

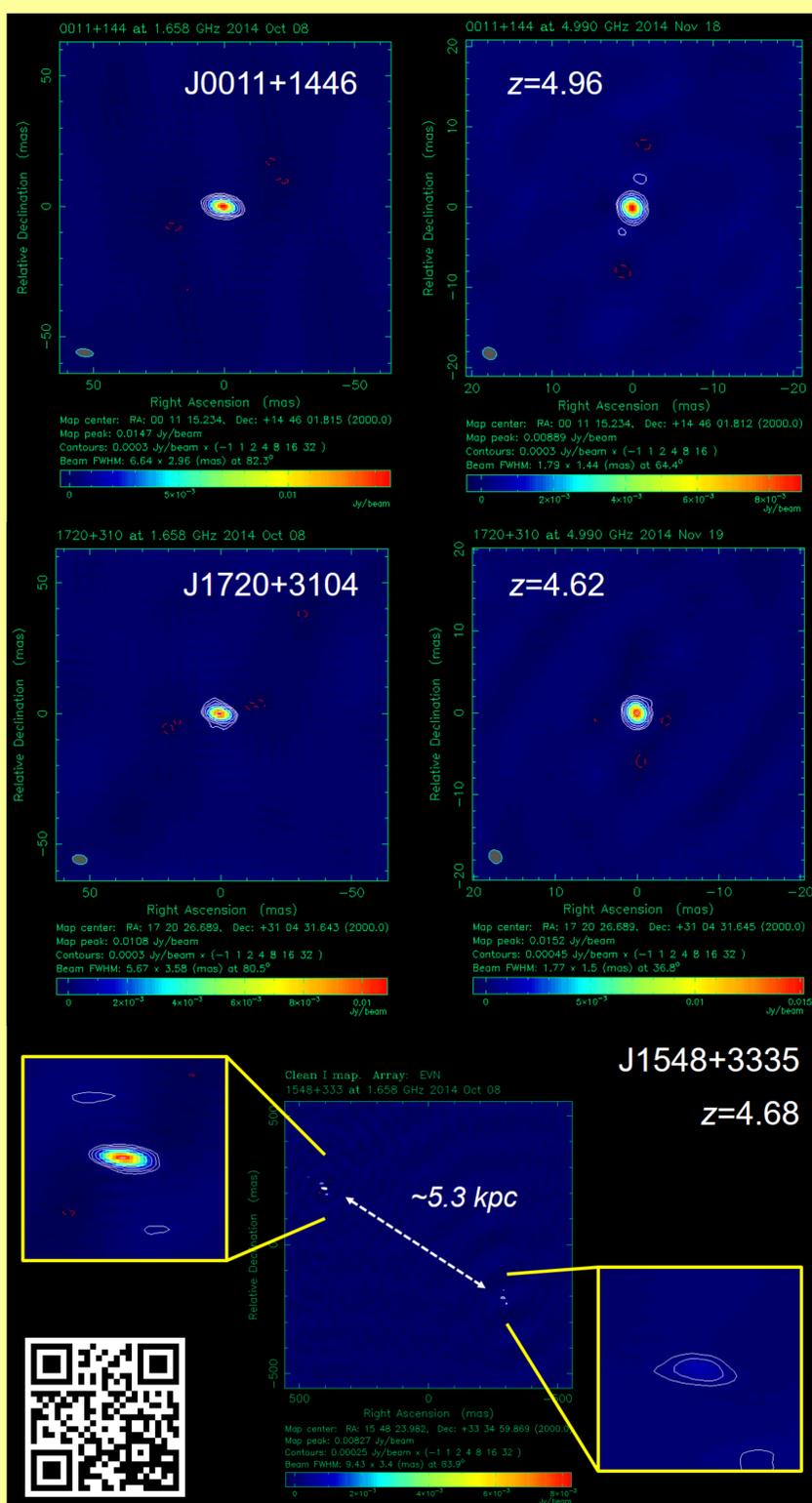
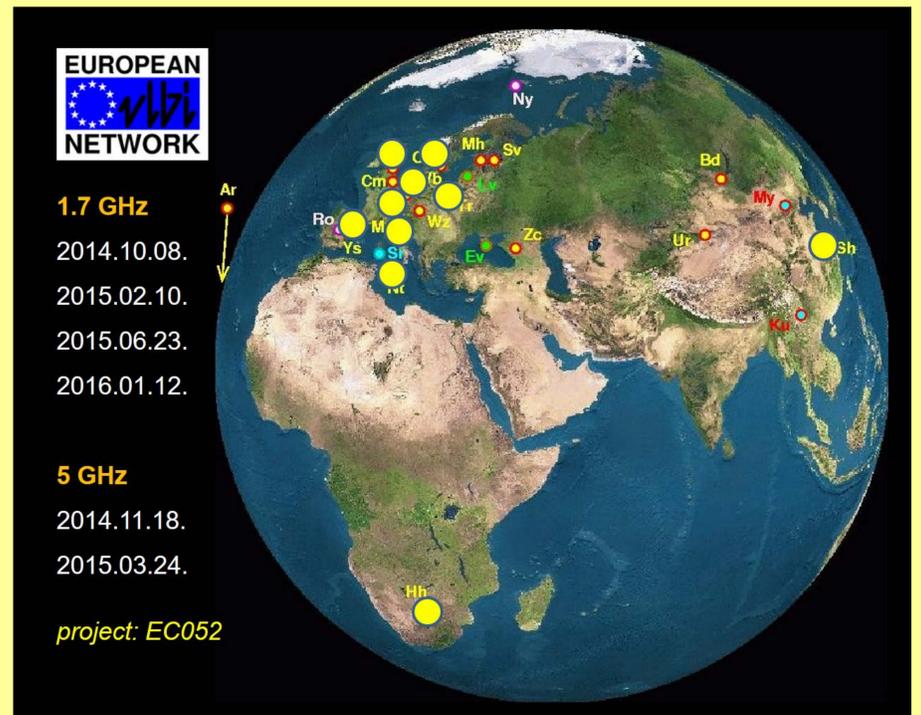
Compact radio quasars at $z > 4.5$ observed with the EVN

Rocco Coppejans¹, Sándor Frey², Dávid Cseh¹, Cornelia Müller¹, Zsolt Paragi³, Krisztina Gabányi^{2,4}, Leonid Gurvits^{3,5}, Tao An^{6,7}, Heino Falcke^{1,8}, Oleg Titov⁹

¹ Department of Astronomy, Radboud University, Nijmegen (NL), ² FÖMI Satellite Geodetic Observatory, Budapest (HU), ³ Joint Institute for VLBI ERIC, Dwingeloo (NL), ⁴ Konkoly Observatory, Hungarian Academy of Sciences, Budapest (HU), ⁵ Department of Astrodynamics and Space Missions, Delft University of Technology (NL), ⁶ Shanghai Astronomical Observatory, Chinese Academy of Sciences (CN), ⁷ Key Laboratory of Radio Astronomy, Chinese Academy of Sciences (CN), ⁸ ASTRON Netherlands Institute for Radio Astronomy, Dwingeloo (NL), ⁹ Geoscience Australia, Canberra (AU)

Why are radio quasars at extremely high redshifts interesting?

- At redshifts above $z > 4.5$, the Universe was younger than $\sim 10\%$ of its present age
- Active galactic nuclei (AGN) powered by supermassive black holes are seen at least until $z \sim 7$
- High-redshift quasars are associated with the youngest supermassive black holes in the Universe
- Some of them are **radio-loud** and thus can be studied with VLBI at the highest possible resolution; radio provides an unobscured view of the nucleus
- They can be used to test the predictions of **cosmological models** (angular size-redshift and apparent proper motion-redshift relation)
- They can be used to set constraints on **black hole growth** in the early Universe and better understand **galaxy evolution**
- Recently we increased the number of VLBI-imaged AGN at $z > 4.5$ by 50%, from 20 to 30, using e-EVN measurements



Sample selection

- Optical quasars with known spectroscopic redshift $z \geq 4.5$ (mostly from SDSS)
 - **Radio emission**, compact on arcsec scales (1.4-GHz flux density $S > 5$ mJy from the VLA FIRST survey)
 - Not yet observed with VLBI
- ⇒ 10 new targets identified with $4.5 \leq z \leq 5$
- Still very few similar objects are known*

Results in a nutshell

- **All 10 targets are detected** with the EVN (9 at both frequencies, 1 at 5 GHz only)
- The images mostly show **featureless objects**, but the measured properties indicate a **mix of resolved and compact sources** on mas scales (*two examples are shown in the images*)
- The most intriguing one of the targets (J1548+3335) has a wide-separation **double structure** seen at 1.7 GHz (*see the image*) the brighter component that coincides with the optical quasar position is also detected at 5 GHz

The sample of $z > 4.5$ radio sources observed with VLBI to date...

- ...consists of 30 objects and is **incomplete and heterogeneous**
- The highest redshift is $z = 6.21$ (J1429+5447)
- The radio emission in one source (J1205-0742) originates from star formation; the others are AGN
- Based on **spectral index, brightness temperature** and (limited) **variability** information, roughly equal numbers of the rest are either **flat-spectrum radio quasars** or **steep-spectrum sources** (including two wide extended ones)
- Objects in the steep-spectrum category are most likely young sources with spectra peaking at GHz or ~ 100 MHz frequencies

Coming next:

- **Larger well-defined high-redshift radio quasar samples**
- **Better understanding of the selection effects**
- **Tackle the question of apparent overabundance of highly-beamed sources** (i.e. blazars) at very high redshifts

More details and further references:

Coppejans R. et al. 2016, *MNRAS*, in press (arXiv:1609.00575)