

# LINEARLY POLARIZED PROPERTIES AND ROTATION MEASURE STUDY OF PARSEC-SCALE AGN JETS

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Kravchenko+2017  
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# KEY ASPECTS OF THE STUDY

## GOAL:

STUDYING MAGNETIC FIELD STRUCTURE  
OF THE AGN JETS

## BY ANALYSING:

FRACTIONAL POLARIZATION:

RE/DEPOLARIZATION EFFECTS

FARADAY ROTATION

TRANSVERSE RM GRADIENTS

INTRINSIC ELECTRIC FIELD STRUCTURE

# SAMPLE & OBSERVATIONAL SETUP

## 20 SOURCES:

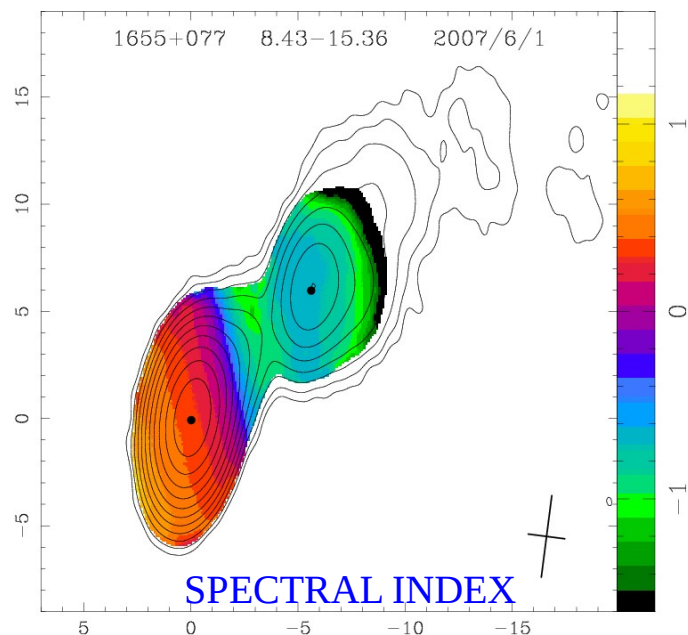
	<i>z</i>	
0148+274	1.260	QSO
0341+147	1.556	QSO
0425+048	0.517	AGN
0507+179	0.416	AGN
0610+260	0.580	QSO
0839+187	1.272	QSO
0952+179	1.478	QSO
1004+141	2.707	QSO
1011+250	1.636	QSO
1049+250	1.300	QSO
1219+285	1.161	BLL
1406-076	1.493	QSO
1458+718	0.904	QSO
1642+690	0.751	QSO
1655+077	0.621	QSO
1803+784	0.680	QSO
1830+285	0.594	QSO
1845+797	0.056	AGN
2201+315	0.298	QSO
2320+506	1.279	QSO

9 VLBA FREQUENCIES: 1.41, 1.66, 2.28, 2.39, 4.60, 5.00, 8.11, 8.43 (16 MHZ BANDWIDTH) & 15.4 (32 MHZ BANDWIDTH) GHZ

SINGLE EPOCH VLBI PARSEC-SCALE OBSERVATIONS IN 2007

SOURCES WITH HIGH CORE SHIFTS (SEE SOKOLOVSKY ET AL. 2011)

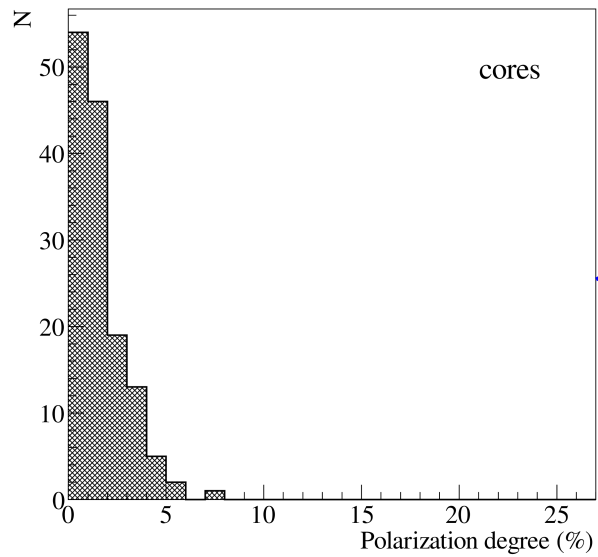
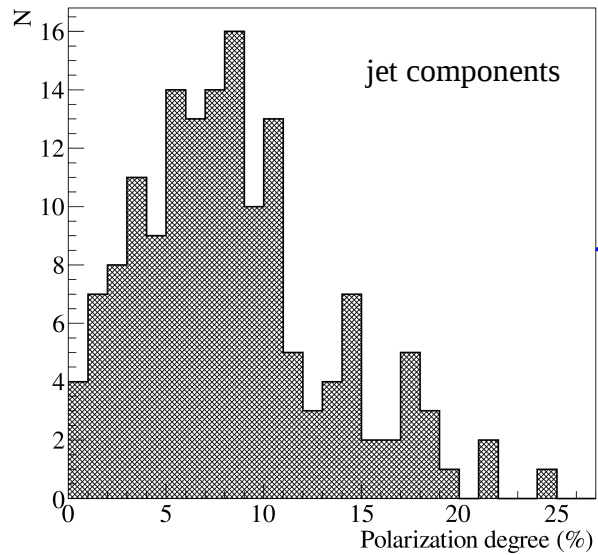
SEPARATE ANALYSIS FOR THE MODELLED OPAQUE CORE AND TRANSPARENT JET COMPONENTS, VISIBLE ACROSS 9 OBSERVABLE FREQUENCIES



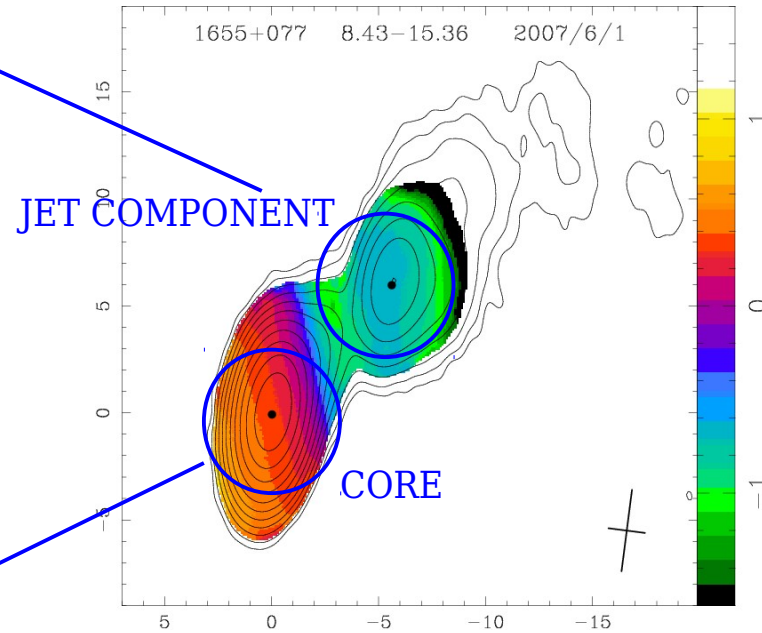
SIMILAR KPC STUDIES:  
FARNES+ 2014  
PASETTO+ 2015  
ANDERSON+ 2016

# FRACTIONAL POLARIZATION

## RESULTS: STATISTICS



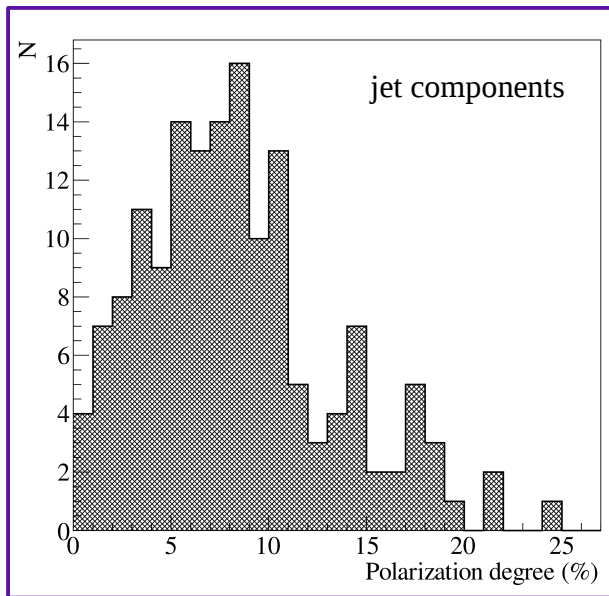
## SPECTRAL INDEX



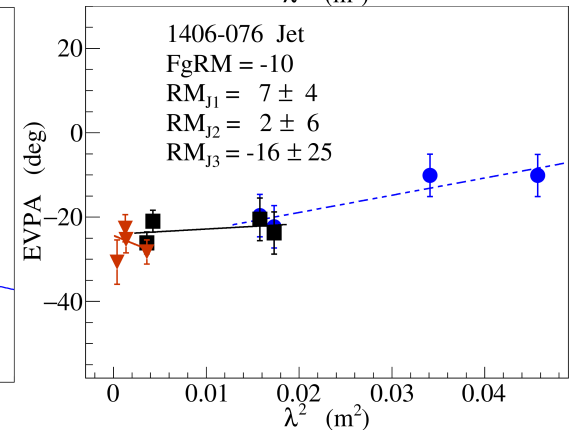
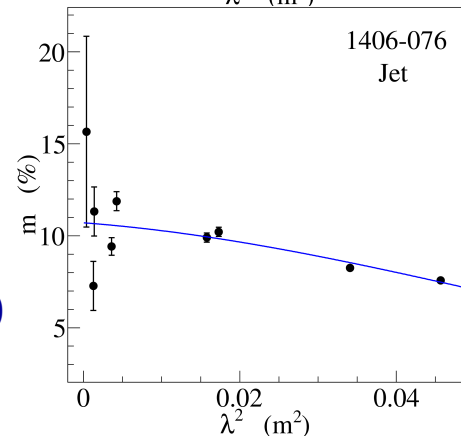
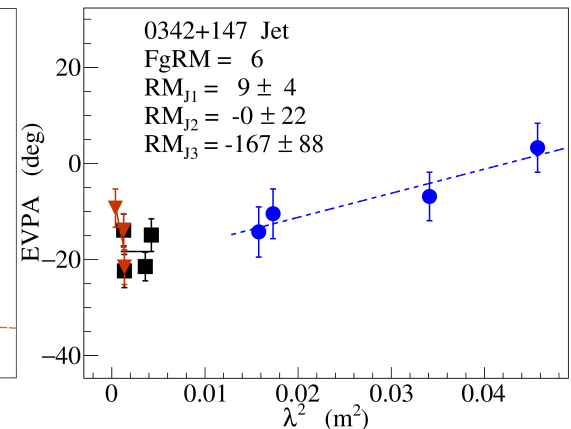
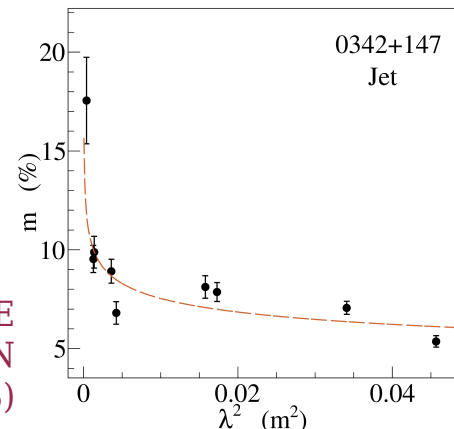
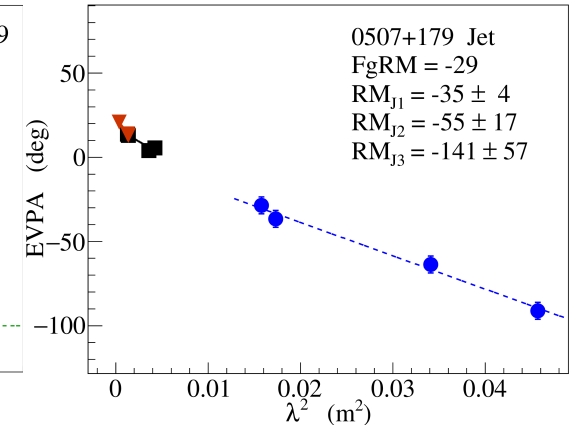
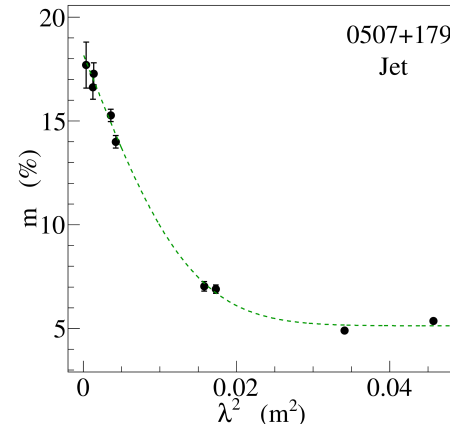
DISTRIBUTION ACROSS 9  
FREQUENCIES IN 20 SOURCES

# FRACTIONAL POLARIZATION & EVPA

EXTERNAL ROTATION: **TRANSPARENT** JET COMPONENTS



THIS DATA AND FIGURES ARE MADE FOR MODELLED OPTICALLY THIN JET COMPONENTS (PIXEL VALUES)

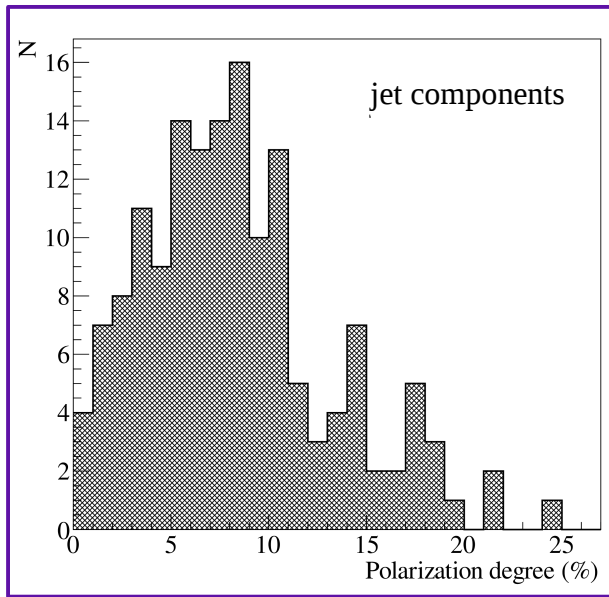


MAJORITY TRANSPARENT JETS EXHIBITS ROTATION ON EXTERNAL FARADAY SCREENS, WHILE FEW SOURCES SHOW INTERNAL ROTATION

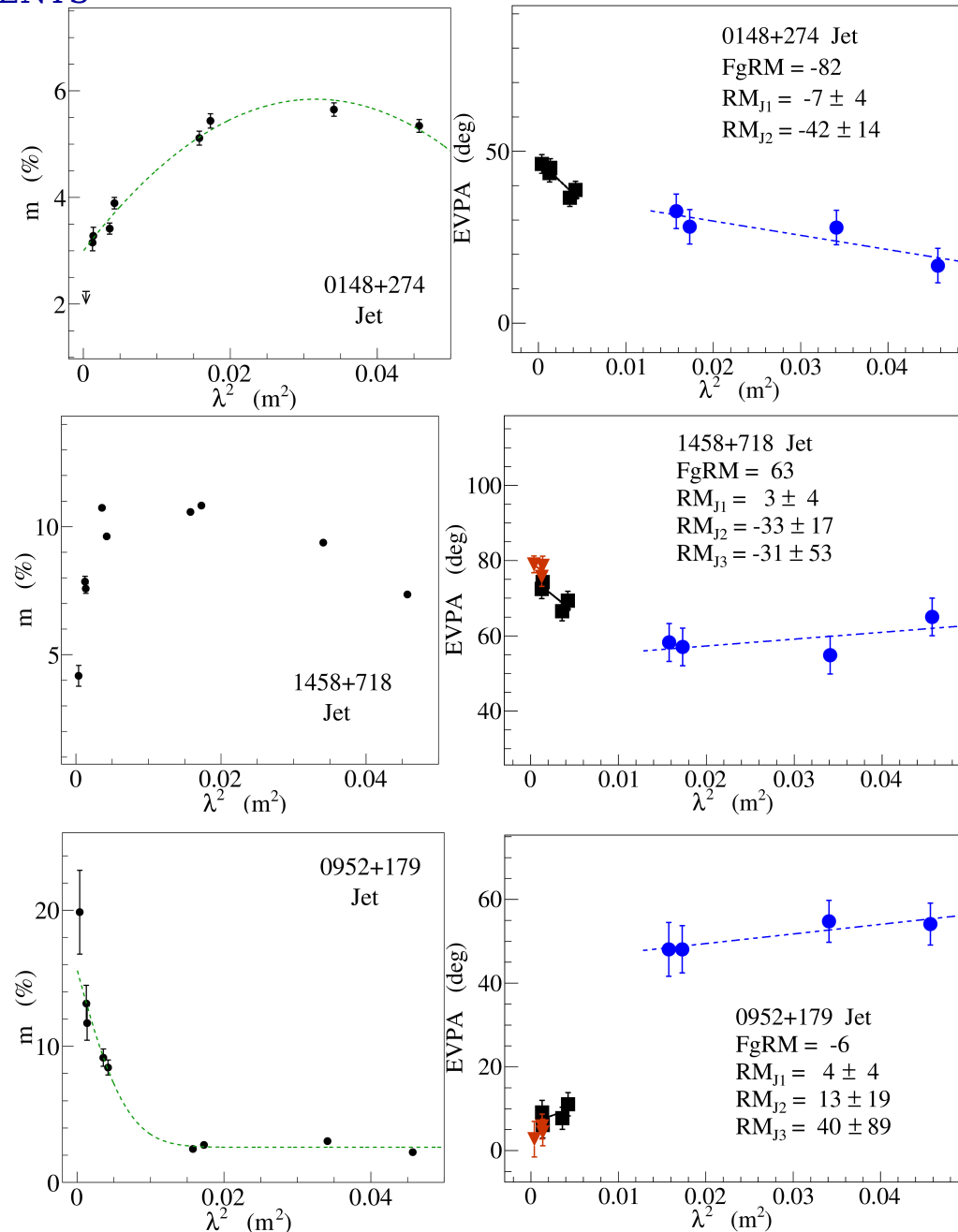
EXTERNAL ROTATION - RANDOM (TURBULENT) OR REGULAR MAGNETIC FIELDS IN INHOMOGENOUS INOZED MEDIUM

# FRACTIONAL POLARIZATION & EVPA

INTERNAL ROTATION: **TRANSPARENT** JET COMPONENTS



THIS DATA AND FIGURES ARE MADE FOR MODELLED OPTICALLY THIN JET COMPONENTS (PIXEL VALUES)



MAJORITY TRANSPARENT JETS EXHIBITS ROTATION ON EXTERNAL FARADAY SCREENS, WHILE FEW SOURCES SHOW INTERNAL ROTATION

EXTERNAL ROTATION - RANDOM (TURBULENT) OR REGULAR MAGNETIC FIELDS IN INHOMOGENOUS IONIZED MEDIUM

INTERNAL ROTATION — HELICAL OR TWISTED MAGNETIC FIELDS

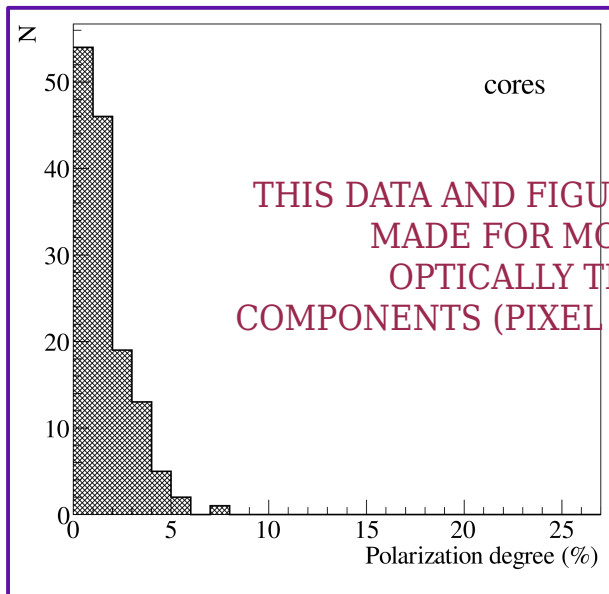
# FRACTIONAL POLARIZATION & EVPA

## OPAQUE JET COMPONENTS

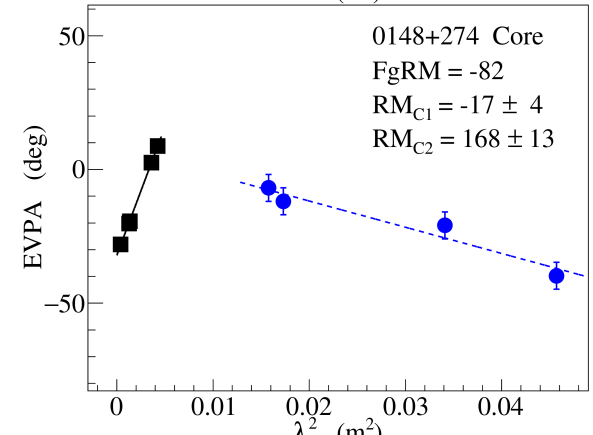
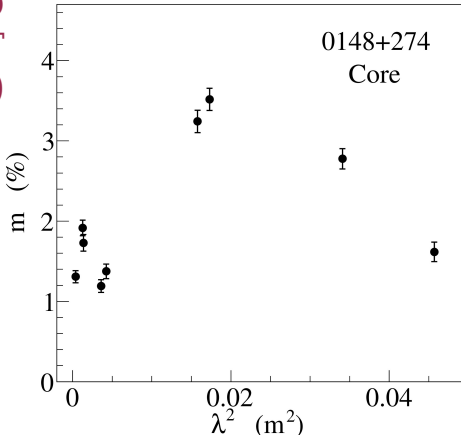
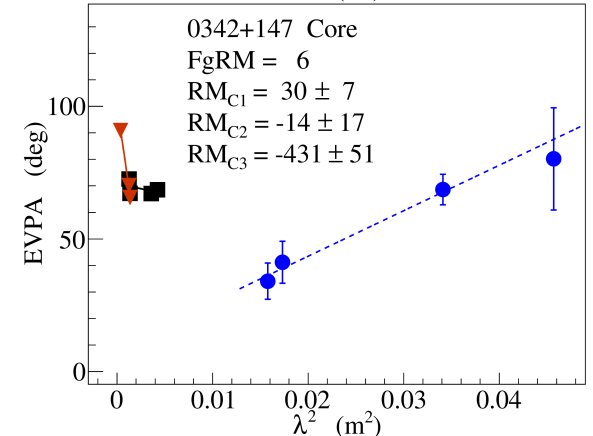
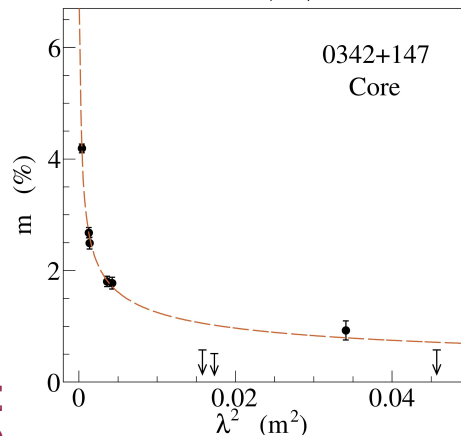
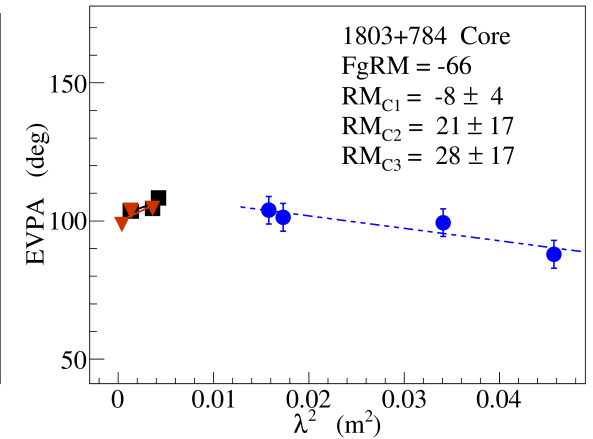
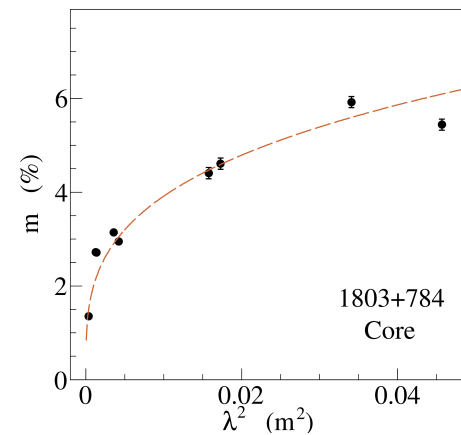
OPAQUE CORES HAVE MORE COMPLEX POLARIZED STRUCTURE, APPLYING THAT POLARIZED FLUX GOES FROM DIFFERENT REGIONS WITHIN A SOURCE

PHYSICAL INTERPRETATION:

HELICAL/TWISTED MAGNETIC FIELDS  
 MULTIPLE ROTATION MEASURE OR JET COMPONENTS SMEARED WITHIN A BEAM  
 DEPOLARIZATION ON EXTERNAL FARADAY SCREEN

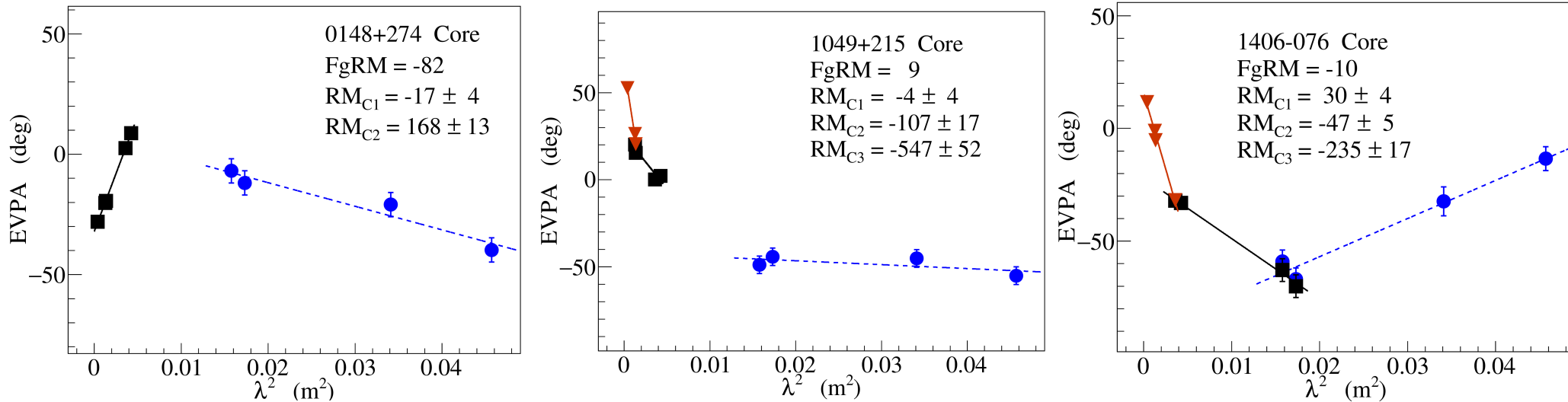


THIS DATA AND FIGURES ARE MADE FOR MODELLED OPTICALLY THICK JET COMPONENTS (PIXEL VALUES)

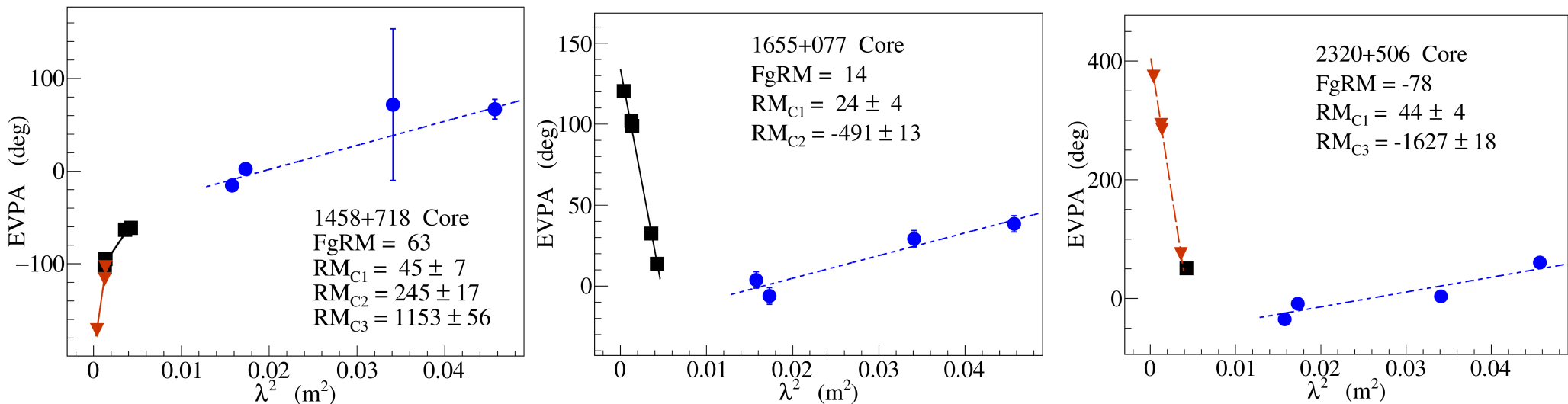


# EVPA IN THE CORES

RM INCREASE TOWARDS SHORTER WAVELENGTHS: CORE SHIFTS



THIS DATA AND FIGURES ARE MADE FOR MODELLED OPTICALLY THICK JET COMPONENTS (PIXEL VALUES)



EVPA<sub>s</sub> ARE CORRECTED FOR GALACTIC RM

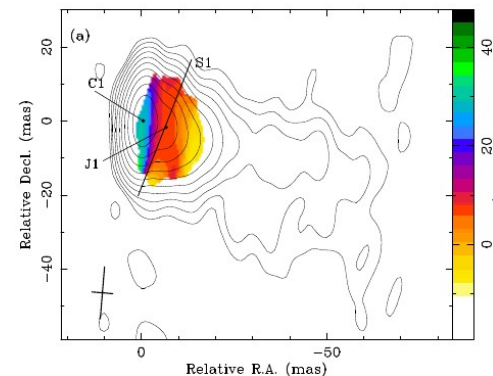
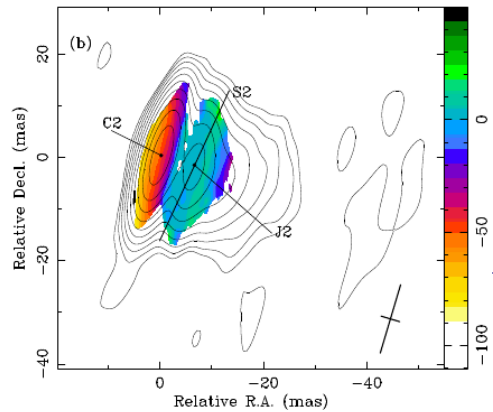
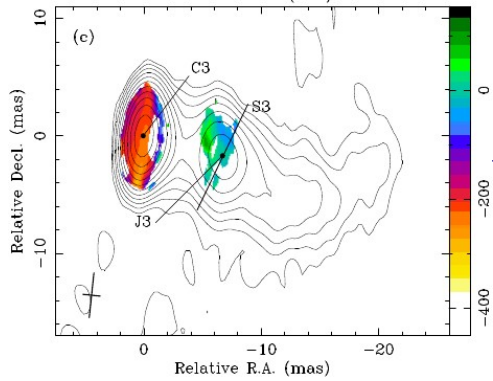


# FARADAY ROTATION MEASURES

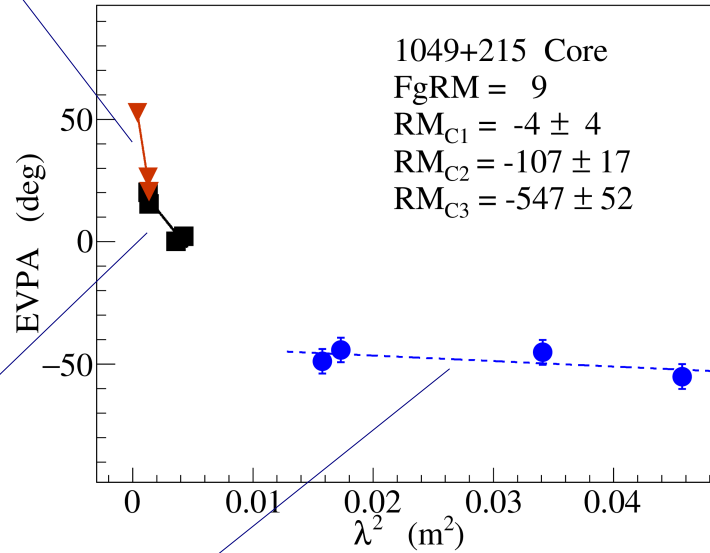
$$RM \sim \int n_e B_{\parallel} dl$$

CONSTRUCTED 43 ROTATION MEASURE MAPS

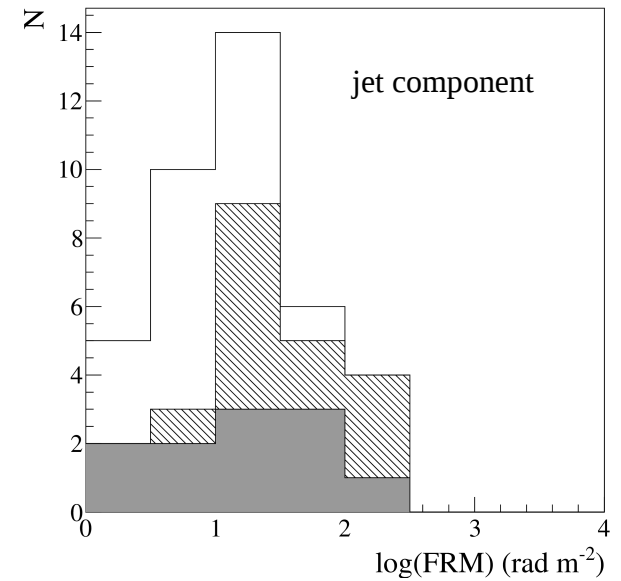
1.4 – 2.4 GHz   
 2.4 – 5.0 GHz   
 4.6 – 15.4 GHz



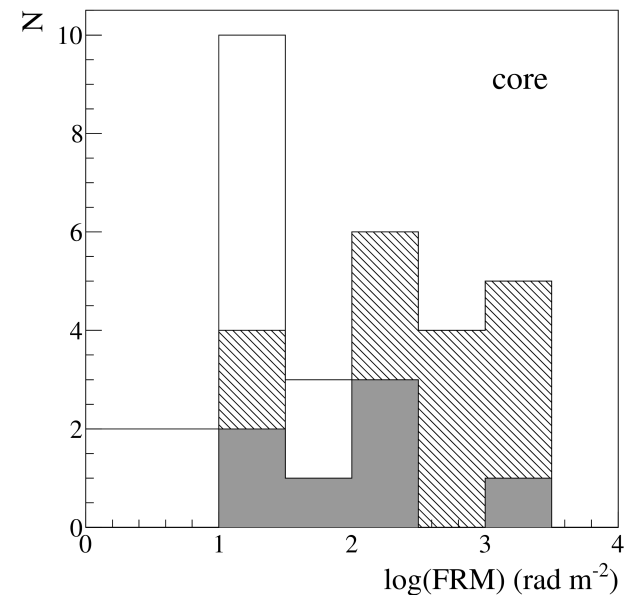
ROTATION MEASURE



ROTATION MEASURE VALUES ARE VARIABLE AT MONTHS – YEARS SCALES, AS SEEN FROM COMPARISON OF OUR VALUES WITH MEASURES MADE BY OTHER AUTHORS

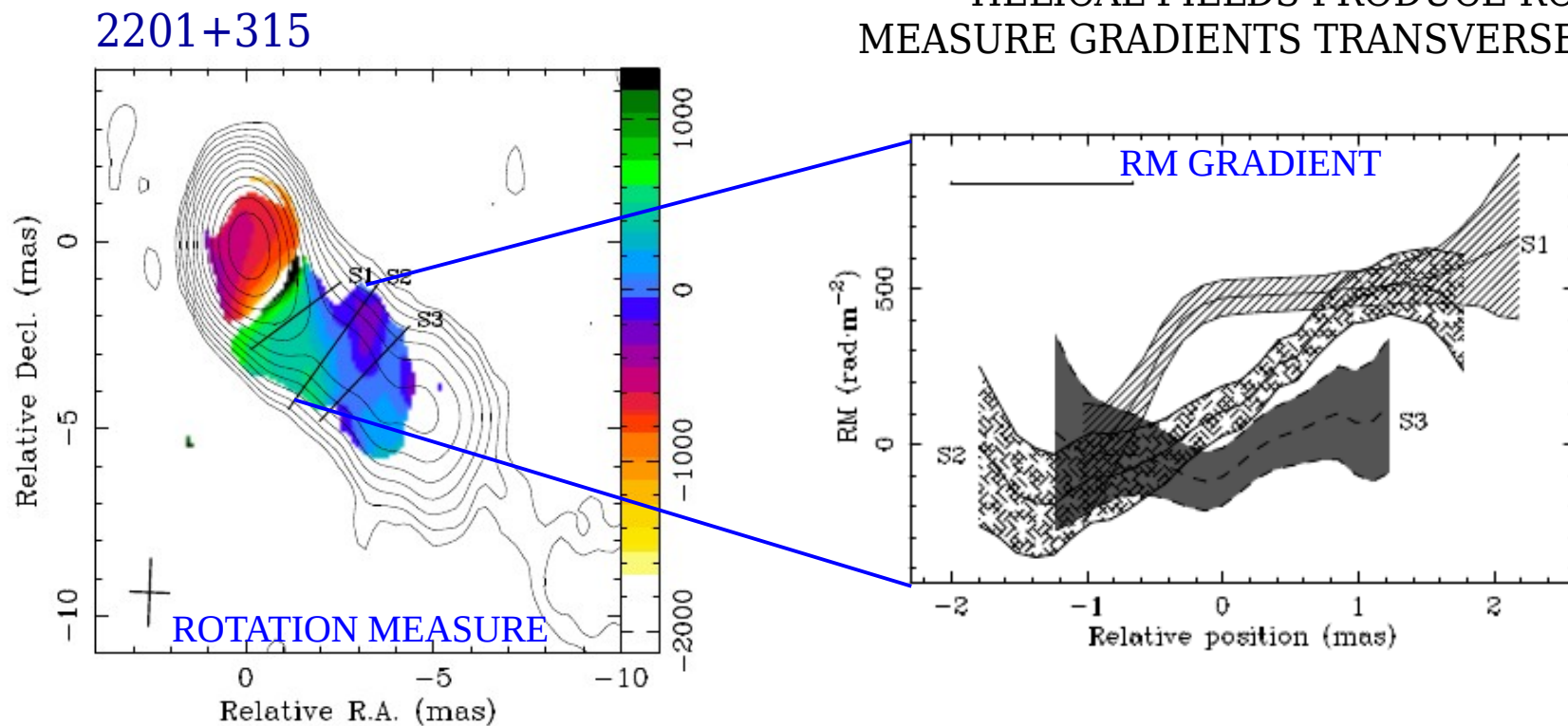


OBSERVABLE ROTATION MEASURE, CORRECTED FOR GALACTIC CONTRIBUTION



# TRANSVERSE ROTATION MEASURE GRADIENT

HELICAL FIELDS PRODUCE ROTATION MEASURE GRADIENTS TRANSVERSE TO JET



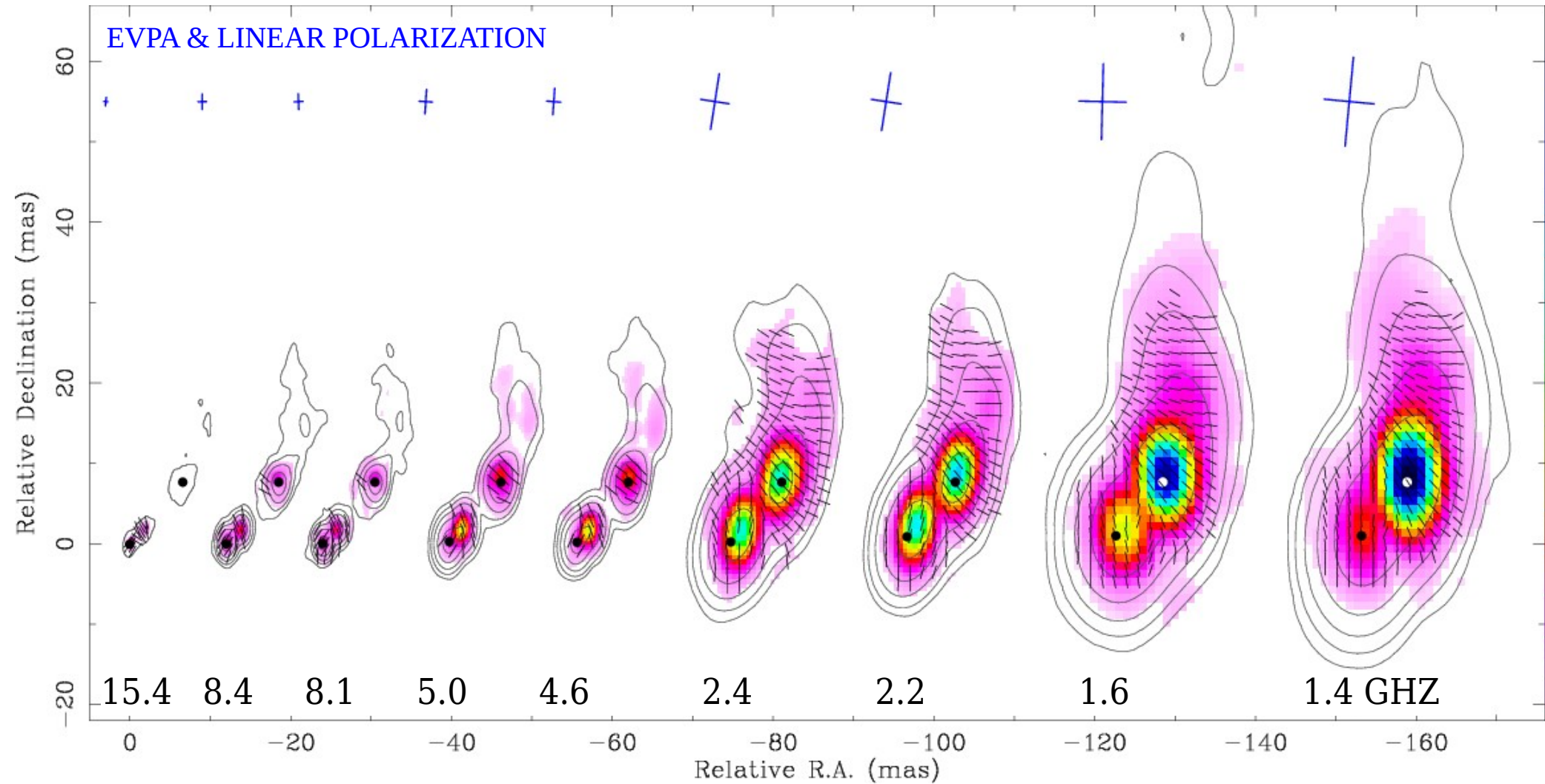
8 OUT OF 20 SOURCES SHOW SIGNIFICANT ROTATION MEASURE GRADIENTS

3 OUT OF THESE 8 MIGHT BE PRODUCED BY  $N_E$  and/or  $B_{\parallel}$  CHANGE RATHER THAN CHANGE IN ORIENTATION OF  $B$

# ELECTRIC FIELD STRUCTURE

CORRECTED FOR FARADAY RM: 0148+274

0148+274

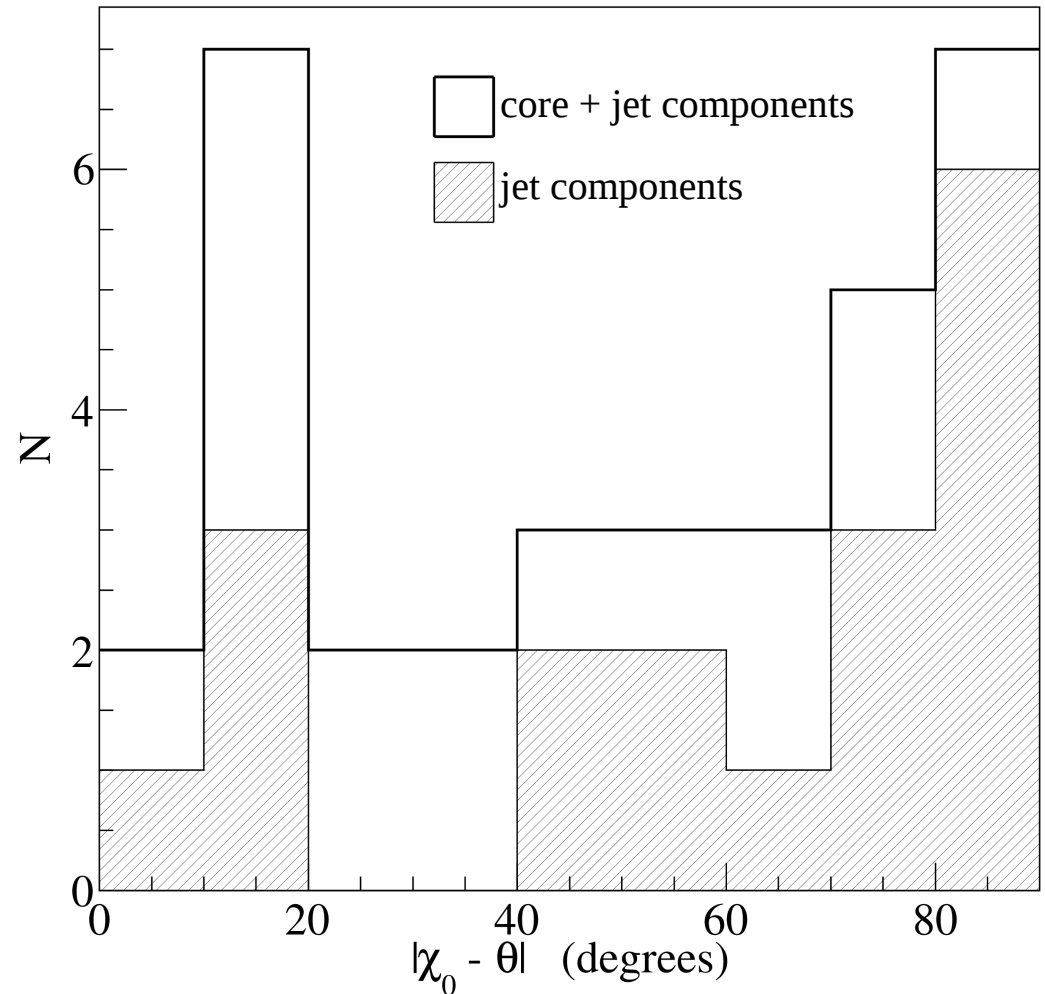


# EVPA vs. JET DIRECTION

CORRECTED FOR FARADAY RM

55% OF THE SOURCES HAVE EITHER  
PARALLEL OR PERPENDICULAR EVPAs  
RELATIVE TO JET DIRECTION

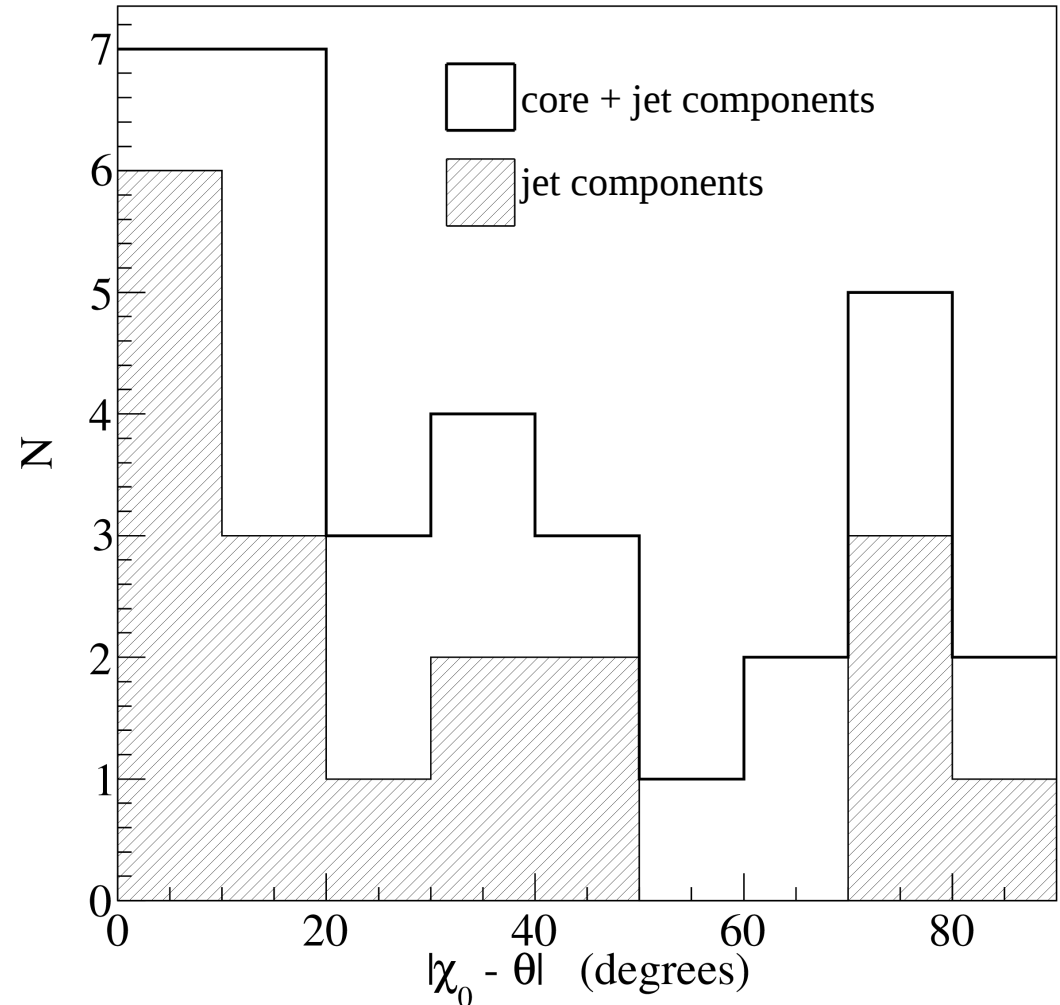
45% OF THE SOURCES INCLINE THEIR  
EVPAs AT ANGLES  $>20^\circ$  &  $<70^\circ$



# EVPA vs. JET DIRECTION

CORRECTED FOR FARADAY RM

JET EVPAs ROTATED BY 90 DEGREES



JET EVPAs ARE ROTATED BY 90 DEGREES  
(TO ALIGN CORE AND JET EVPAs,  
CHANGE IN OPACITY)

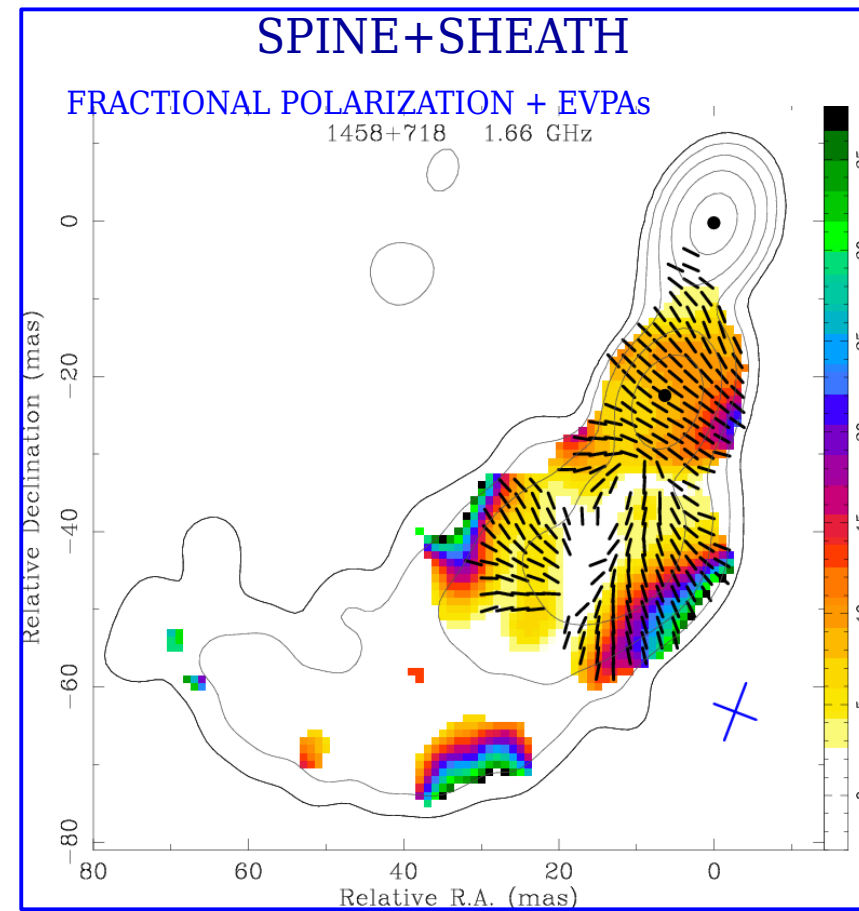
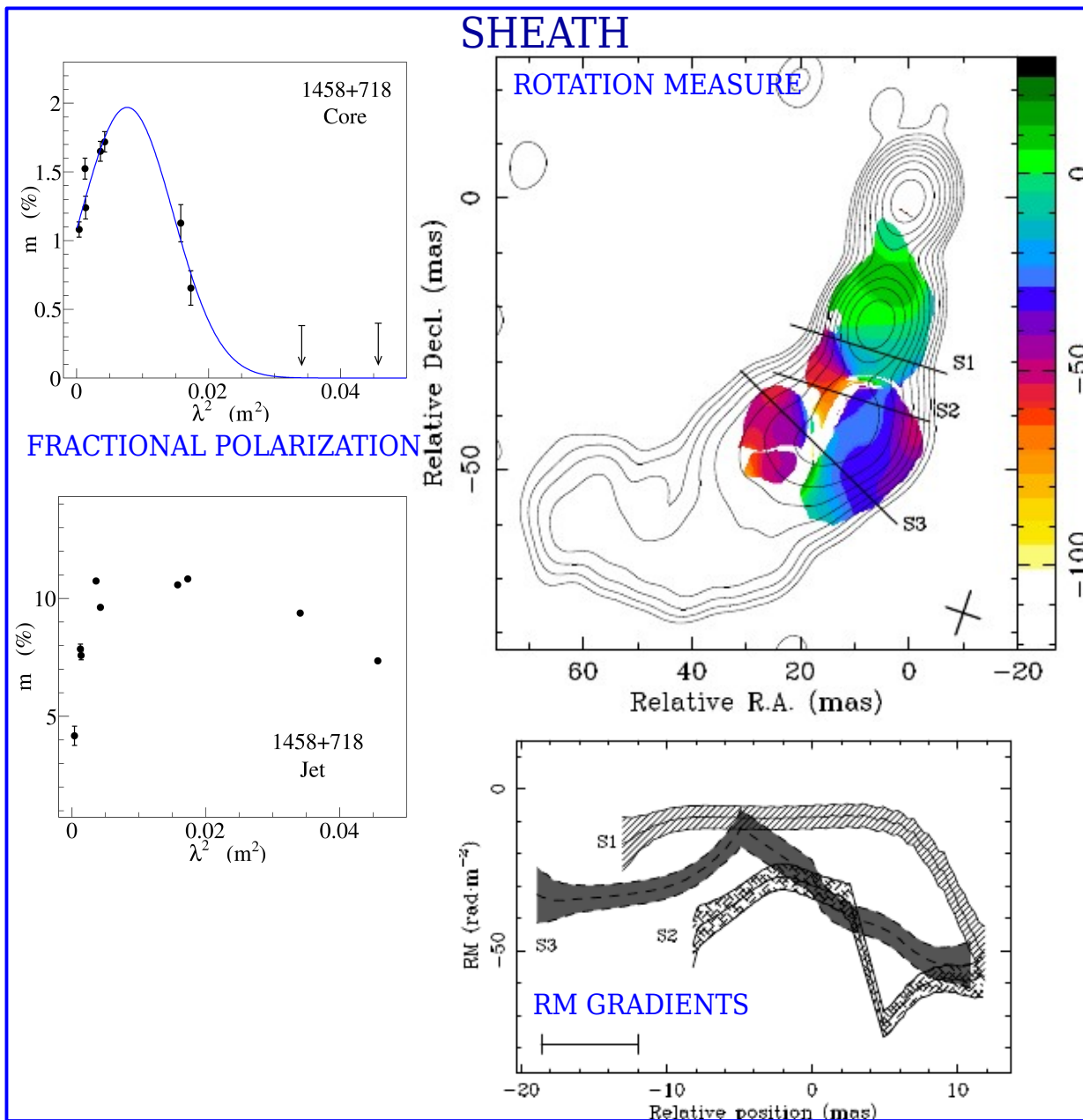
SMALL TENDENCY TO ALIGN EVPAs  
WITH THE JET DIRECTION

ALIGNMENT DOES NOT IMPLY  
EXISTENCE OF POLOIDAL MAGNETIC  
FIELD IN THE JET

SUCH DISTRIBUTION HAS BEEN  
OBSERVED BEFORE (E.G.  
POLLACK+2003, LISTER&HOMAN 2005,  
AGUDO+2014, )

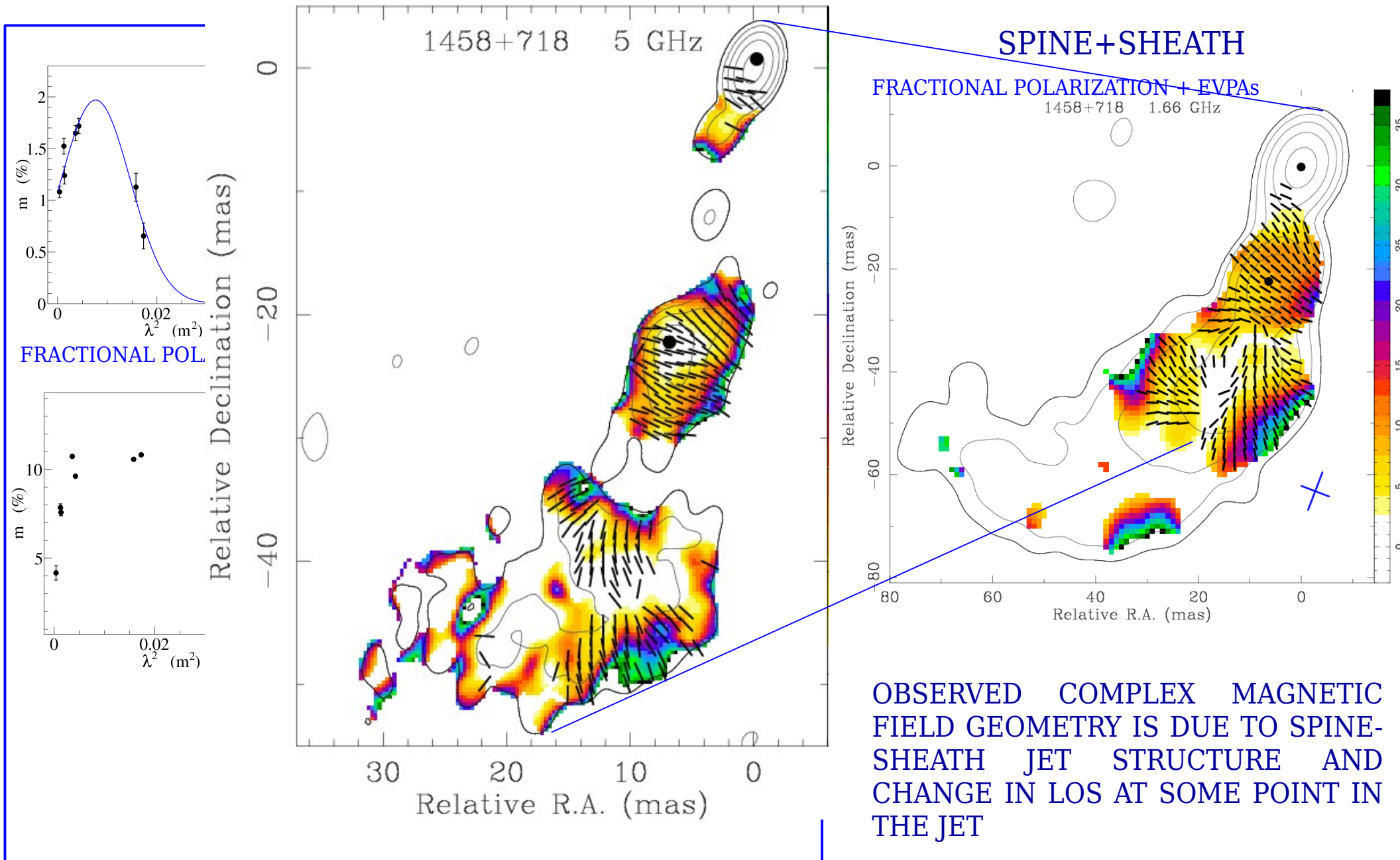
# ELECTRIC FIELD STRUCTURE

CORRECTED FOR FARADAY RM: 1458+718



# ELECTRIC FIELD STRUCTURE

CORRECTED FOR FARADAY RM: 1458+718



# SUMMARY

MAJORITY OF THE SOURCES ARE CONSISTENT WITH THE MODEL OF EXTERNAL FARADAY SCREEN, LOCATED CLOSE TO THE JET (MIGHT BE ITS OUTER LAYER, SHEATH, ETC.)

## 38 CASES (18 SOURCES). OBSERVED MAGNETIC FIELD STRUCTURE:

POLOIDAL in the JET	+++++
RANDOM/REGULAR in the SCREEN	+++++
HELICAL in the JET	+++++
HELICAL in the SCREEN	+++++
UNKNOWN in the JET/SCREEN	++++

THERE IS NO SINGLE, ULTIMATE MODEL OF THE MAGNETIC FIELD STRUCTURE ABLE TO DESCRIBE ALL OBSERVED SOURCES

SPINE (CENTRE) - SHEATH (SURROUNDING MEDIA) STRUCTURE, LOS AND RELATIVISTIC EFFECTS MAY DESCRIBE OBSERVED VARIETY OF POLARIZED CHARACTERISTICS. IN THIS CONTEXT, SHEATH HOLDS TURBULENT, REGULAR POLOIDAL OR TOROIDAL MAGNETIC FIELD. SPINE CONTAINS WEEL-ORDERED, LARGE-SCALE MAGNETIC FIELDS.

POOR SPATIAL RESOLUTION, INSTABILITIES, SHOCKS, JET BENDS, CHANGE IN JET GEOMETRY, JET ASYMMETRY, SOURCE FLARING ACTIVITY AFFECT INTRINSIC MAGNETIC FIELD ORIENTATION