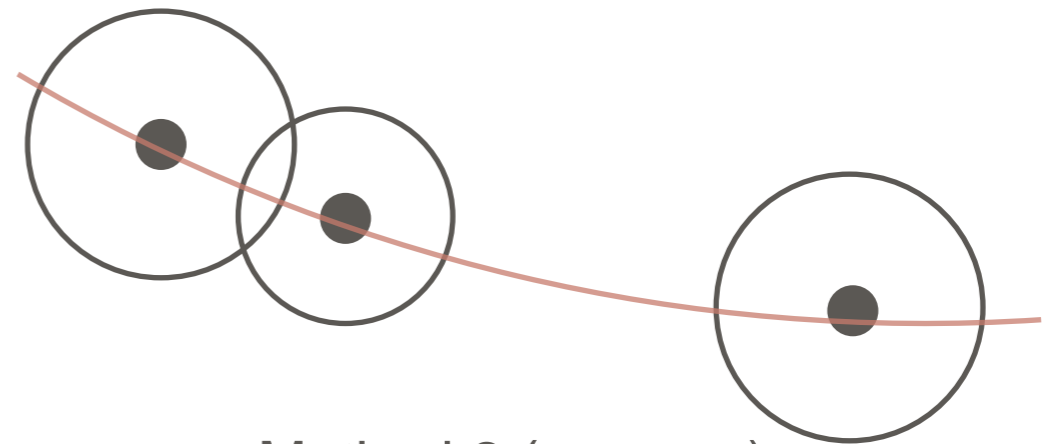


Orbit improvement : DR2 + other observations

- A factor ~ 2 (**only**) average improvement by using DR2 + all other data
- But: most other data are affected by systematic (zonal) errors of the pre-Gaia catalogs
- **Debiasing required!**
 - exploit the catalog information provided by MPC with the asteroid data
 - compute local biases per catalog (positions and pm) and correct observations



Method 1 : Farnocchia et al. (2015)
all catalogs referred to PPMXL
corrections computed on a grid
(F2015 scheme)

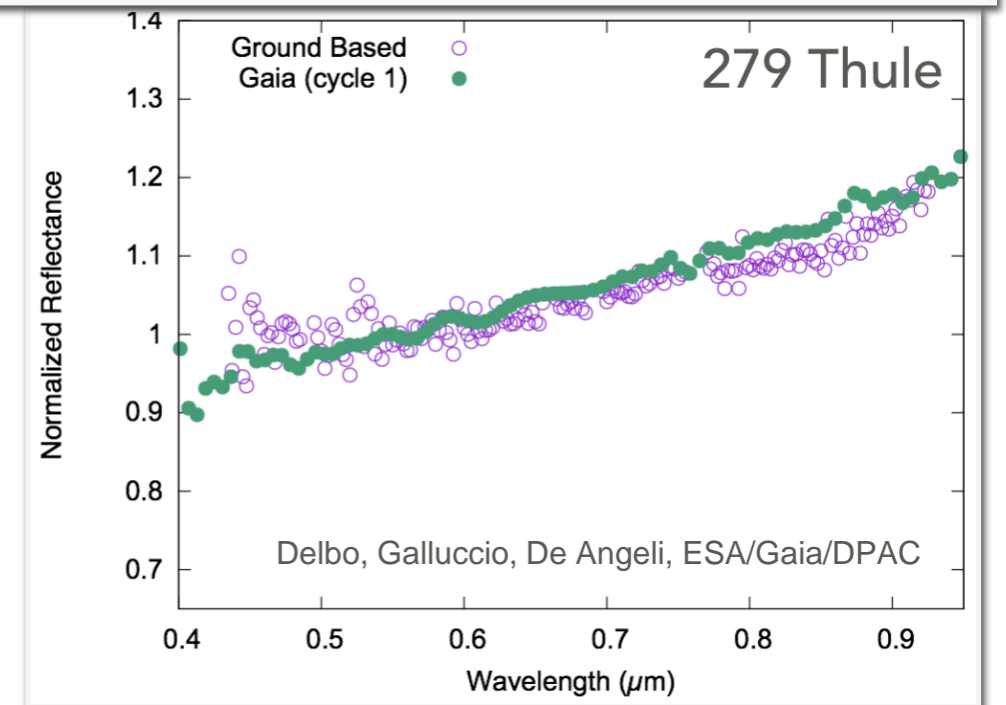
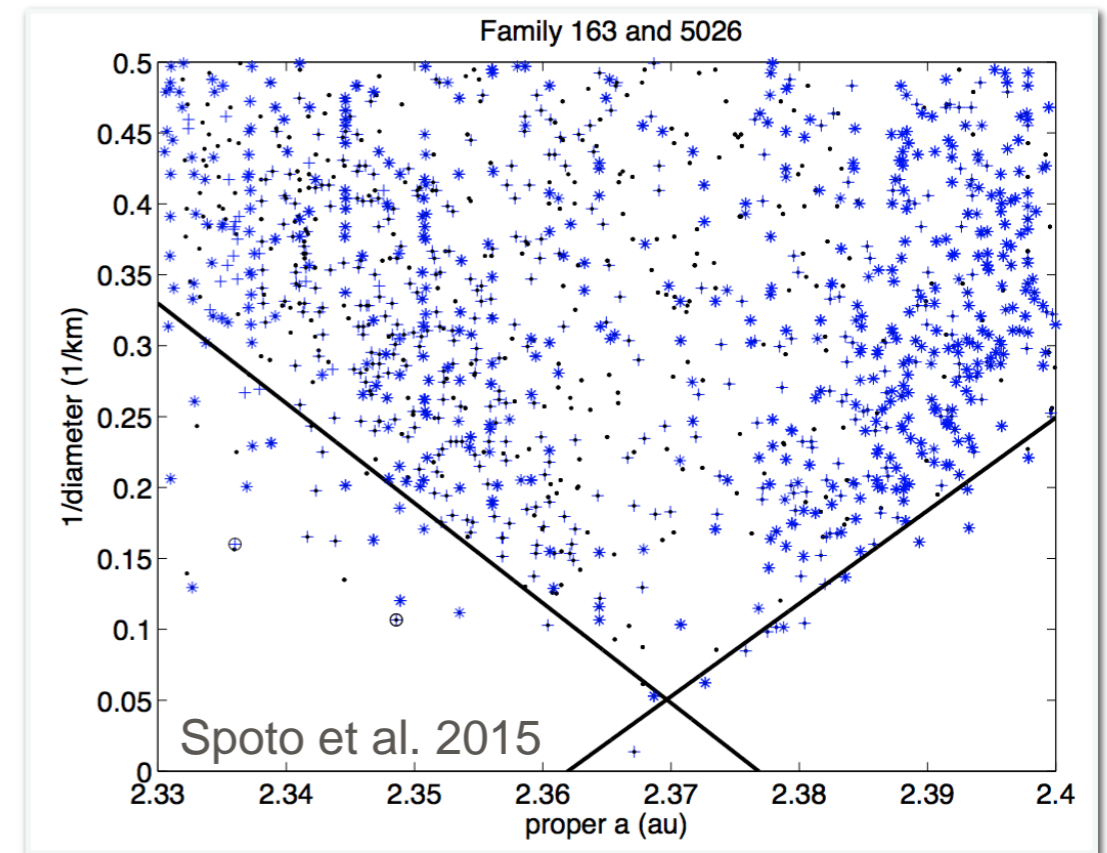


Method 2 (our own) :
areas centered on archive asteroid observations
Reference: Gaia DR2
(DR2 scheme)



Yarkovsky effect: why we care

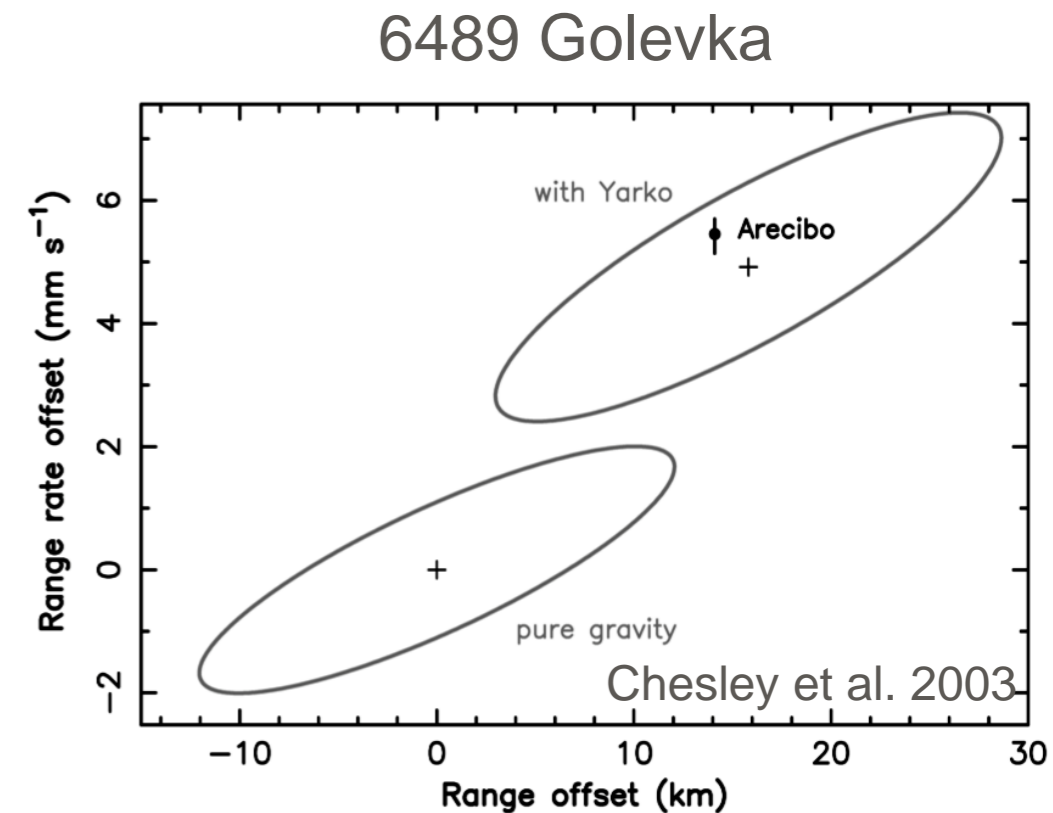
- Yarkovsky applications:
 - NEO and meteorite transport
 - physical properties (spin, density...)
 - family dispersion, ages
- connection to spectro- photometry by Gaia:
 - low-res spectra: taxonomy in the visible
 - *mmag* photometry: shape determinations
- Large sample by end of mission: $\approx 350k$





The problem of measuring Yarkovsky

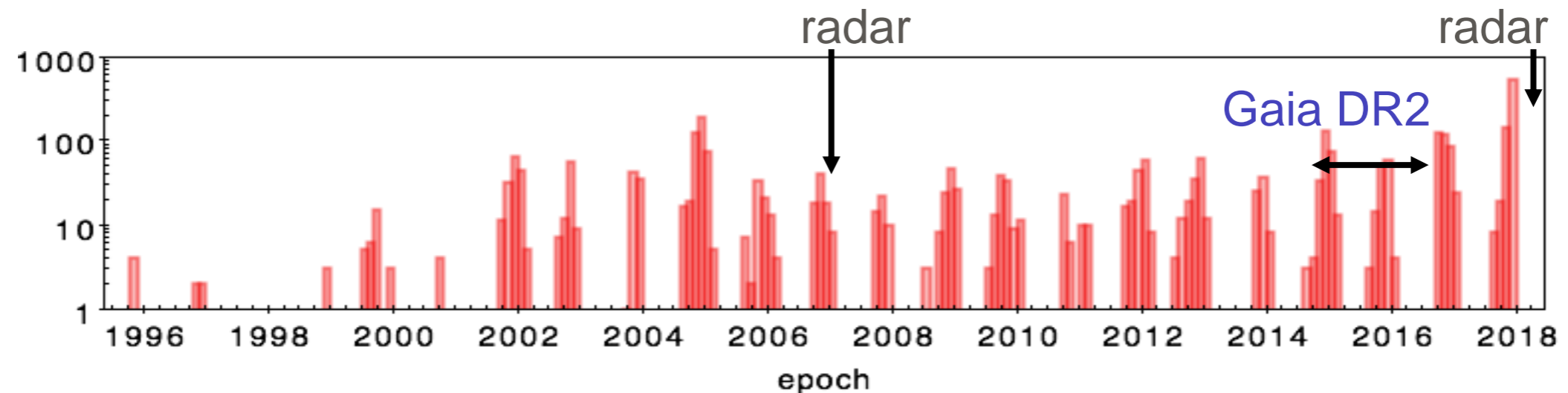
- Weak semi-major axis drift $\sim 10^{-4}$ au/Myr (at 1 au)
 - decreasing with size ($1/D$) and distance ($1/r_2$)
- It requires:
 - accurate observations
 - ...spanning a « large » interval of time
 - accurate weighting
- Results:
 - **36 NEOs** with *valid* detections (SNR>3)
 - **4 with SNR > 10** (Vokrouhlický et al. 2015)
 -radar observations essential !!
- for a small fraction of asteroids, only, $\sigma_a < 10^{-9}$ au





Another important case: (3200) Phaethon

- B-type, associated to the Geminid meteor stream
- Close encounter in Dec. 2017 (radar + many optical data)



- Without Gaia : SNR ~ 3.4 (JPL)
- By using Gaia DR2 observations:
$$da/dt = -9.22 \pm 1.27 (10^{-4} \text{ au/Myr}) - \text{SNR} \sim 16$$
- A large amount of low-quality observations in 2017 « pollutes » the result